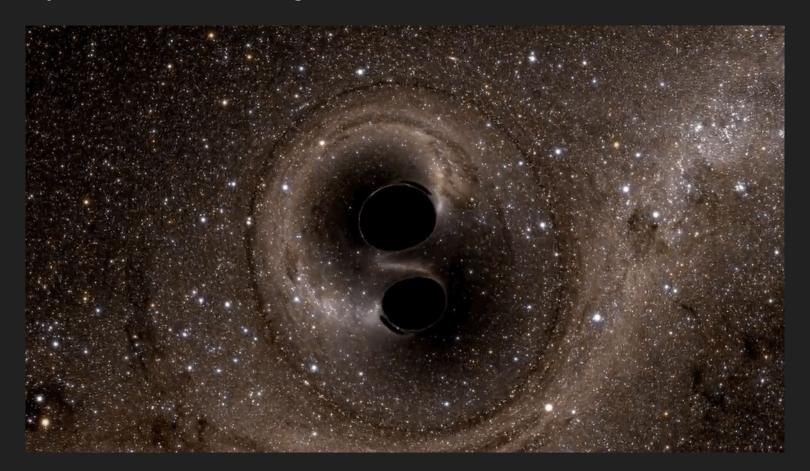




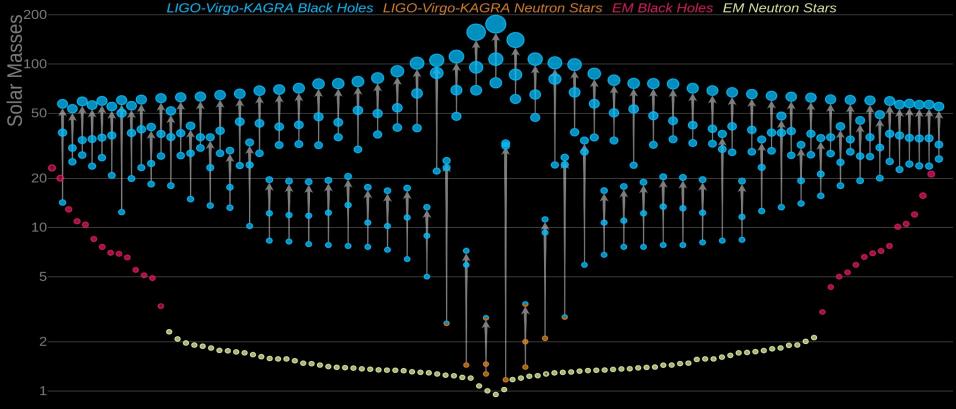
# Black hole mergers from star clusters

Winter Meeting 2022

## Binary black hole merger



# Masses in the Stellar Graveyard



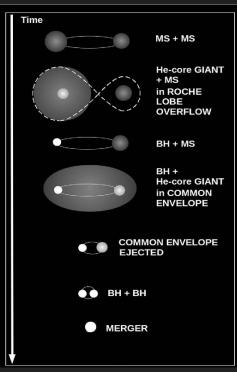
LIGO-Virgo-KAGRA | Aaron Geller | Northwestern

 $\mathcal{R} \sim 44 \; \mathrm{Gpc^{-3}yr^{-1}}$ 

How do black holes get so close to each other?

#### Mechanisms for binary black hole merger formation

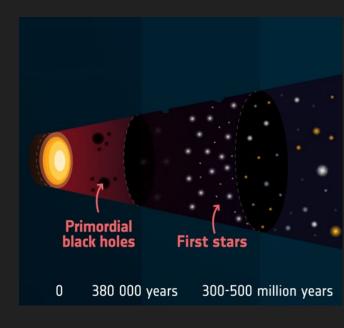
#### Binary stellar evolution



#### Dynamics in star clusters



#### Primordial black holes

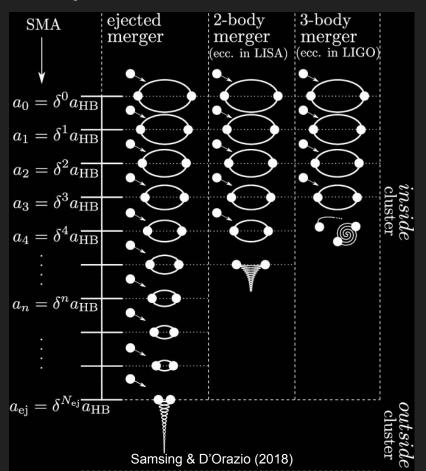


Mapelli (2021)

Isaac Newton Telescopes, La Palma

ESA (2021)

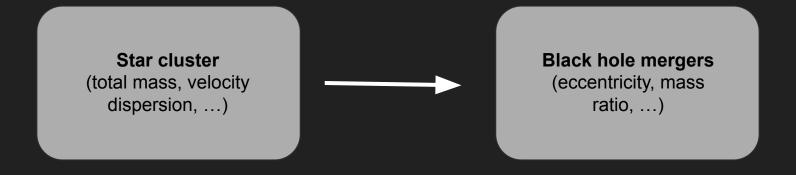
#### Binary black hole dynamics in star clusters



### Separate mechanisms using GW data

	Binary stellar evolution	Dynamical interactions	Primordial binaries
Spin	Aligned	Randomly misaligned	Non spinning
Redshift dependence	Following SFR	Skewed towards z	Steep z dependence
Eccentricity	No	Yes	No

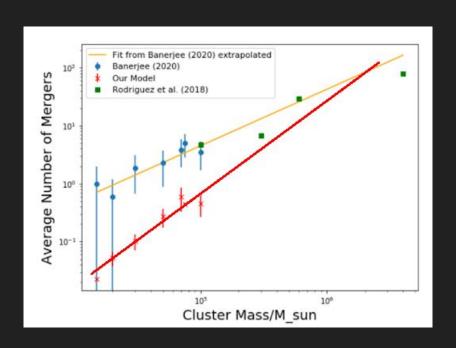
#### Understanding the dynamical origin of black hole mergers



#### How do we simulate star clusters?

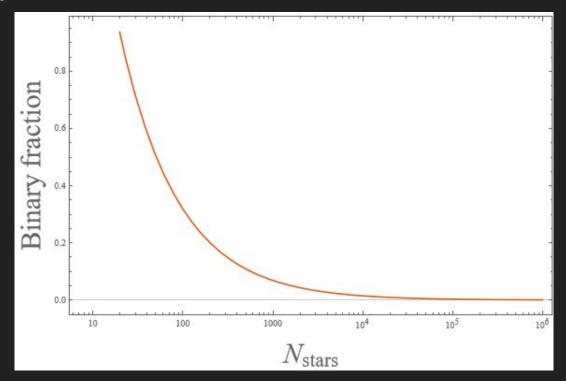
	N-Body integration	Monte Carlo	Fast codes
Speed	Slowest	Slow	Fast
Accuracy	Best	Good	Worst
Examples	NBODY7 PeTar	СМС	cBHBd FASTCLUSTER

#### Underprediction at low cluster masses



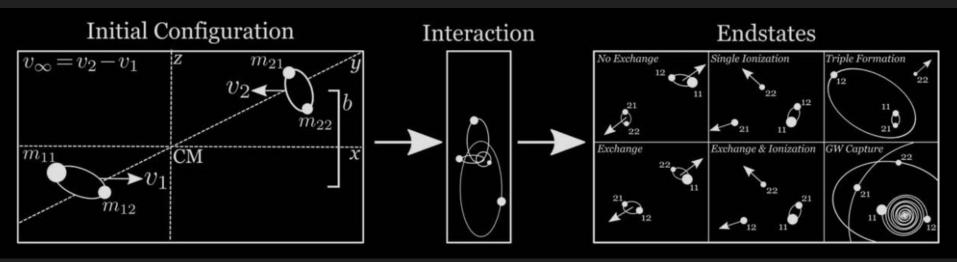
Is the assumption of a single active binary correct?

#### Dynamically-formed binaries



$$N_{\rm binaries} \sim 30 N_{\rm stars}^{-1/3}$$

#### Binary-binary interactions are more likely to merge



Zevin, M et. al. (2019)

# $N_{ m binaries} \propto W_0 N_{BH}^{-1/3}$

A new key ingredient in untangling the merger rate: binary-binary interactions