

Discovery of non-equilibrium ionization plasma around the Fermi Bubble

; new evidence of past activity of the Galactic Center

WASEDA



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Abstract

Fermi bubbles are giant gamma-ray structure toward the Galactic center (GC) with symmetrical north-south extension perpendicular to the galactic plane. Such giant structures toward the GC are also observed in various wavelengths from radio to X-rays, such as WMAP haze, North Polar Spur (NPS), and most recently, eROSITA bubbles. We investigated the detailed plasma condition of the NPS/Loop I around Fermi bubble using archival Suzaku data. In previous research collisional ionization equilibrium (CIE) have been assumed for plasma state, but we also assume non-equilibrium ionization (NEI) to check the plasma condition in more detail. We found that most of the plasma in the NPS/Loop I favors the state of NEI, and has the density-weighted ionization timescale of $n_e t \sim 10^{11-12} \text{ s cm}^{-3}$ and the electron number density $n_e \sim \text{a few} \times 10^{-3} \text{ cm}^{-3}$. The plasma shock age, t , or the time elapsed after the shock front passed through the plasma, is estimated to be on the order of a few Myr for the NPS/Loop I, which puts a strict lower limit to the age of the whole NPS/Loop I structure. We found that NEI results in significantly higher temperature and lower emission measure than those currently derived under CIE assumption. The electron temperature under NEI is estimated to be as high as 0.5-keV toward the brightest X-ray NPS ridge at $\Delta\theta = -20^\circ$, which decreases to 0.3 keV at -10° , and again increases to $\sim 0.6 \text{ keV}$ towards the outer edge of Loop I at $\Delta\theta \sim 0^\circ$, about twice the currently estimated temperatures. Here, $\Delta\theta$ is the angular distance from the outer edge of Loop I. We discuss the implication of introducing NEI for the research in plasma states in astrophysical phenomena.

1. Introduction

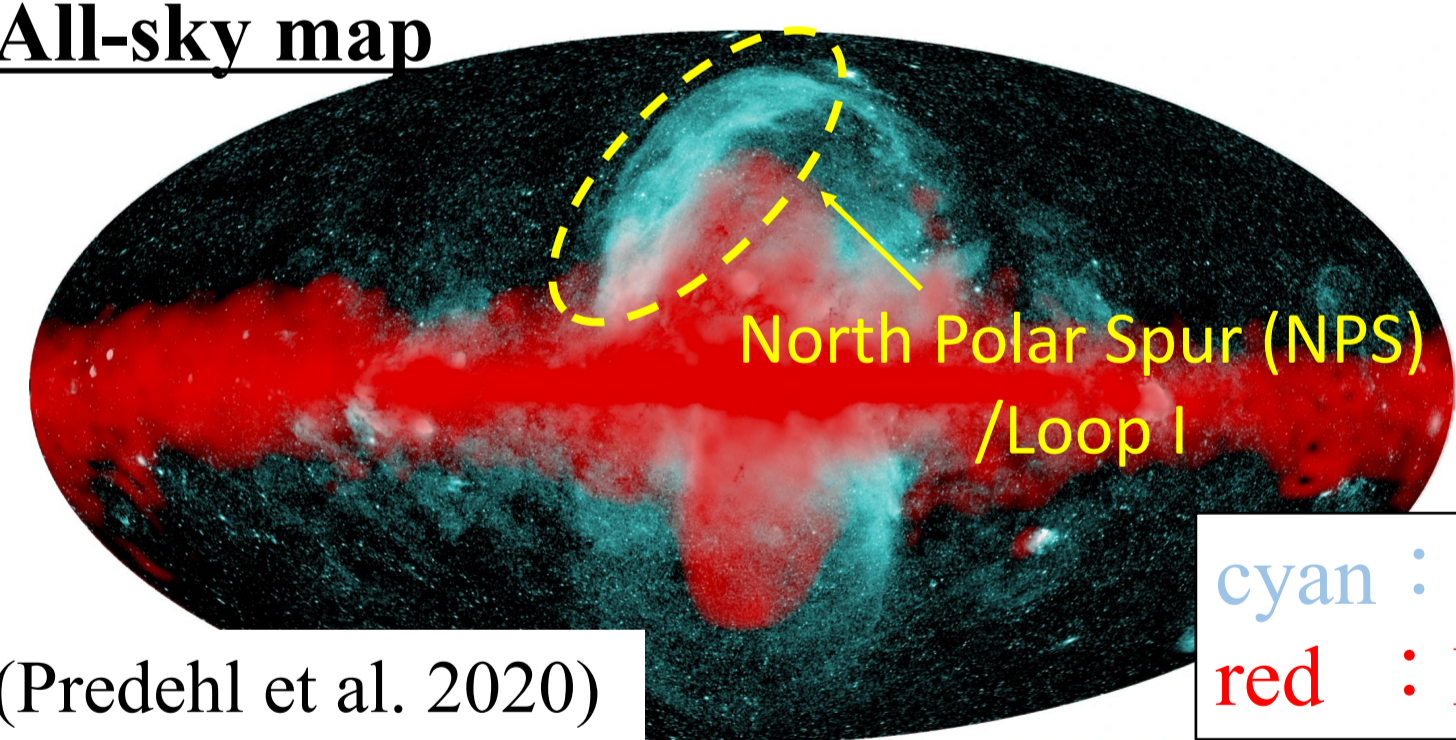
NPS (North Polar Spur) / Loop I structure

- ◆ Large structure towards GC (Galactic Center)
- ◆ Observation with various wavelengths
ex) microwave : WMAP haze (Finkbeiner 2004)
numerous loop structure
X-ray : NPS, e-Rosita Bubble
 γ -ray : Fermi Bubble (Su et al. 2010)

Hypotheses on 2 prevailing & opposing origins

GC explosion @ $\sim 10 \text{ kpc}$ VS nearby SNR @ $\sim 100 \text{ pc}$
 \Rightarrow The origin of these structures is under discussion

All-sky map



cyan : eROSITA 0.6-1.0 keV
red : Fermi bubbles

Previous research

- ◆ Plasma is assumed to be in ionization equilibrium (CIE)
 $T_i(\text{proton temp}) = T_e(\text{electron temp}) = T_z(\text{ionization temp})$

In this research

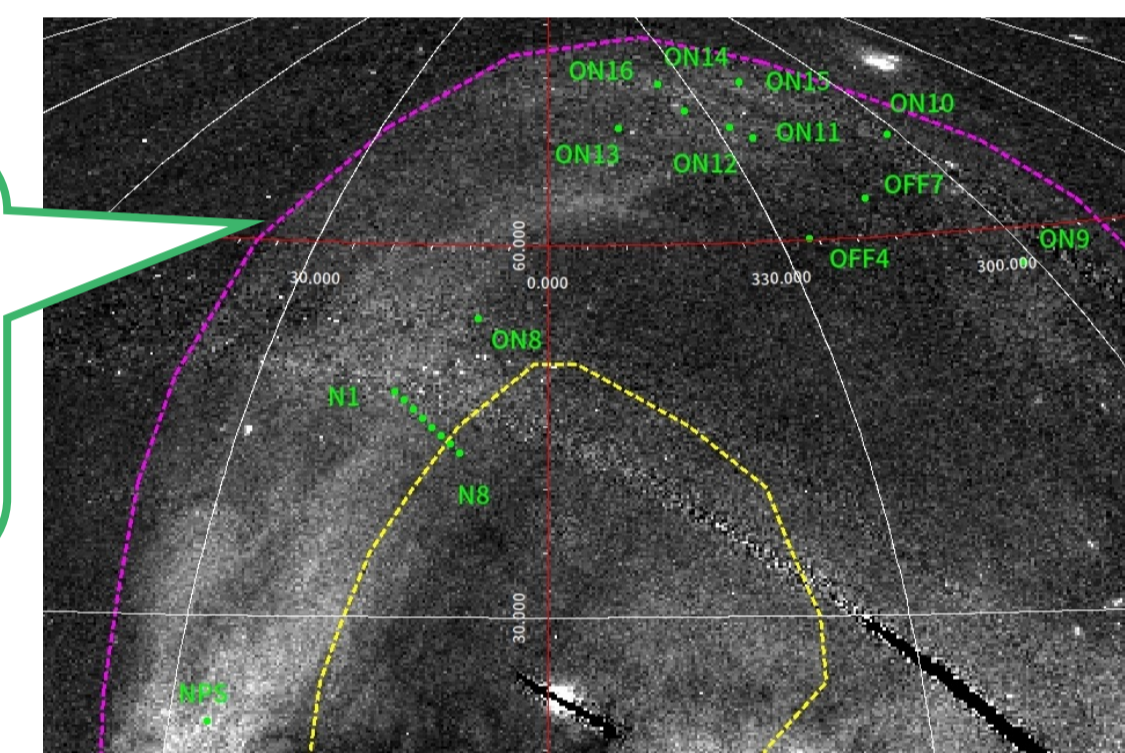
- ◆ Consider ionization non-equilibrium (NEI)
 \Rightarrow ionization timescale $n_e t$: electron density \times shock age

Investigate the origin by estimating plasma temperature, density, and time scale

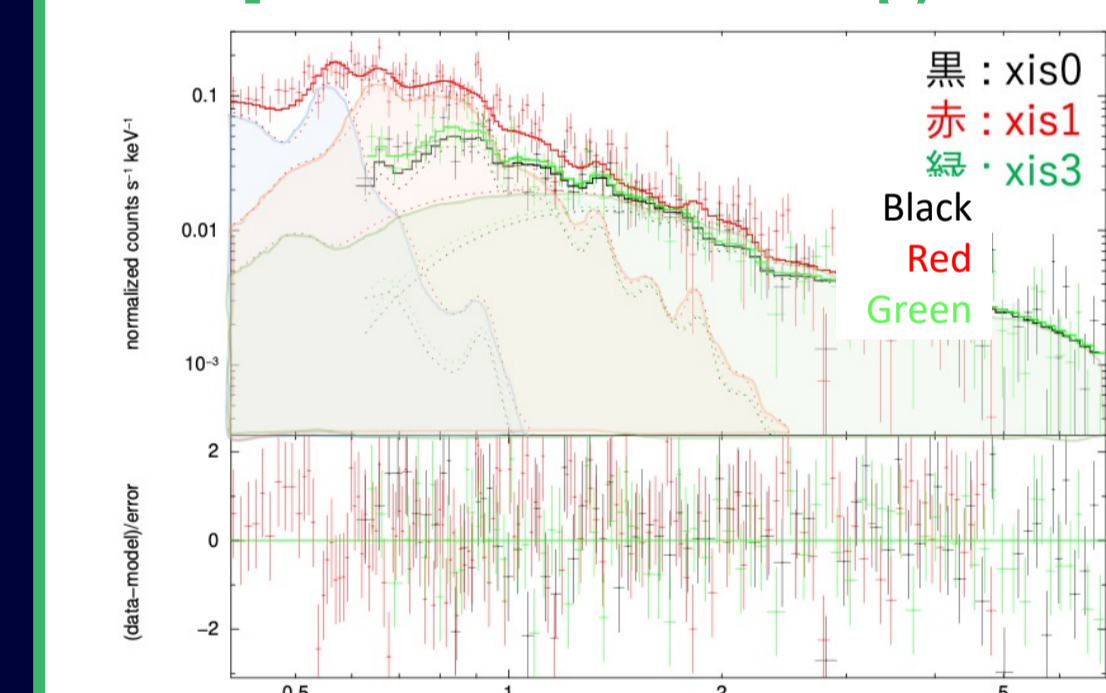
2. observation region & spectrum analysis

Observation region

- ◆ Archive data from Suzaku XIS detector
Low-latitude ($b < 50^\circ$) : 9 regions
High-latitude ($b > 50^\circ$) : 11 regions



Spectrum fitting model



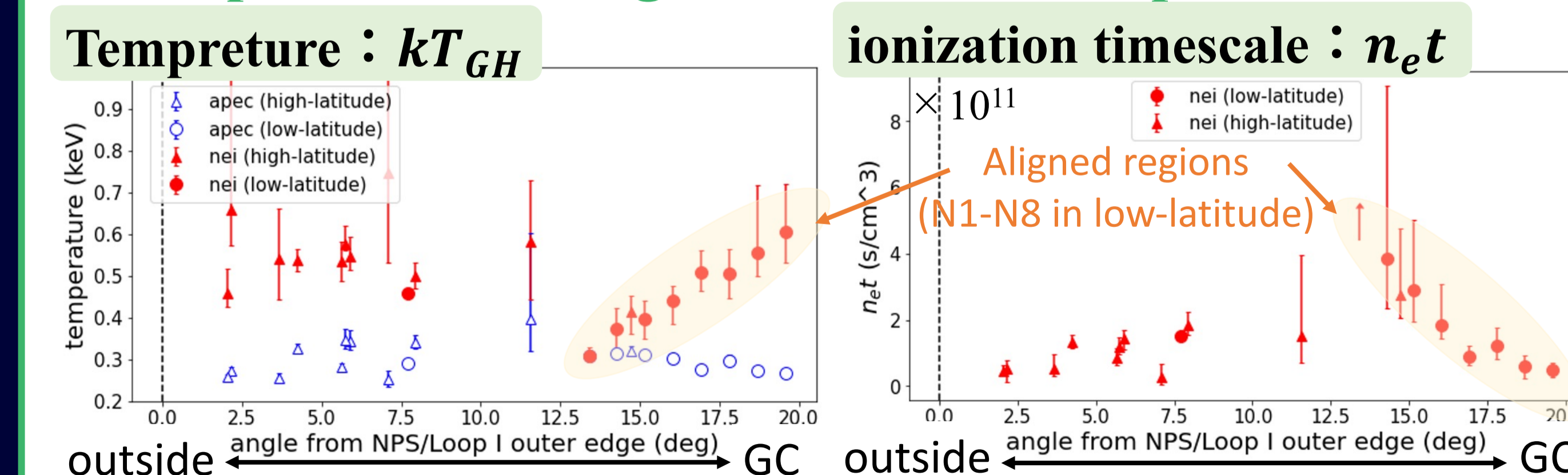
Apec + wabs*(Apec/Nei + Powerlaw)

※Apec model : CIE, Nei model : NEI

- ① LHB / SWCX
- ② GH (Galactic Halo)
 \Rightarrow CIE (Apec model) or NEI (NEI model)
- ③ CXB (Cosmic X-ray Background)

3. Results

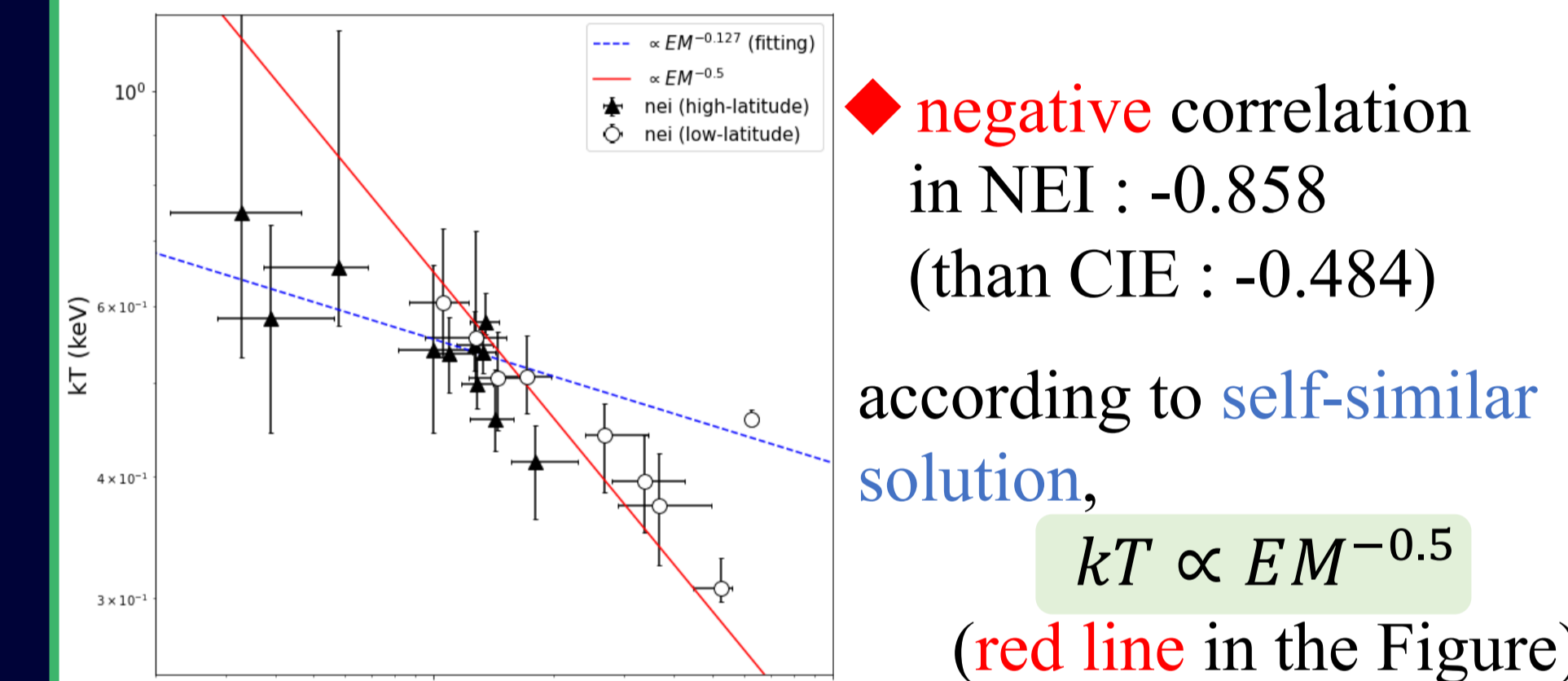
Comparison of fitting results of GH component



- ◆ Plasma temperature is $kT_e \sim 0.6 \text{ keV}$ assuming NEI (hotter than CIE)
- ◆ Ionization timescale is $n_e t \sim 10^{11} \text{ s/cm}^3 \Rightarrow$ plasma is in state of NEI
- ◆ NEI model is favored than CIE with more than 95% confidence level

4. Discussion

Correlation between EM & kT



- ◆ negative correlation in NEI : -0.858 (than CIE : -0.484)
- ◆ according to self-similar solution,
 $kT \propto EM^{-0.5}$ (red line in the Figure)

Origin of the NPS/Loop I

GC explosion (@ \sim few kpc)

- ◆ shock age : $t \sim 1 \text{ Myr}$
- ◆ electron density : $n_e \sim (2 - 5) \times 10^{-3} \text{ cm}^{-3}$
- ◆ energy : $E_g \sim 10^{54-55} \text{ erg}$

nearby SNR (@ $\sim 100 \text{ pc}$)

- ◆ shock age : $t \sim 300 \text{ kyr}$
- ◆ electron density : $n_e \sim 2 \times 10^{-2} \text{ cm}^{-3}$
- ◆ energy : $E_g \sim 10^{52} \text{ erg}$

- ◆ Older than typical SNR
 - ◆ Thinner than ISM ($n_e \sim 1.0 \text{ cm}^{-3}$)
 - ◆ Equivalent energy of several dozen type-II SNRs
- \therefore NPS/Loop I is interpreted as a structure from GC

Considering the shock front temperature as $kT \sim 0.6 \text{ keV}$,
 $v_{sh} \geq 460 - 530 \text{ km/s}$

5. Conclusion & Future work

Conclusion

- ◆ NPS/Loop I plasma is NEI status, close to CIE.
 $\Rightarrow \sim \text{Myr-scale event}$
- ◆ The origin of NPS/Loop I is more likely to be an explosion near the GC.

Future work

- ◆ Systematic analysis with increased region
ex) use other satellite data, analyze south region
 \Rightarrow Unified understanding of the giant structure
Understanding the origin and explosion model