

Physics Beyond the Standard Model

Flavour and Dark Photons

Jordi Folch Eguren

w/ E. Stamou (TUD), M. Tabet (TUD), R. Ziegler (KIT)

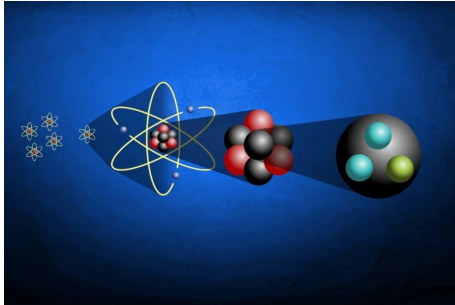
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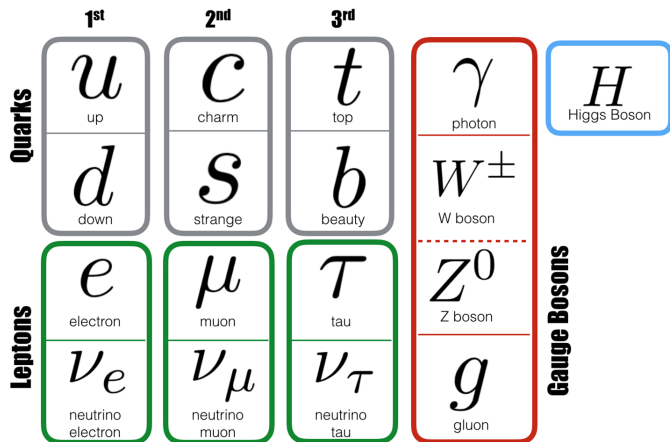
Part I

Standard Model, Flavour and Beyond

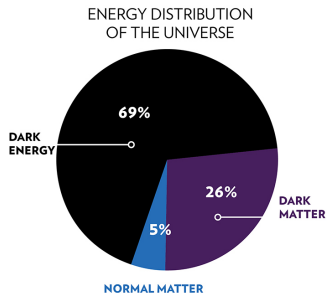
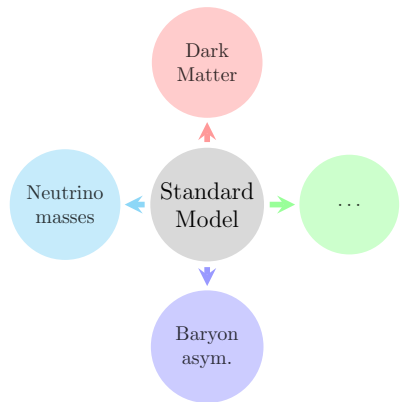


Standard Model: Flavour

Quarks come in three copies (generations), and 6 flavours: $\underbrace{u, c, t}_{u^i}, \underbrace{d, s, b}_{d^i}$.

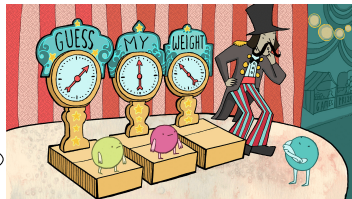


Beyond the Standard Model (BSM)



- Seesaw models
- SUSY
- ...

- WIMPs
- Dark Photon



How can flavour physics help?

Flavour physics allows to probe new phenomena → BSM searches!

But why?

How can flavour physics help?

Flavour physics allows to probe new phenomena → BSM searches!

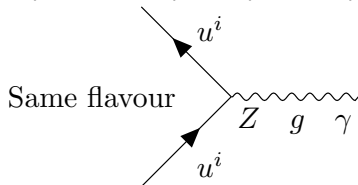
But why?

Because of the **suppression** of flavour-changing neutral currents (**FCNCs**)
in the SM

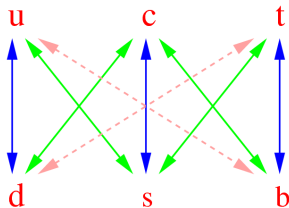
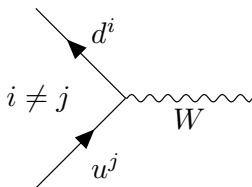
SM and flavour changing currents

There are neutral (γ, g, Z) and charged (W^\pm) interactions with quarks

- (γ, g, Z) do not change flavour \rightarrow **no** flavour changing **neutral** currents (FCNCs) $u^i \equiv (u, c, t), d^i \equiv (d, s, b)$

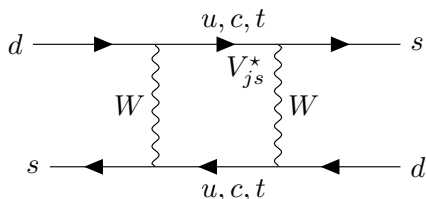


- W^\pm bosons change flavour \rightarrow flavour changing **charged** currents

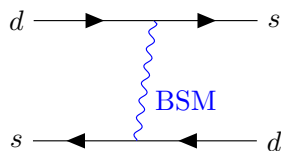


SM, FCNCs and BSM

FCNCs can occur at loop level



New BSM field with FCNCs



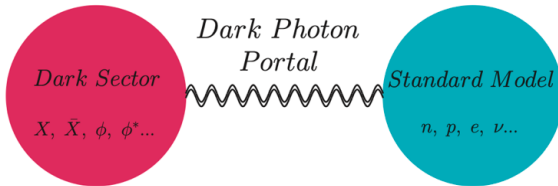
FCNC processes are very suppressed in the SM:

- Arise at loop level ($\sim 1/16\pi^2$)
- Smallness of CKM elements ($V_{ij} \ll 1, i \neq j$)
- GIM mechanism ($\sim (m_u - m_c)^2 / M_W^2; V_{td}, V_{ts} \ll 1$)

FCNCs are a good probe for BSM physics

Part II

Dark Photon



The Dark Photon (DP) with kinetic mixing

Consider adding a **neutral spin 1** field (A'_μ) to QED (A_μ)

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}^2 - \frac{1}{4}F'_{\mu\nu}{}^2 + eJ_\mu A^\mu + e'J'_\mu A'^\mu + \frac{m_{\gamma'}^2}{2}A'_\mu A'^\mu$$

J_μ SM matter

J'_μ dark sector (DS) matter

The Dark Photon (DP) with kinetic mixing

Consider adding a **neutral spin 1** field (A'_μ) to QED (A_μ)

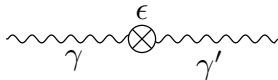
$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}^2 - \frac{1}{4}F'_{\mu\nu}{}^2 + eJ_\mu A^\mu + e'J'_\mu A'^\mu + \frac{m_{\gamma'}^2}{2}A'_\mu A'^\mu$$

J_μ SM matter

J'_μ dark sector (DS) matter

We can also write a kinetic mixing term!

$$\mathcal{L}_{KM} = -\frac{\epsilon}{2}F^{\mu\nu}F'_{\mu\nu}, \quad \epsilon \ll 1$$



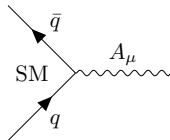
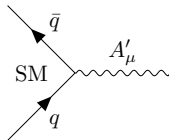
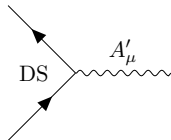
Due to kinetic mixing the DP can interact with SM matter

DP and SM matter

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}^2 - \frac{1}{4}F'_{\mu\nu}{}^2 - \frac{\epsilon}{2}F^{\mu\nu}F'_{\mu\nu} + eJ_{\mu}A^{\mu} + e'J'_{\mu}A'^{\mu} + \frac{m_{\gamma'}^2}{2}A'_{\mu}A'^{\mu}$$

Diagonalisation of the kinetic mixing

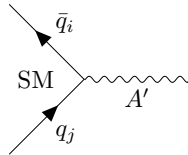
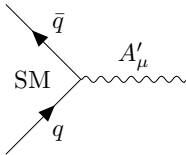
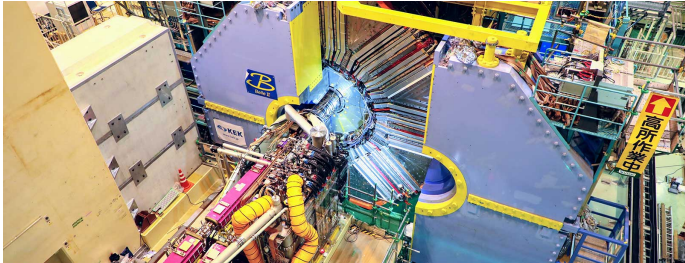
$$\mathcal{L}_{\text{int}} = (e'J'_{\mu} - \epsilon eJ_{\mu})A'^{\mu} + eJ_{\mu}A^{\mu}$$



Minimal model with kinetic mixing and no flavour-changing couplings

Part III

Flavour and Dark Photon



We considered kinetic mixing, now we go beyond!

Why flavour and the Dark Photon

- SM has non-trivial flavour interactions, DP could behave similarly
- Rich phenomenology with experimental searches at colliders (2-body decays)

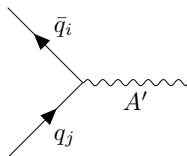
- Belle II ($B \rightarrow \rho$)
- NA62 ($K \rightarrow \pi$)
- BaBar ($B \rightarrow \pi$)
- ...



- Future upgrades and searches: KOTO, KLEVER, BESIII, ...

⇒ Potential discovery via FCNCs

Flavour-changing couplings with the Dark Photon



$$\mathcal{L} = \mathcal{L}_D + \mathcal{L}_V$$

$$\mathcal{L}_D = \frac{1}{\Lambda} \bar{q}_i \sigma^{\mu\nu} \left(C_{ij}^D + i\gamma_5 C_{ij}^{5D} \right) q_j F'_{\mu\nu} \quad \text{Dipole (dim-5)}$$

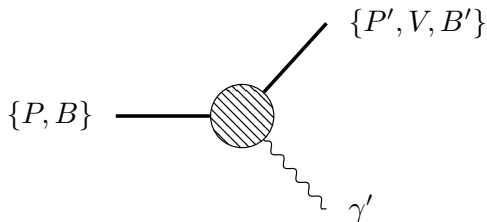
$$\mathcal{L}_V = \left(\frac{m_{\gamma'}}{\Lambda} \right)^2 \bar{q}_i \left(C_{ij} + \gamma_5 C_{ij}^5 \right) q_j A' \quad \text{Vector (dim-4)}$$

Objective

Constrain $\{C_{ij}^D, C_{ij}^{5D}, C_{ij}, C_{ij}^5\}$ with flavour-changing 2-body decays
($m_{\gamma'} \ll m_{EW}$; EFT)

2-body decays

$P^{(\prime)} \equiv$ pseudoscalar, $V \equiv$ vector, $B^{(\prime)} \equiv$ baryon

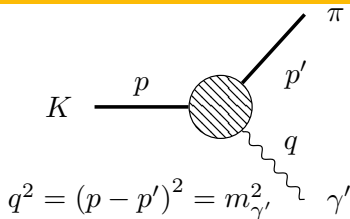


Type	Example	\mathbb{C}_{ij}^D	\mathbb{C}_{ij}^{5D}	\mathbb{C}_{ij}	\mathbb{C}_{ij}^5	Sector	Example
$P \rightarrow P' \gamma'$	$K \rightarrow \pi \gamma'$	✓	✗	✓	✗	<i>sd</i>	$K \rightarrow \pi \gamma'$
$P \rightarrow V \gamma'$	$B \rightarrow K^* \gamma'$	✓	✓	✓	✓	<i>bs</i>	$B \rightarrow K^* \gamma'$
$B \rightarrow B' \gamma'$	$\Sigma \rightarrow p \gamma'$	✓	✓	✓	✓	<i>bd</i>	$B \rightarrow \pi \gamma'$
						<i>cu</i>	$D \rightarrow \pi \gamma'$

Form factors (FFs): $K \rightarrow \pi\gamma'$ example

$$\mathcal{L}_V = \left(\frac{m_{\gamma'}}{\Lambda}\right)^2 \bar{q}_i A' C_{ij} q_j$$

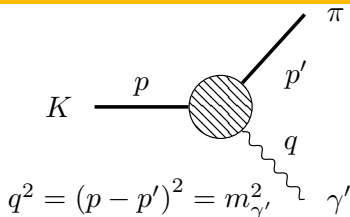
$$\mathcal{M} = \langle \gamma' \pi(p') | \mathcal{L}_V | K(p) \rangle$$



Form factors (FFs): $K \rightarrow \pi \gamma'$ example

$$\mathcal{L}_V = \left(\frac{m_{\gamma'}}{\Lambda}\right)^2 \bar{q}_i A' C_{ij} q_j$$

$$\mathcal{M} = \langle \gamma' \pi(p') | \mathcal{L}_V | K(p) \rangle$$



- $\mathcal{M} = \left(\frac{m_{\gamma'}}{\Lambda}\right)^2 C_{ij} \langle \pi(p') \gamma' | \bar{q}_i A'_\mu \gamma^\mu q_j | K(p) \rangle$

$\langle \pi(p') | \bar{q}_i \gamma^\mu q_j | K(p) \rangle \sim f(m_{\gamma'}^2)$ is the form factor

$$\langle \pi(p') | \bar{q}_i \gamma^\mu q_j | K(p) \rangle = (p + p')^\mu \mathbf{f}_{ij}^+ (m_{\gamma'}^2) + (p - p')^\mu \mathbf{f}_{ij}^- (m_{\gamma'}^2)$$

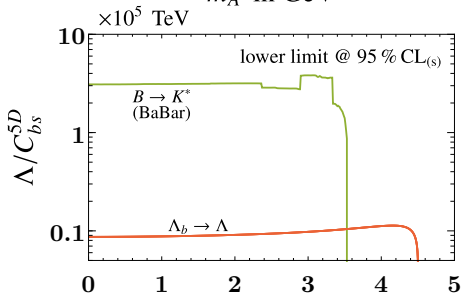
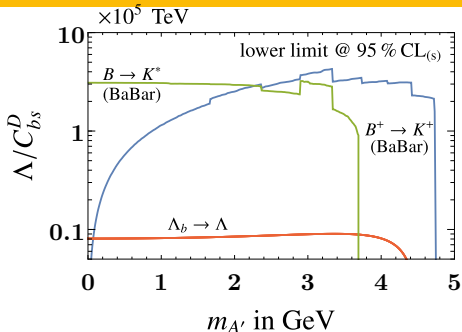
Mass dependence and non-perturbative effects in FFs

Equivalently for $\langle B(p') | \bar{q}_i \gamma_\mu q_j | B(p) \rangle$, $\langle V(p', \epsilon_\alpha) | \bar{q}_i \sigma_{\mu\nu} \gamma_5 q_j | P(p) \rangle$, etc

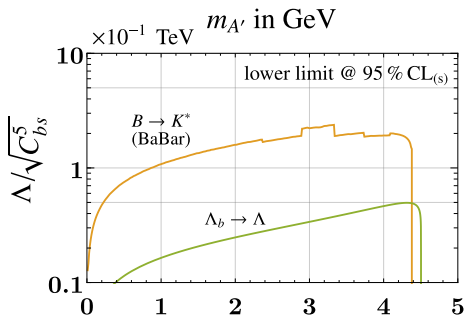
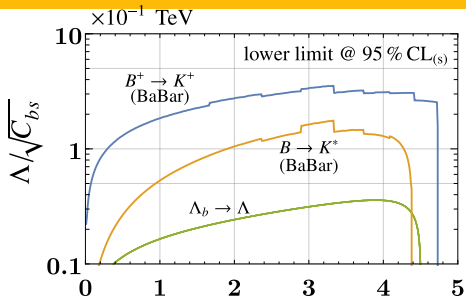
Part IV

Constraining Dark Photon FCNCs

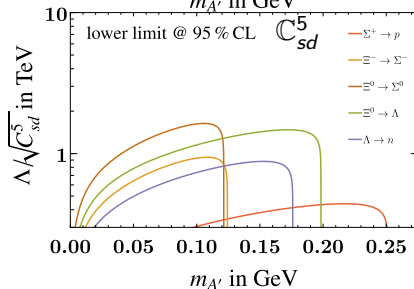
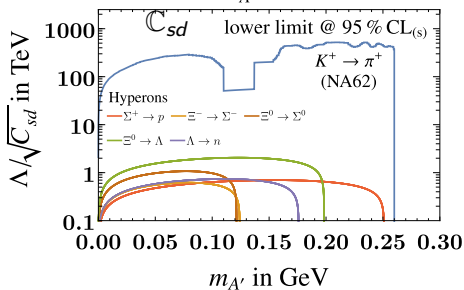
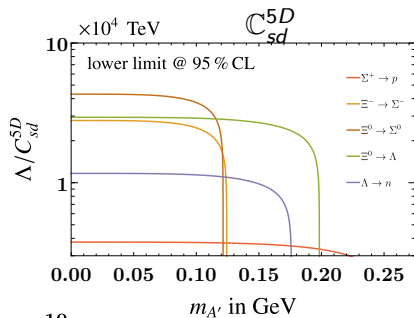
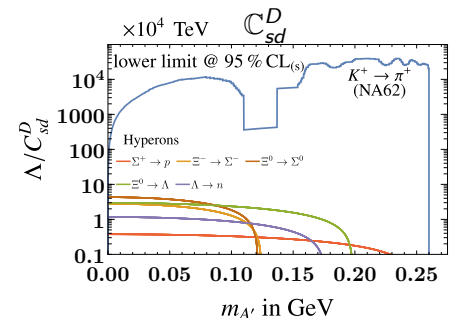
Constraints bs sector Dipole $\mathbb{C}_{bs}^D, \mathbb{C}_{bs}^{5D}$



Constraints bs sector Vector $\mathbb{C}_{bs}, \mathbb{C}_{bs}^5$



Constraints sd sector



Towards a global picture of FCNC bounds

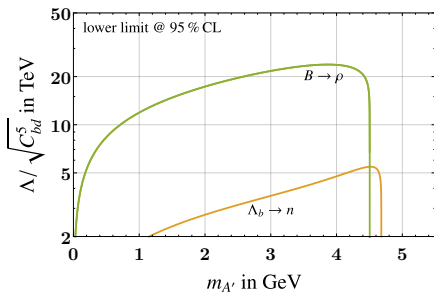
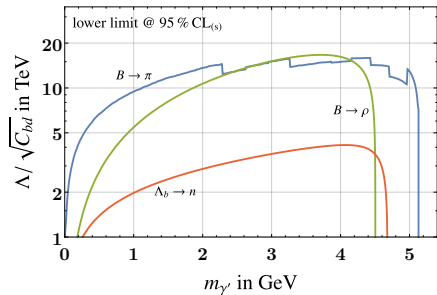
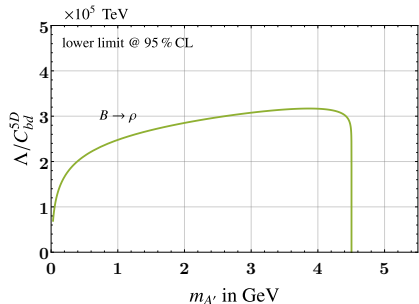
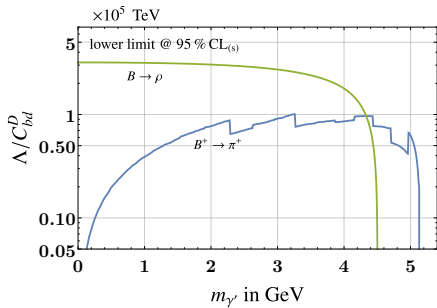
- Recast of data, experiments needed!
- DP mass dependence, with $m_{A'} \ll m_{EW}$
- $K \rightarrow \pi$ sets the strongest constraint (sd sector)
- Baryon decays are the least constraining but sometimes the only available (no current experimental searches, but planned ones).
- Two more sectors: bd and cu

Conclusions and outlook

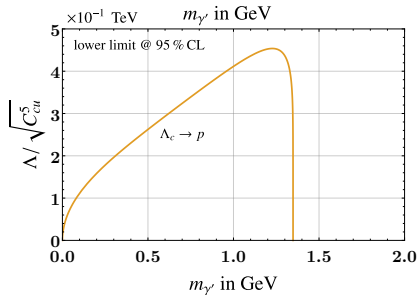
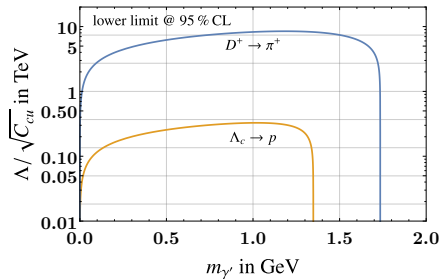
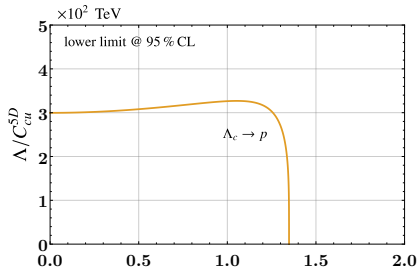
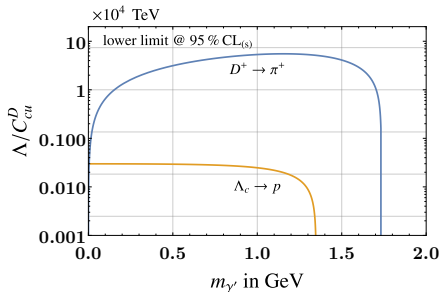
- SM must be extended, flavour gives us a guide
- Dark Photon is a minimal extension of the SM and a DM candidate
- Discovery potential via FCNCs searches: constrained flavour-violating couplings with 2-body decays
 - $P \rightarrow P'\gamma'$
 - $P \rightarrow V\gamma'$
 - $B \rightarrow B'\gamma'$
- Future:
 - Astrophysical and cosmological constraints
 - Dark sector matter

Backup slides

bd sector



cu sector



Dark Photon mass and gauge invariance

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} + \frac{m^2}{2}A_\mu A^\mu \quad \text{Proca theory}$$

The mass term $\frac{m^2}{2}A_\mu A^\mu$ is gauge invariant in a **hidden** manner

Consider doing a field redefinition $A_\mu \rightarrow A_\mu + \partial_\mu\phi$, then the theory is invariant under

$$\begin{aligned}A_\mu &\rightarrow A_\mu + \partial_\mu\Lambda \\ \phi &\rightarrow \phi - \Lambda\end{aligned}$$

This is known as the Stueckelberg procedure, which is nothing else than the **affine Higgs mechanism** (i.e. Higgs is decoupled)

DP can get mass through a Dark Higgs

Kinetic mixing with SM hypercharge boson

$$\mathcal{L} = \mathcal{L}_{EW} + \mathcal{L}_{Higgs} + \mathcal{L}_{KM}$$

Kinetic mixing term for Dark Photon and SM $U(1)_Y$ boson

$$\mathcal{L}_{KM} = -\frac{\epsilon}{2} B^{\mu\nu} F'_{\mu\nu}$$

Diagonalisation+SSB+gauge mass basis

$$\begin{pmatrix} B_\mu \\ W_\mu^3 \\ A'_\mu \end{pmatrix} = \begin{pmatrix} 1 & 0 & -\epsilon t \\ 0 & 1 & 0 \\ 0 & 0 & t \end{pmatrix} \begin{pmatrix} c_W & -s_W c_\xi & s_W s_\xi \\ s_W & c_W c_\xi & c_W s_\xi \\ 0 & s_\xi & c_\xi \end{pmatrix} \begin{pmatrix} A_\mu \\ Z_\mu \\ Z'_\mu \end{pmatrix}$$

$$\tan 2\xi = -\frac{2\eta s_W}{1 - s_W^2 \eta^2 - \delta} \quad \text{with} \quad t = 1/\sqrt{1 - \epsilon^2}, \quad \eta = \epsilon t, \quad \delta = m_{A'}^2/m_Z^2$$

Interactions

$$\mathcal{L}_D = \frac{1}{\Lambda} \bar{q}_i \sigma^{\mu\nu} \left(C_{ij}^D + i\gamma_5 C_{ij}^{5D} \right) q_j F'_{\mu\nu} \quad \text{Dipole}$$

- dim-5 operator, DP can be massive/massless

$$\mathcal{L}_V = \bar{q}_i A' \left(C_{ij} + \gamma_5 C_{ij}^5 \right) q_j \quad \text{Vector}$$

- dim-4 operator, DP **cannot** be massless, Ward identity is not satisfied ($p_\mu M^{\mu\nu} \neq 0$). Rescale $\mathcal{L}_V \rightarrow \left(\frac{m_{\gamma'}}{\Lambda}\right)^2 \mathcal{L}_V$ so $\mathcal{L}_V = 0$ if $m_{\gamma'} = 0$

	\mathcal{L}_D	\mathcal{L}_V
Massive	✓	✓
Massless	✓	✗

Massless DP couples to SM only through higher dimensional operators

Origin of flavour violating couplings

- \mathcal{L}_V comes from the interaction $A'_\mu J^\mu$

$$J^\mu = \sum_{ij} \bar{Q}^i Y_Q'^{ij} \gamma^\mu Q^j + \sum_{ij} \bar{u}_R^i Y_u'^{ij} \gamma^\mu u_R^j + \sum_{ij} \bar{d}_R^i Y_d'^{ij} \gamma^\mu d_R^j$$

Going to the Yukawa mass basis we infer:

FV couplings are induced if the hypercharges Y'_x are **not** universal

- \mathcal{L}_D comes from the interaction $\frac{1}{\Lambda^2} F'_{\mu\nu} J^{\mu\nu}$

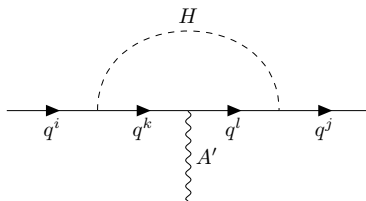
$$J^{\mu\nu} = \sum_{ij} \bar{Q}^i \tilde{H} C_u^{ij} \sigma^{\mu\nu} u_R^j + \sum_{ij} \bar{Q}^i H C_d^{ij} \sigma^{\mu\nu} d_R^j + \text{h.c.}$$

Going to the Yukawa mass basis we infer:

FV couplings are induced if couplings C_x are **not** aligned with SM Yukawas

Flavour-changing couplings from RGEs

FCNCs can be induced from the couplings RGEs (1310.4838v3)


$$\mu \frac{dC}{d\mu} = \gamma^T C$$

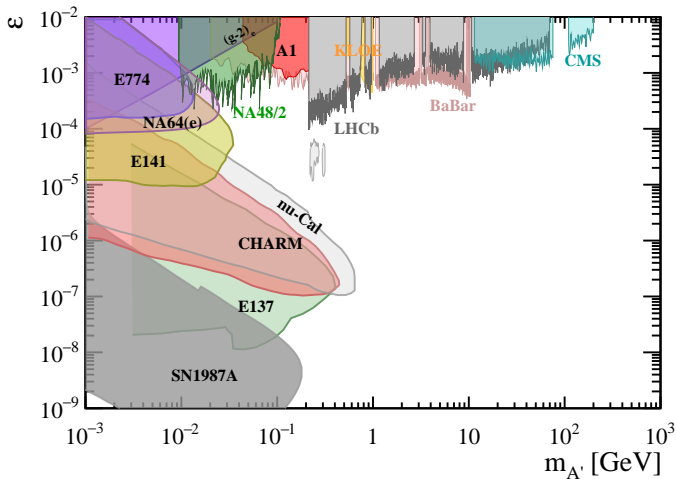
Taking into account 1-loop Yukawa corrections we find

Starting with flavour-diagonal interactions at a high scale Λ FCNCs are induced at the low scale μ

Top contributions yield

$$C_{ij}(\mu) \sim \delta_{ij} C_{ij}(\Lambda) + m_t^2 V_{tj} V_{ti}^* \log\left(\frac{\mu}{\Lambda}\right)$$

Constraints



Di-lepton searches (LHCb, NA48, BaBar, etc); Beam dump (NA64, E774 at Fermilab, etc); Supernova (1987A)