



# **TEOBResumS:**

## an advanced waveform model for O4

#### R. Gamba for the TEOBResumS (coding) team

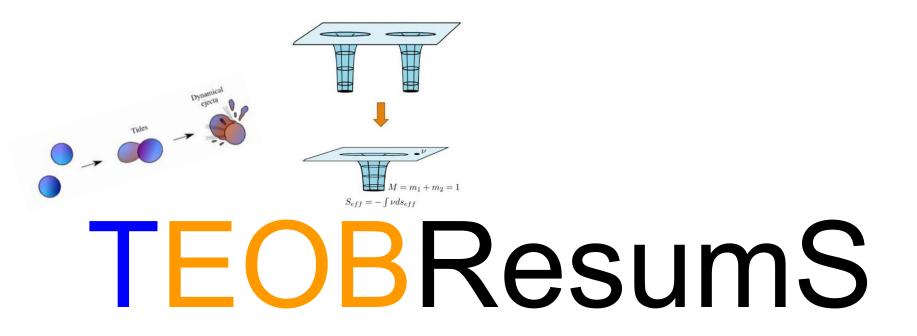
UB, 10.10.2022



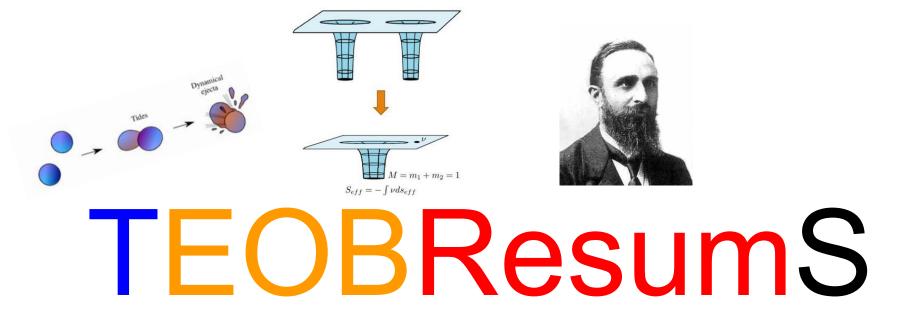
# TEOBResumS



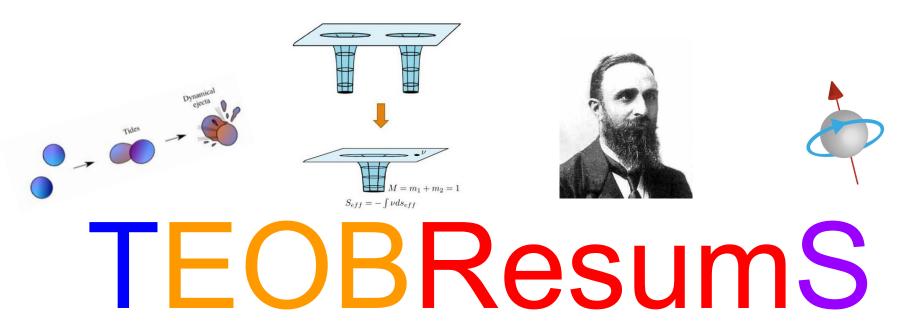




## **Effective One Body**



## (Padé) Resummed









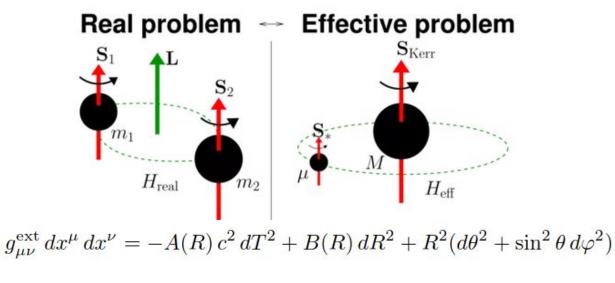


Quasi-circular/Eccentric



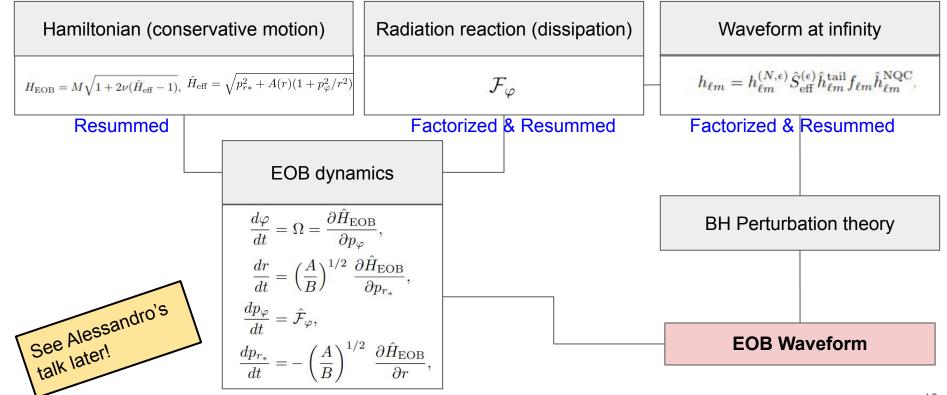
#### **Effective One Body**

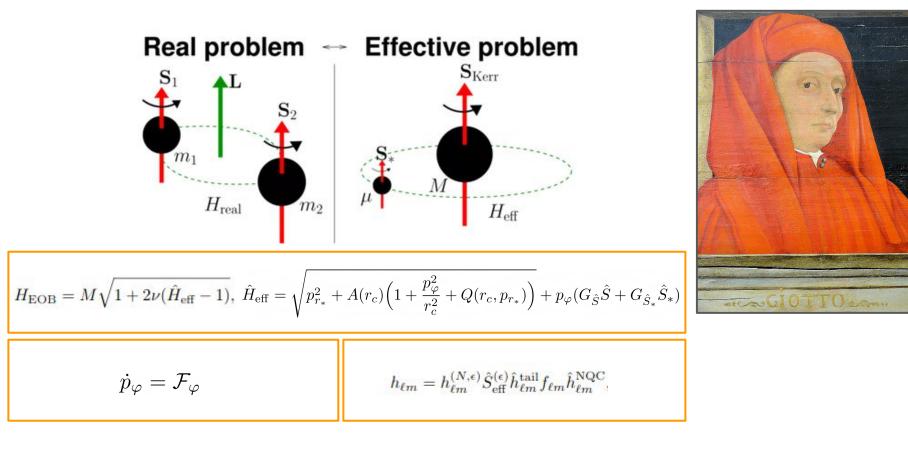
- $\Box$  Newtonian gravity: two body problem  $\rightarrow$  one body moving in effective potential
- $\Box$  General relativity: two body problem  $\rightarrow$  one test particle moving in an effective metric



Polar coordinates (r, arphi) and associated momenta  $(p_{r*}, p_arphi)$ 

### **Effective One Body**



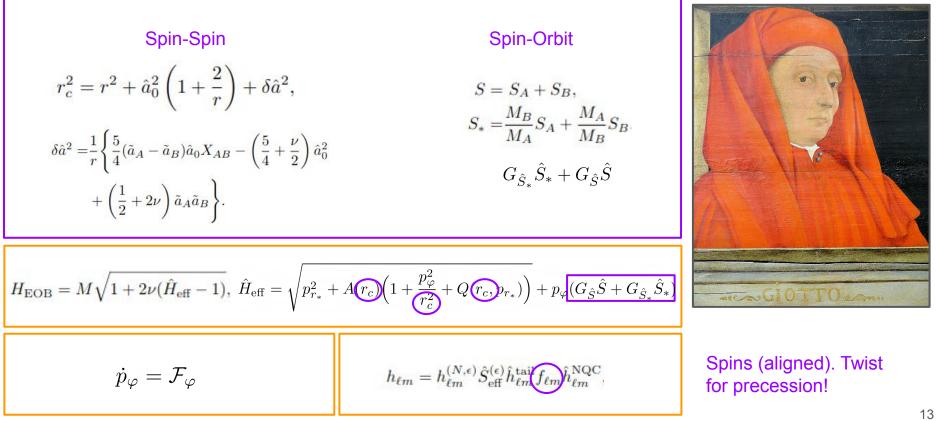


$$A_{4\text{PN}}(u,v) = 1 - 2u + 2vu^3 + v\left(\frac{94}{3} - \frac{41}{32}\pi^2\right)u^4 + v\left(a_5^c(v) + a_5^{\ln}(v)\ln u\right)u^5 \quad \text{4 PN analytical}$$
$$D(u,v) \equiv \frac{1}{1 + 6vu^2 + 2(26 - 3v)vu^3} \quad \text{3 PN analytical}$$

flm also use high order PN information \*and\* test mass info

$$\begin{aligned} H_{\rm EOB} &= M\sqrt{1+2\nu(\hat{H}_{\rm eff}-1)}, \ \hat{H}_{\rm eff} = \sqrt{p_{r_*}^2 + A(r_c)\left(1+\frac{p_{\varphi}^2}{r_c^2} + Q(r_c)p_{r_*})\right)} + p_{\varphi}(G_{\hat{S}}\hat{S} + G_{\hat{S}_*}\hat{S}_* \\ \\ \dot{p}_{\varphi} &= \mathcal{F}_{\varphi} \end{aligned}$$
$$h_{\ell m} = h_{\ell m}^{(N,\epsilon)}\hat{S}_{\rm eff}^{(\epsilon)}\hat{h}_{\ell m}^{\rm tail}f_{\ell m}\hat{h}_{\ell m}^{\rm NQC}, \end{aligned}$$





$$\begin{split} A &= A_0 + A_T^{(+)} \quad A_T^{(+)}(u;\nu) \equiv -\sum_{\ell=2}^4 \left[ \kappa_A^{(\ell)} u^{2\ell+2} \hat{A}_A^{(\ell^+)} + (A \leftrightarrow B) \right] \\ \hat{A}_A^{(2^+)}(u) &= 1 + \frac{3u^2}{1 - r_{\rm LR}u} + \frac{X_A \tilde{A}_1^{(2^+)1\rm SF}}{(1 - r_{\rm LR}u)^{7/2}} + \frac{X_A^2 \tilde{A}_2^{(2^+)2\rm SF}}{(1 - r_{\rm LR}u)^p} \\ \hat{a}_Q^2 &= C_Q A \tilde{a}_A^2 + 2\tilde{a}_A \tilde{a}_B + C_Q B \tilde{a}_B^2 \quad \text{Spin-spin} \end{split}$$

$$\begin{split} H_{\rm EOB} &= M \sqrt{1 + 2\nu(\hat{H}_{\rm eff} - 1)}, \ \hat{H}_{\rm eff} = \sqrt{p_{r_*}^2 + \hat{A}(p_c) \left(1 + \frac{p_{\varphi}^2}{r_c^2} + Q(r_c, p_{r_*})\right)} + p_{\varphi}(G_{\hat{S}} \hat{S} + G_{\hat{S}_*} \hat{S}_*) \\ \dot{p}_{\varphi} &= \mathcal{F}_{\varphi} \end{split}$$

$$\begin{split} h_{\ell m} &= h_{\ell m}^{(N,\epsilon)} \hat{S}_{\rm eff}^{(\epsilon)} \hat{h}_{\ell m}^{\rm tail} f_{\ell m} \hat{h}_{\ell m}^{\rm NQC} + \hat{h}_{\ell m}^N \hat{h}_{\ell m}^T \end{split}$$



#### Matter effects



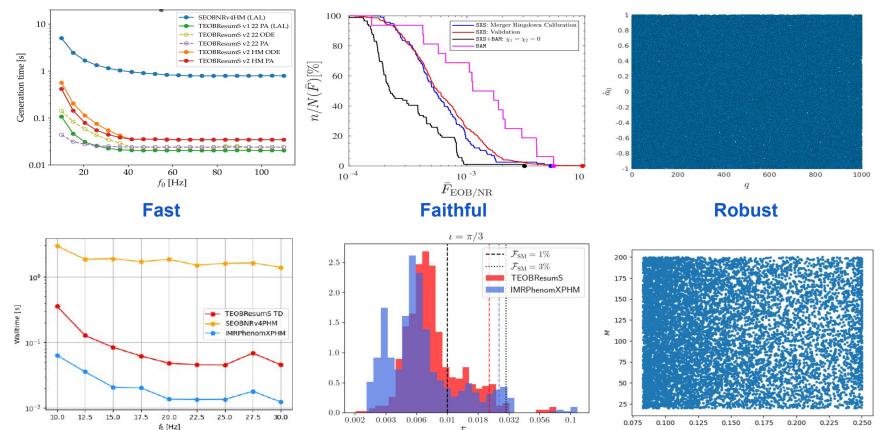
(explicit) NR info

Quasi-circular EOB model for generic-spins [2111.03675] :

- BNS [<u>1812.02744,1812.07923</u>]:
  - □ Electric ell=2,..,8 interactions, magnetic ell=2
  - GSF resummation ell=2,3
  - □ f-mode resonance model
  - Spin-quadrupole terms included (@NNLO)
  - □ <u>Higher modes (up to ell = 8)</u>
  - □ Frequency Domain model (SPA)
- BBH [2001.09082, 2104.07533]:
  - □ Inspiral-Merger-Ringdown
  - Higher modes, with MR up to ell=5
  - □ NR-informed NQCs and merger/ringdown
  - High-consistency between radiation reaction and waveform

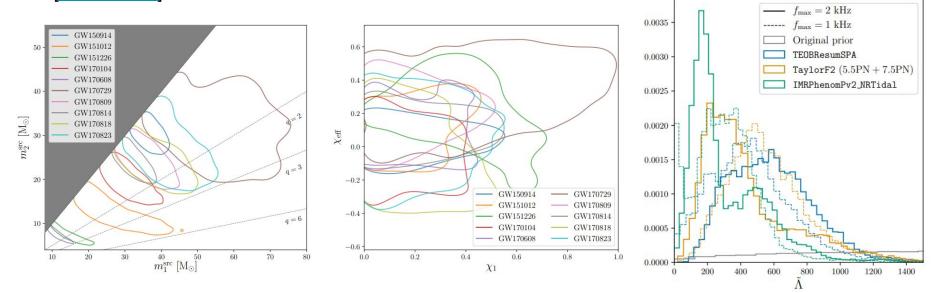


**Takeaway**: TEOBResumS is one of the most physically complete models available!



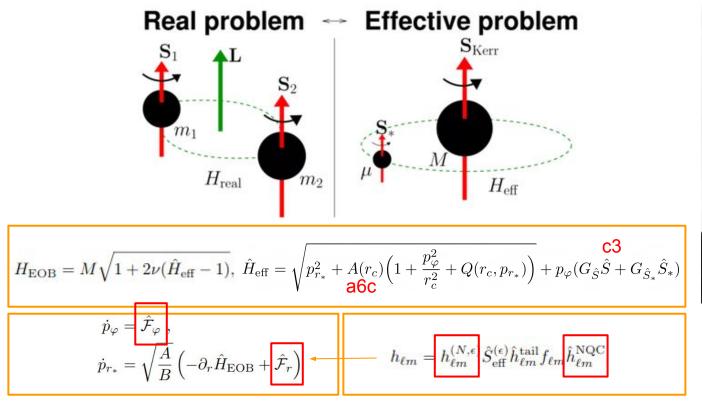
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2102.00017



Fast enough to be used in real PE!

#### TEOBResumS – DALI'





Also, new initial conditions!

#### TEOBResumS – DALI'

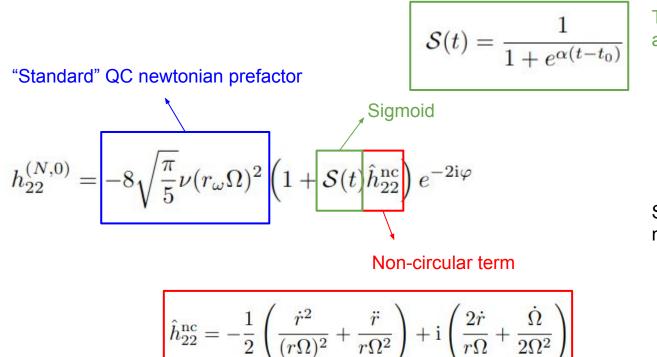
EOB model for generic orbit, <u>aligned spins</u> :

#### BNS:

- □ All features of GIOTTO, except...
- □ Frequency Domain model (SPA)
- **BBH** [2001.11736 , 2108.02043] :
  - □ All features of GIOTTO, but...
  - □ New treatment of dissipative terms
  - Different initial conditions
  - □ New determinations of a6c and c3



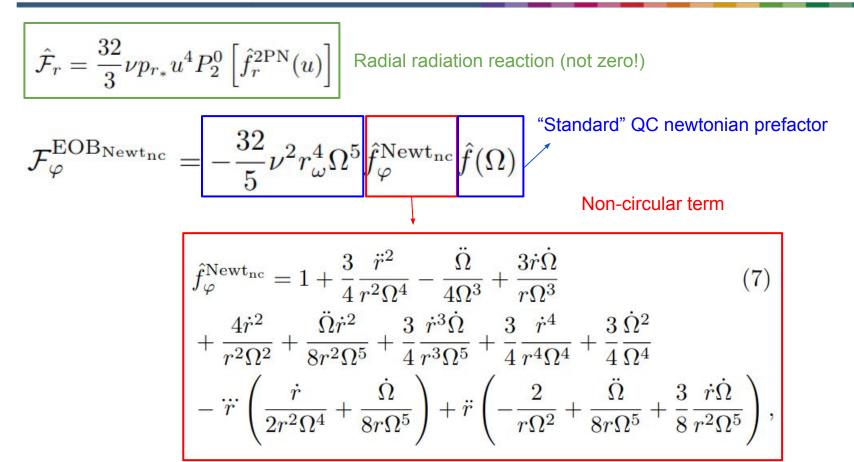
#### **TEOBResumS – DALI': Waveforms**



The same sigmoid activates the NQCs

Similar expression for other multipoles with (I,m) > (2,2)

#### **TEOBResumS – DALI': Fluxes**



#### **TEOBResumS – DALI': Initial conditions**

$$(e_0, \Omega_0, \xi_0) \to (r_0, p_{\varphi}^0, p_{r_*}^0)$$

• Assume:

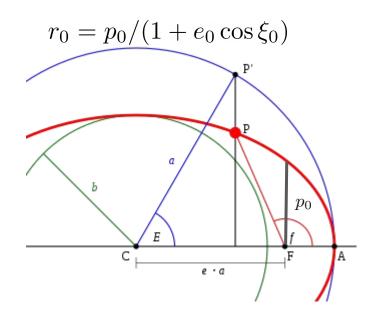
$$r_0 = p_0/(1+e_0)$$
 or  $r_0 = p_0/(1-e_0)$ 

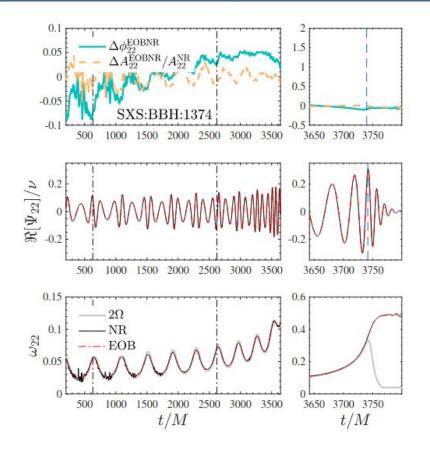
• Solve (numerically)

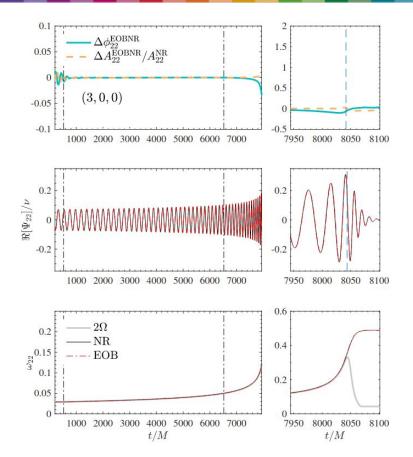
 $\partial_{p_{\varphi}} H(r_0(p_0), j_0(p_0), p_{r_*} = 0) = \Omega_0$  $\hat{H}^0_{\text{eff}}(p_0, j_0, \xi = 0) = \hat{H}^0_{\text{eff}}(p_0, j_0, \xi = \pi)$ 

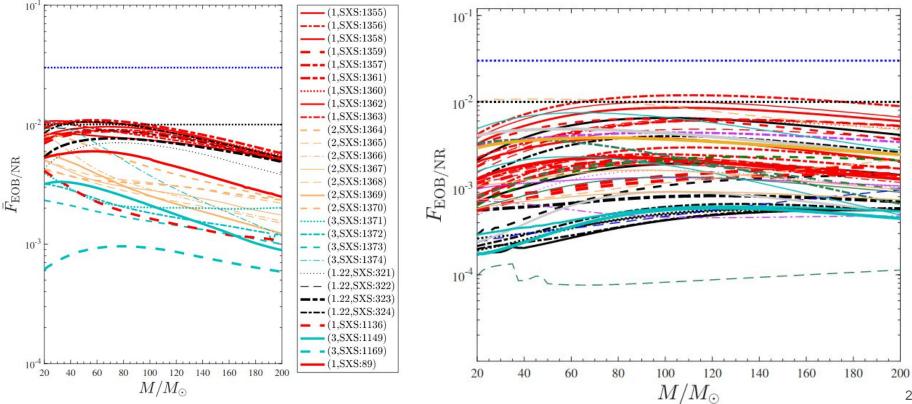
This procedure works, but is only adiabatic! (Giotto: 2PA)

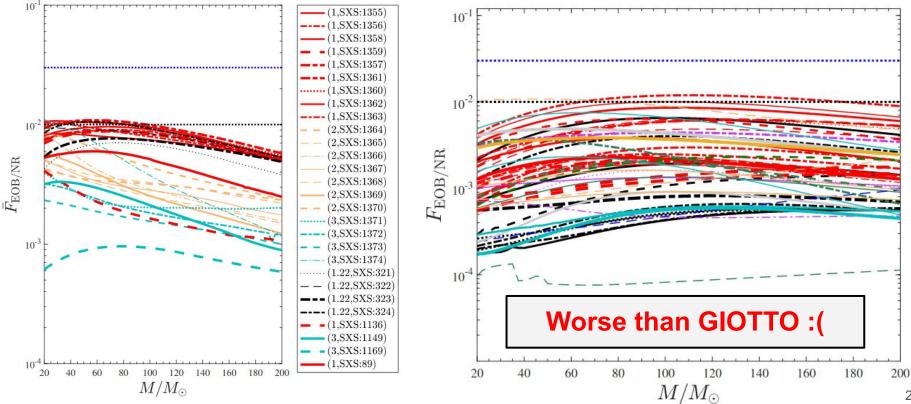
Alternative way of giving initial conditions: via  $(p_{\varphi}, E_{EOB}, r_0) \rightarrow$  Works for scatterings and encounters!

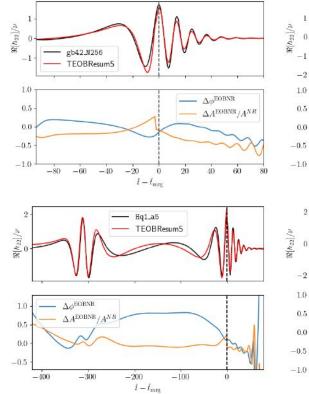


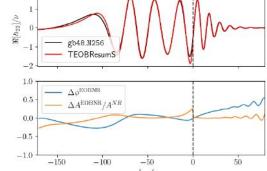












Hq1\_b5

TEOBResumS

 $\Delta \phi^{\text{EOBNR}}$ 

-60

 $\Delta A^{\text{EOBNR}} / A^{NR}$ 

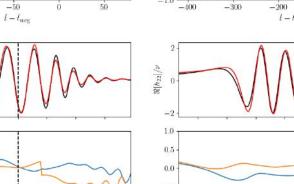
-40

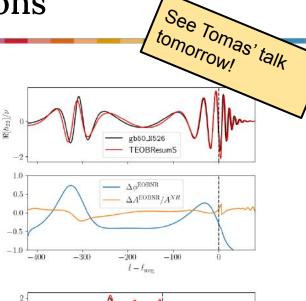
-20

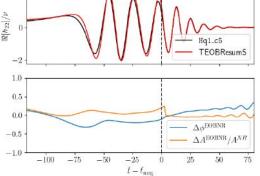
 $t - t_{\rm mrg}$ 

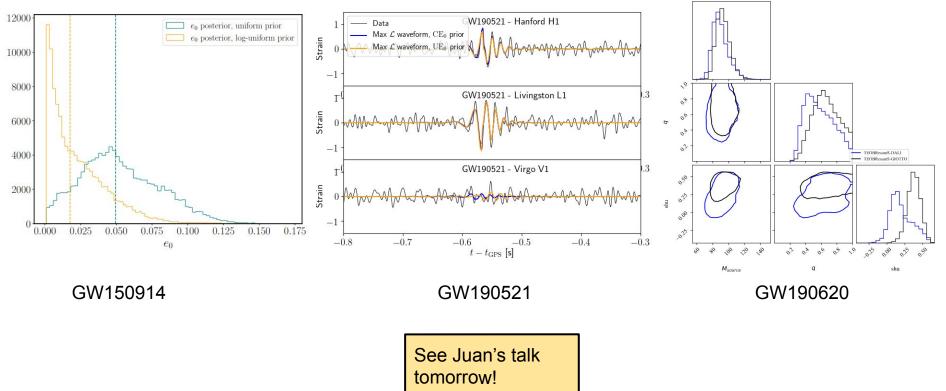
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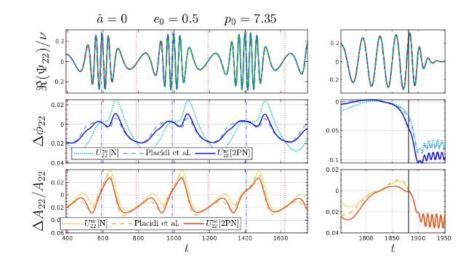


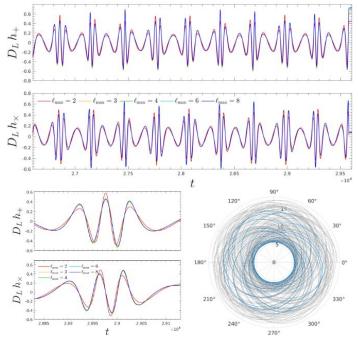




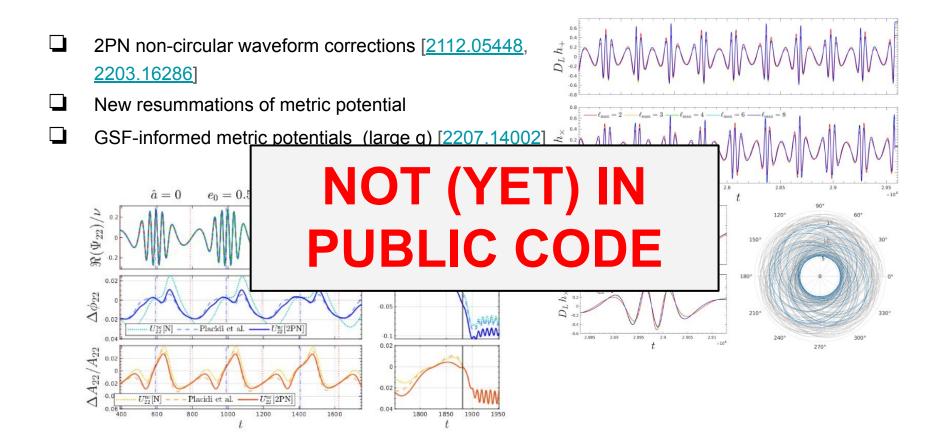
#### TEOBResumS – DALI': future perspective

- 2PN non-circular waveform corrections [2112.05448, 2203.16286]
- New resummations of metric potential
- GSF-informed metric potentials (large q) [2207.14002]

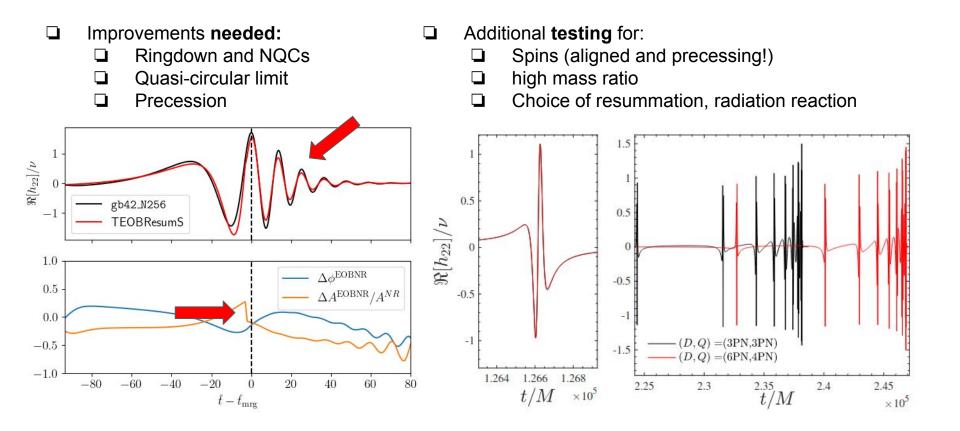




#### TEOBResumS – DALI': future perspective



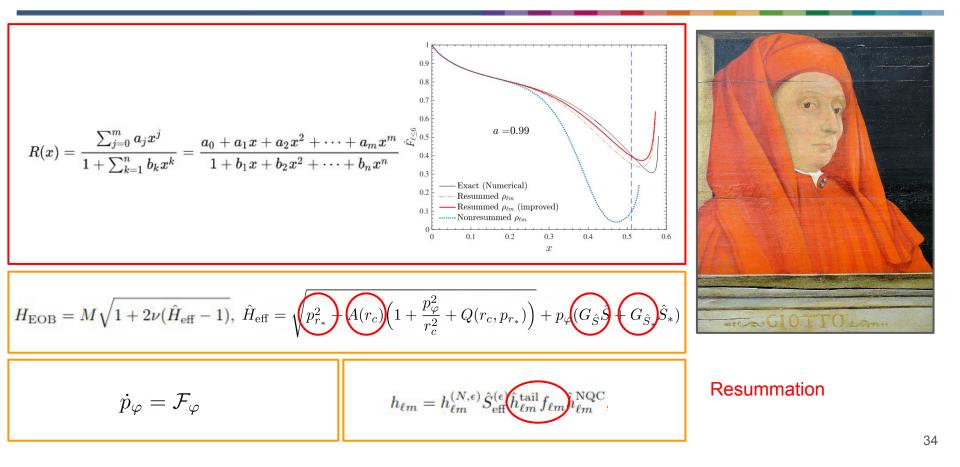
### TEOBResumS – DALI': future perspective





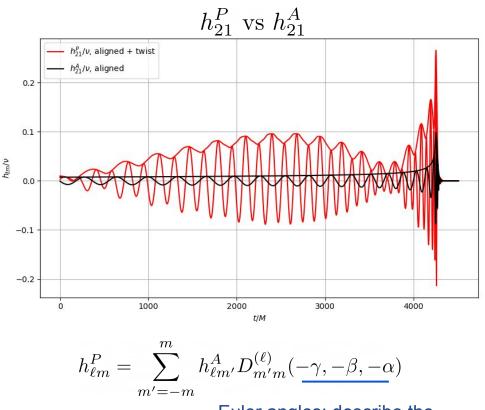
- TEOBResumS is an advanced EOB waveform approximant which is fast, faithful, and robust
- Both GIOTTO and DALI' are publicly available <u>here</u>, and have been used in real PE of LIGO-Virgo sources
- Physical completeness (BBH, BNS, BHNS, eccentricity, precession, bound and unbound dynamics) is at reach!

# Supplementary Slides



#### Precession

- If spins are not (mis)aligned to L → Orbital plane is not fixed
- In the <u>non inertial</u> frame which "follows" the precession of L (co-precessing frame), the WFs can be approximated with aligned spins waveforms
- Precessing waveforms in an inertial frame can be obtained <u>twisting</u> co-precessing (aligned) WFs



Euler angles: describe the evolution of L w.r.t inertial frame

$$\hat{h}_{\ell m}^{\rm NQC} = \left(1 + a_1^{\ell m} n_1^{\ell m} + a_2^{\ell m} n_2^{\ell m}\right) e^{i\left(b_1^{\ell m} n_3^{\ell m} + b_2^{\ell m} n_4^{\ell m}\right)},$$

#### Define then

$$t_{\rm NQC}^{\rm EOB} \equiv t_{\Omega}^{\rm peak} - \Delta t_{\rm NQC} \qquad t_{\rm NQC}^{\rm EOB} \leftrightarrow t_{\rm NQC}^{\rm NR} \equiv t_{A_{22}^{\rm max}}^{\rm NR} + 2. \qquad t_{\rm NQC-\ell m}^{\rm NR} \equiv t_{A_{\ell m}^{\rm max}}^{\rm NR} + 2. \qquad t_{A_{22}^{\rm max}}^{\rm EOB} \equiv t_{\rm NQC}^{\rm EOB} - 2,$$

With these definitions, one can solve the following system of equations to compute a1, a2, b1, b2

$$\begin{aligned} A_{\ell m}^{\rm EOB} \left( t_{\rm NQC}^{\rm EOB} + \Delta t_{\ell m}^{\rm NR} \right) &= A_{\ell m}^{\rm NR} \left( t_{\rm NQC-\ell m}^{\rm NR} \right), \\ \dot{A}_{\ell m}^{\rm EOB} \left( t_{\rm NQC,}^{\rm EOB} + \Delta t_{\ell m}^{\rm NR} \right) &= \dot{A}_{\ell m}^{\rm NR} \left( t_{\rm NQC-\ell m}^{\rm NR} \right), \\ \omega_{\ell m}^{\rm EOB} \left( t_{\rm NQC}^{\rm EOB} + \Delta t_{\ell m}^{\rm NR} \right) &= \omega_{\ell m}^{\rm NR} \left( t_{\rm NQC-\ell m}^{\rm NR} \right), \\ \dot{\omega}_{\ell m}^{\rm EOB} \left( t_{\rm NQC}^{\rm EOB} + \Delta t_{\ell m}^{\rm NR} \right) &= \dot{\omega}_{\ell m}^{\rm NR} \left( t_{\rm NQC-\ell m}^{\rm NR} \right), \end{aligned}$$

Note: NQCs affect the waveform amplitude, and therefore the radiation reaction  $\rightarrow$  different NQCs, different inspiral!

After computing them once, one needs to use those values and rerun the entire model. Iterative procedure.

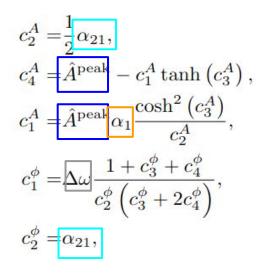
### Ringdown template

 $\bar{h}(\tau) = A_{\bar{h}}(\tau) e^{\mathrm{i}\phi_{\bar{h}}(\tau)},$ 

$$\begin{aligned} A_{\bar{h}}(\tau) = & c_1^A \tanh\left(c_2^A \tau + \overline{c_3^A}\right) + c_4^A, \\ \phi_{\bar{h}}(\tau) = & -c_1^{\phi} \ln\left(\frac{1 + c_3^{\phi} e^{-c_2^{\phi} \tau} + \overline{c_4^{\phi} e^{-2c_2^{\phi} \tau}}}{1 + c_3^{\phi} + \overline{c_4^{\phi}}}\right) \end{aligned}$$

Each multipole depends on 10 parameters fitted from NR:

- Two quantities at merger (A\_peak, omega\_peak)
- QNMs fundamental complex frequency (alpha\_1 + i omega\_1) and alpha of the first overtone (alpha\_2) → alpha\_21 = alpha\_2 - alpha\_1. These depend on the final mass and spin of the remnant BH
- Three model parameters (c3A, c3phi, c4phi)
- Remnant mass and spin



#### **TEOBResumS vs GW190521**

