

Rare Processes and Precision Measurements

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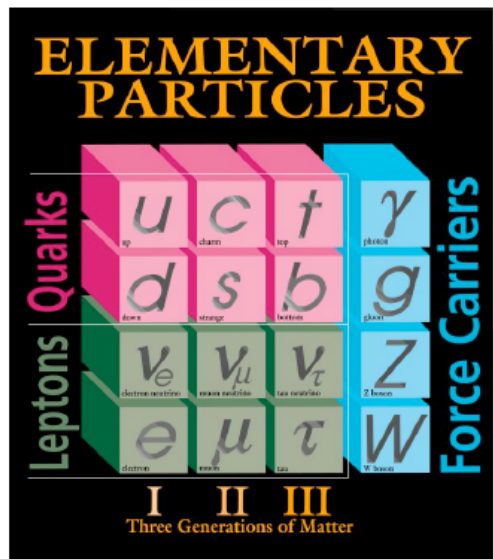
Outline

1. Introduction and Motivation
2. Testing the Standard Model with the anomalous magnetic moment of the Muon
3. Looking for dark matter with a Higgs-mixed Scalar
4. Conclusion and Outlook

1. Introduction and Motivation

1.1 The Standard Model

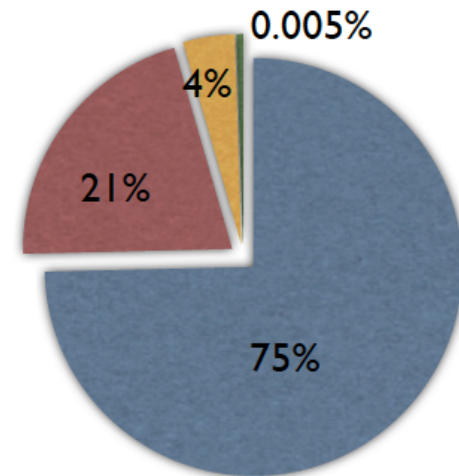
- The Standard Model, theory to describe electroweak and strong interactions very successful to explain experimental findings
- But hints for New Physics (SM: theory valid up to a certain energy scale Λ)
 - Neutrino Oscillations: $m_\nu \neq 0$
 - The Standard Model fails to explain the observed cosmology:



+ Higgs boson

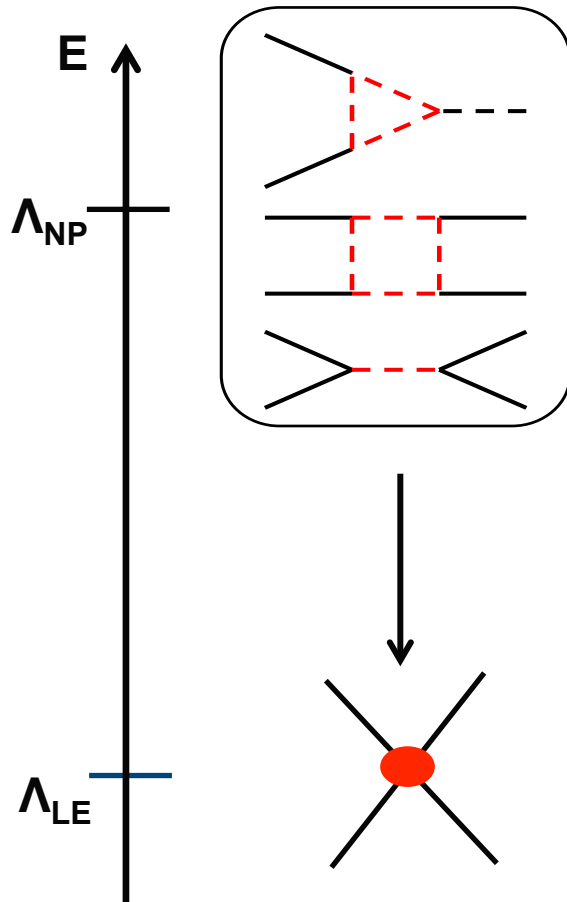


Energy density budget



● Dark Energy ● Dark Matter
● Baryons ● Radiation

1.2 Indirect searches of New Physics



- 2 ways to look for new physics:
Directly by producing new states at very high energy at colliders → LHC

Indirectly by precision measurements:

- Kaon physics: $\frac{s\bar{d}s\bar{d}}{\Lambda^2} \Rightarrow \Lambda \gtrsim 10^5 \text{ TeV}$
[ϵ_K]

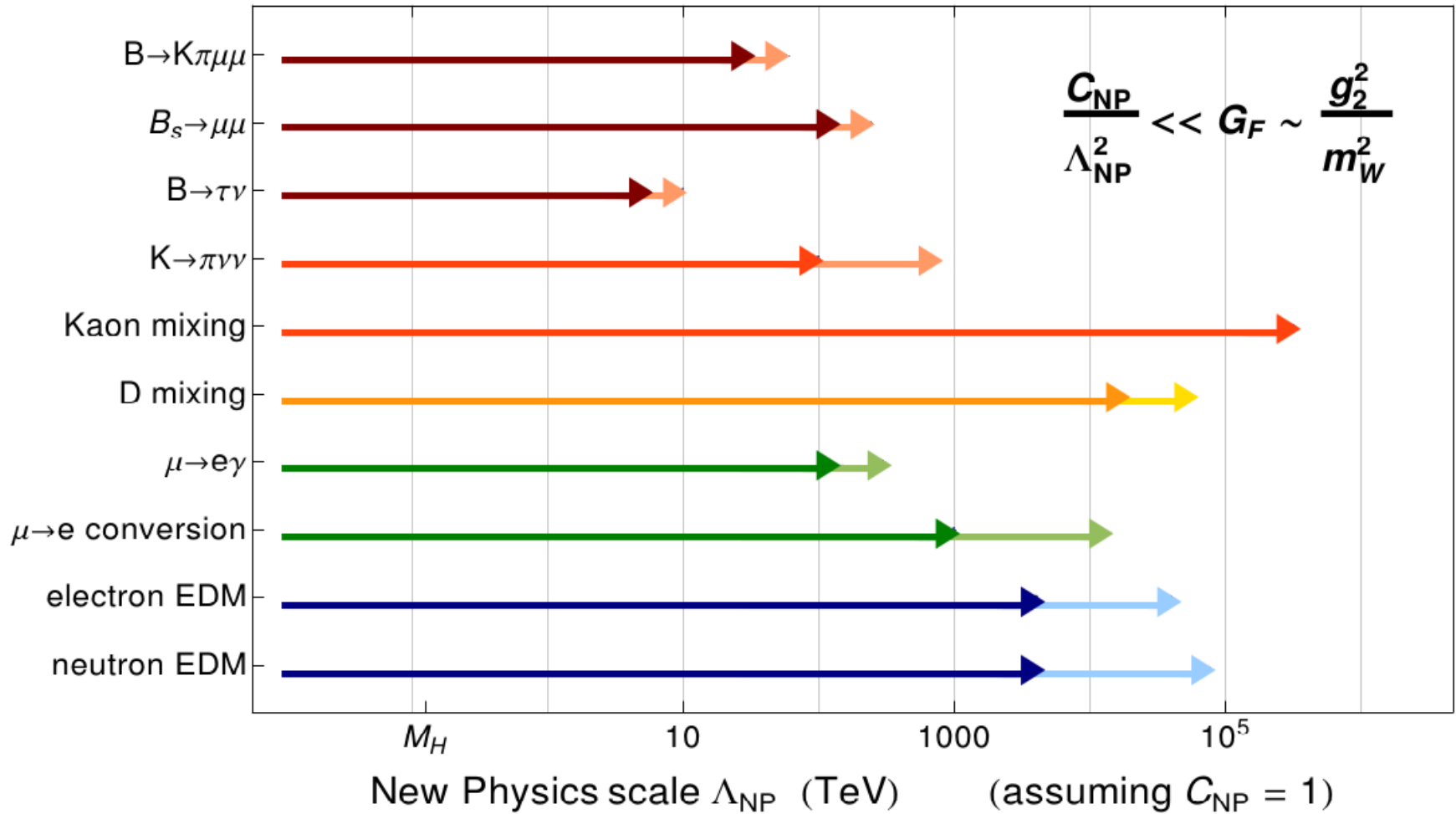
- Charged Leptons: $\frac{\mu\bar{e}f\bar{f}}{\Lambda^2} \Rightarrow \Lambda \gtrsim 10^4 \text{ TeV}$
[$\mu \rightarrow e\gamma$]

- At low energy: lots of experiments e.g., *MEG, COMET, Mu2e, E-969, NA62, FNAL g-2, BaBar, Belle-II, BESIII, LHCb, NA62, JLab* → huge improvements on measurements and bounds obtained and more expected
- For some modes accurate calculations of hadronic uncertainties essential

1.3 New Physics and Flavour sector

- Sensitivity of New Physics in the flavour sector

W. Altmannshofer

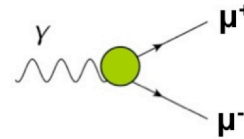


2. Testing the Standard Model with $g-2$ of the Muon

2.1 Anomalous magnetic moment of the muon

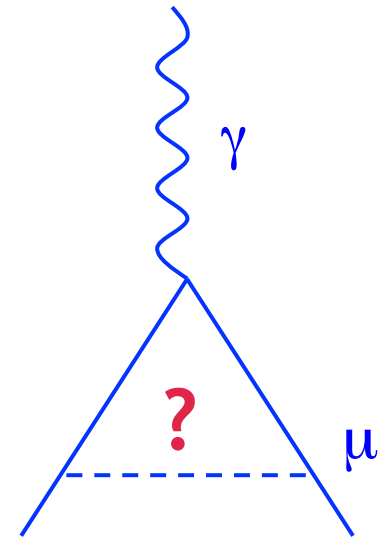
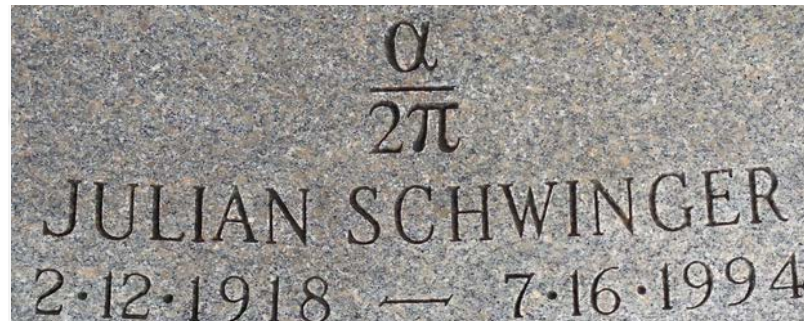
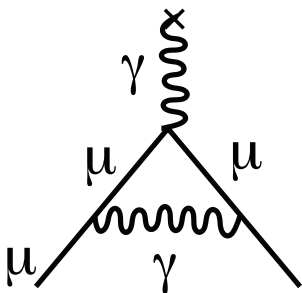
$$a_\mu = \frac{(g - 2)_\mu}{2}$$

Anomalous
magnetic moment



- The gyromagnetic factor of the muon is modified by loop contribution
- Predicted by Dirac to be 2
- Schwinger computed the first order correction

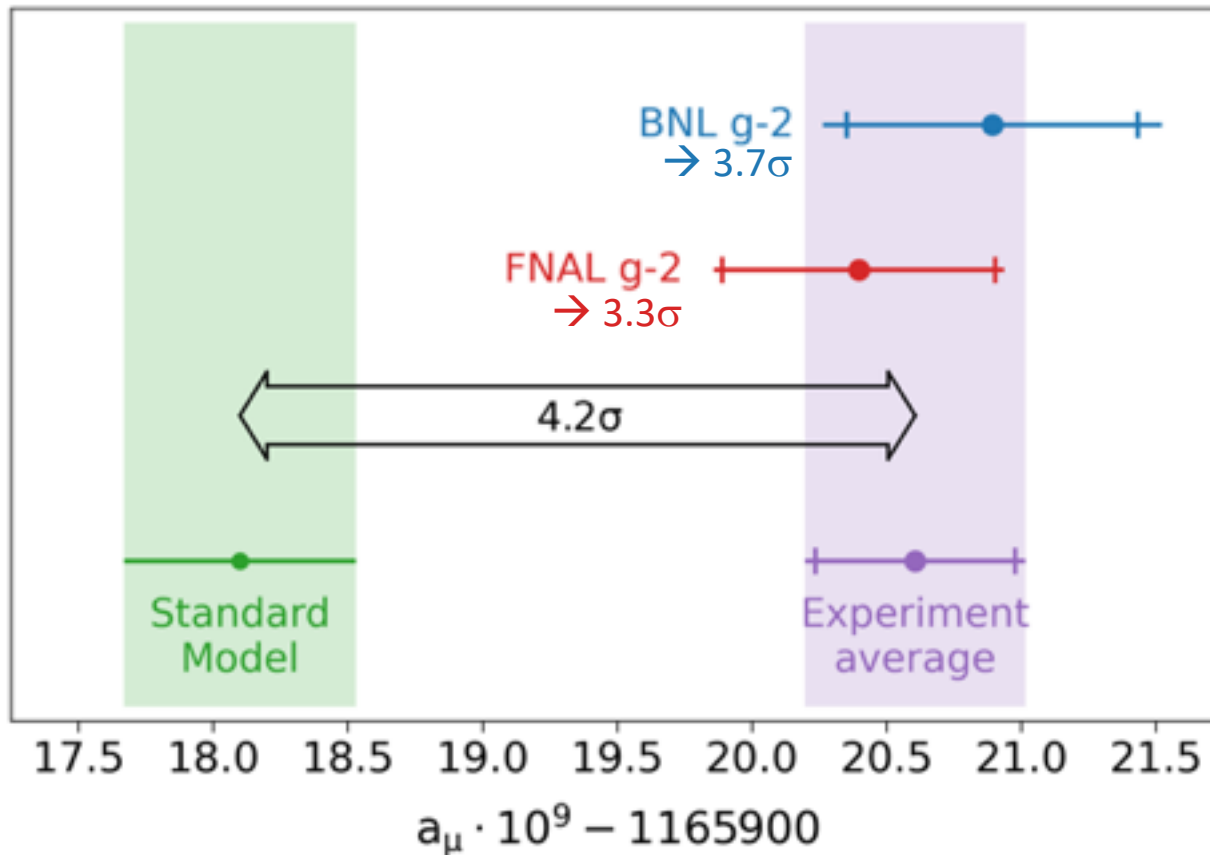
QED



2.1 Anomalous magnetic moment of the muon

FNAL g-2
Chris Polly'21

- In 2021 $a_\mu(\text{SM}) = 0.00116591810(43) \rightarrow 368 \text{ ppb}$



- Individual tension with SM

– BNL: 3.7σ

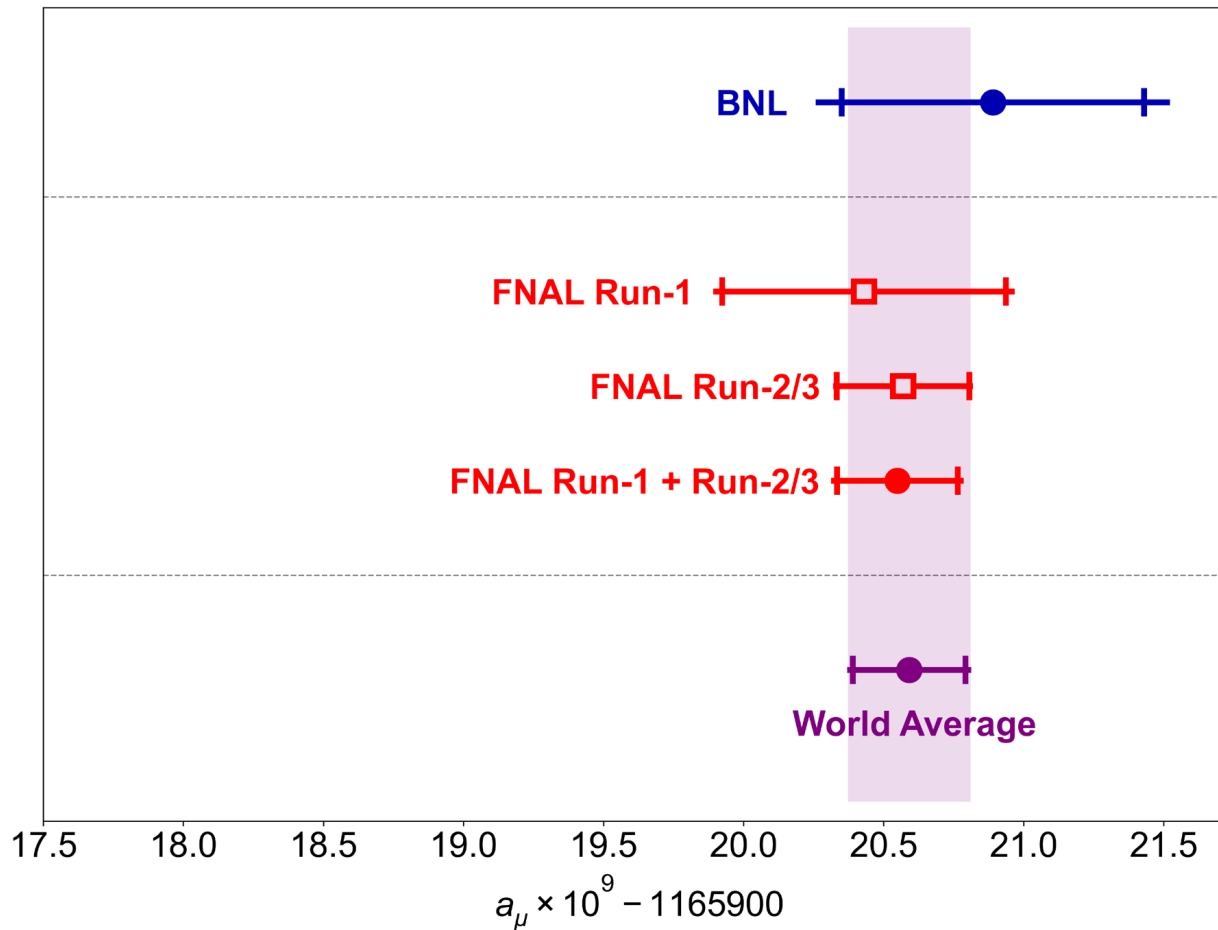
– FNAL: 3.3σ

$$a_\mu(\text{Exp}) - a_\mu(\text{SM}) = 0.00000000251(59) \rightarrow 4.2\sigma$$

2.1 Anomalous magnetic moment of the muon

FNAL g-2
James Mott'23

- In 2023 $a_\mu(\text{FNAL}) = 0.00\ 116\ 592\ 055(24)$ [203 ppb]



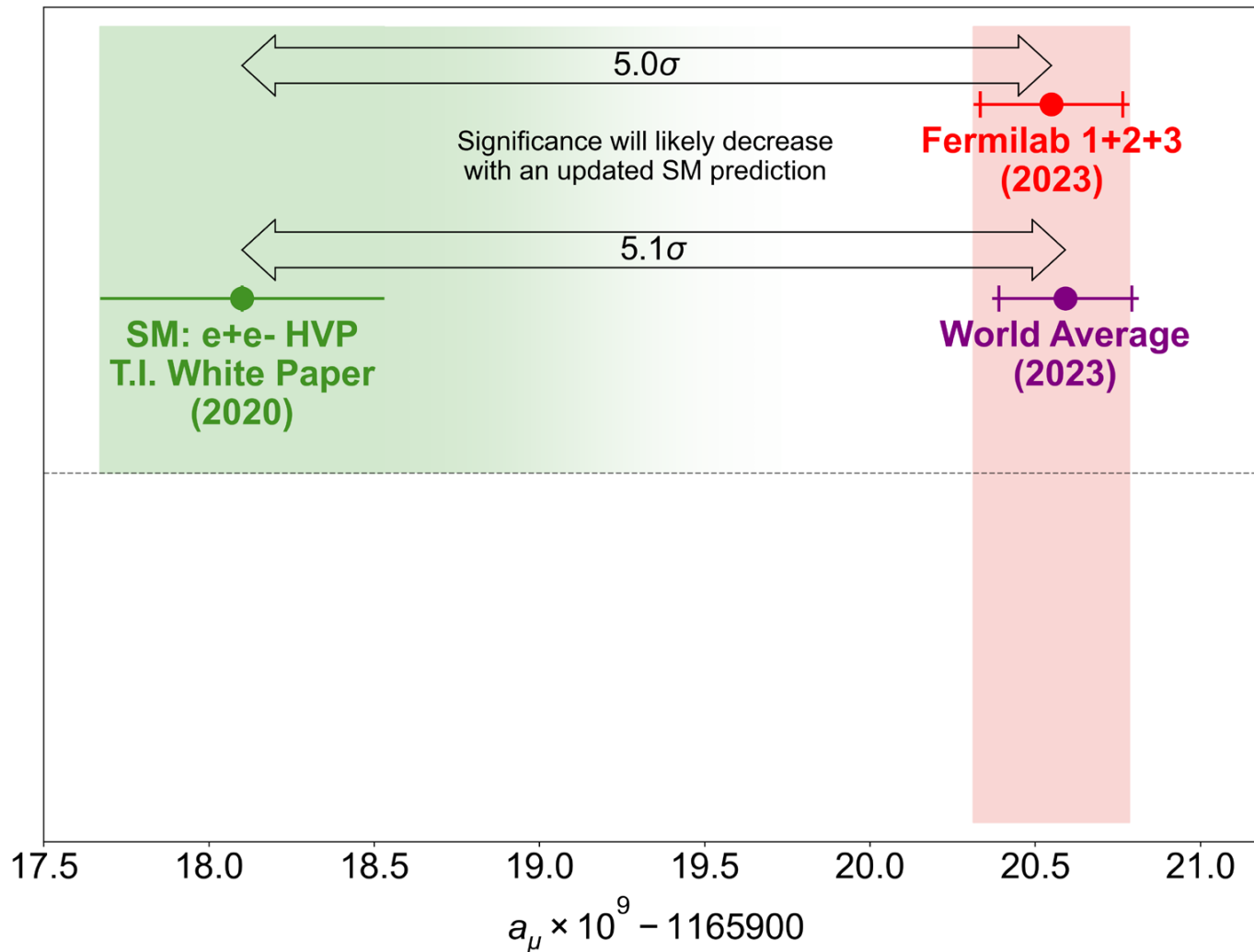
See talk tomorrow
by A. Driutti

$$a_\mu(\text{Exp}) = 0.00\ 116\ 592\ 059(22)$$
 [190 ppb]

2.1 Anomalous magnetic moment of the muon

FNAL g-2
James Mott'23

- In 2023



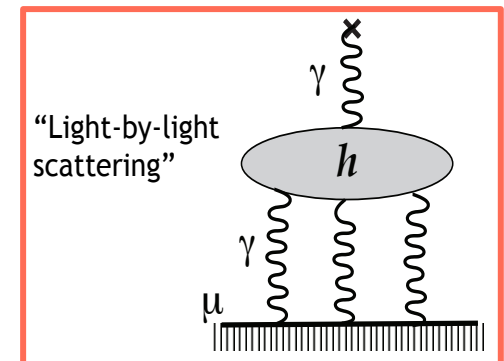
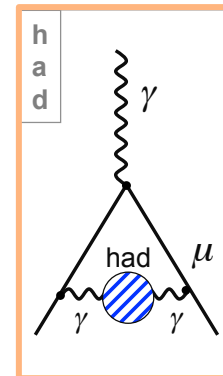
2.2 Confronting measurement and prediction

- Theoretical Prediction:

*Colangelo et al.
Snowmass 2022*

Contribution	Result in 10^{-10} units
QED(leptons)	11658471.885 ± 0.004
HVP(leading order)	693.1 ± 4.0
HVP(higher order) (*)	-8.59 ± 0.07
HLBL	9.2 ± 1.8
EW	15.4 ± 0.1
Total	11659181.0 ± 4.3

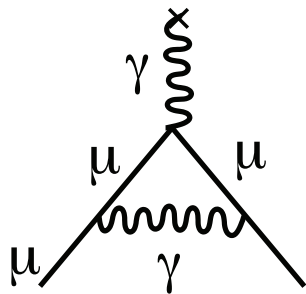
- Important contribution comes from virtual hadrons in the loop!
- Tackled using :
 - Models
 - Dispersion Relations
 - Lattice QCD



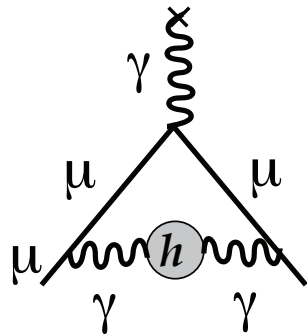
2.3 Contribution to $(g-2)_\mu$

Hoecker'11

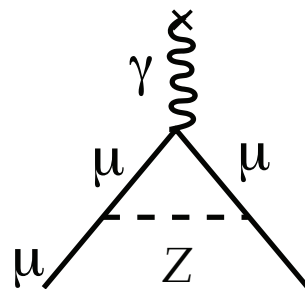
QED



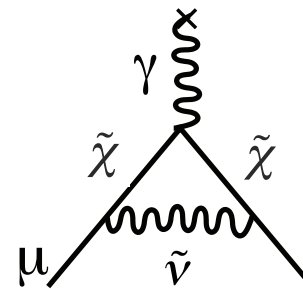
Hadronic



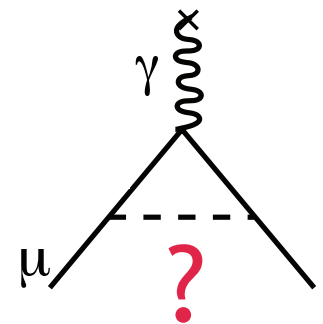
Weak



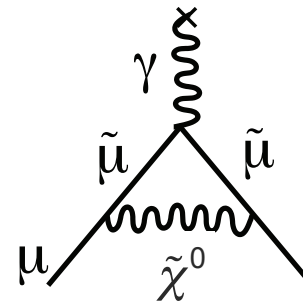
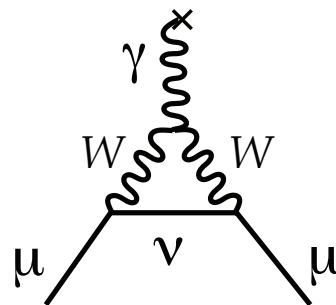
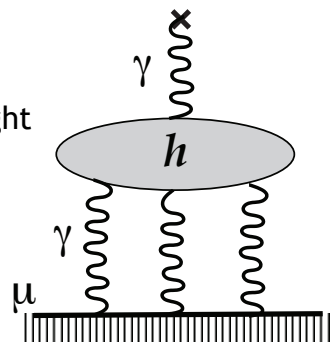
SUSY... ?



... or some unknown type of new physics ?



“Light-by-light scattering”



... or no effect on a_μ , but new physics at the LHC? That would be interesting as well !!

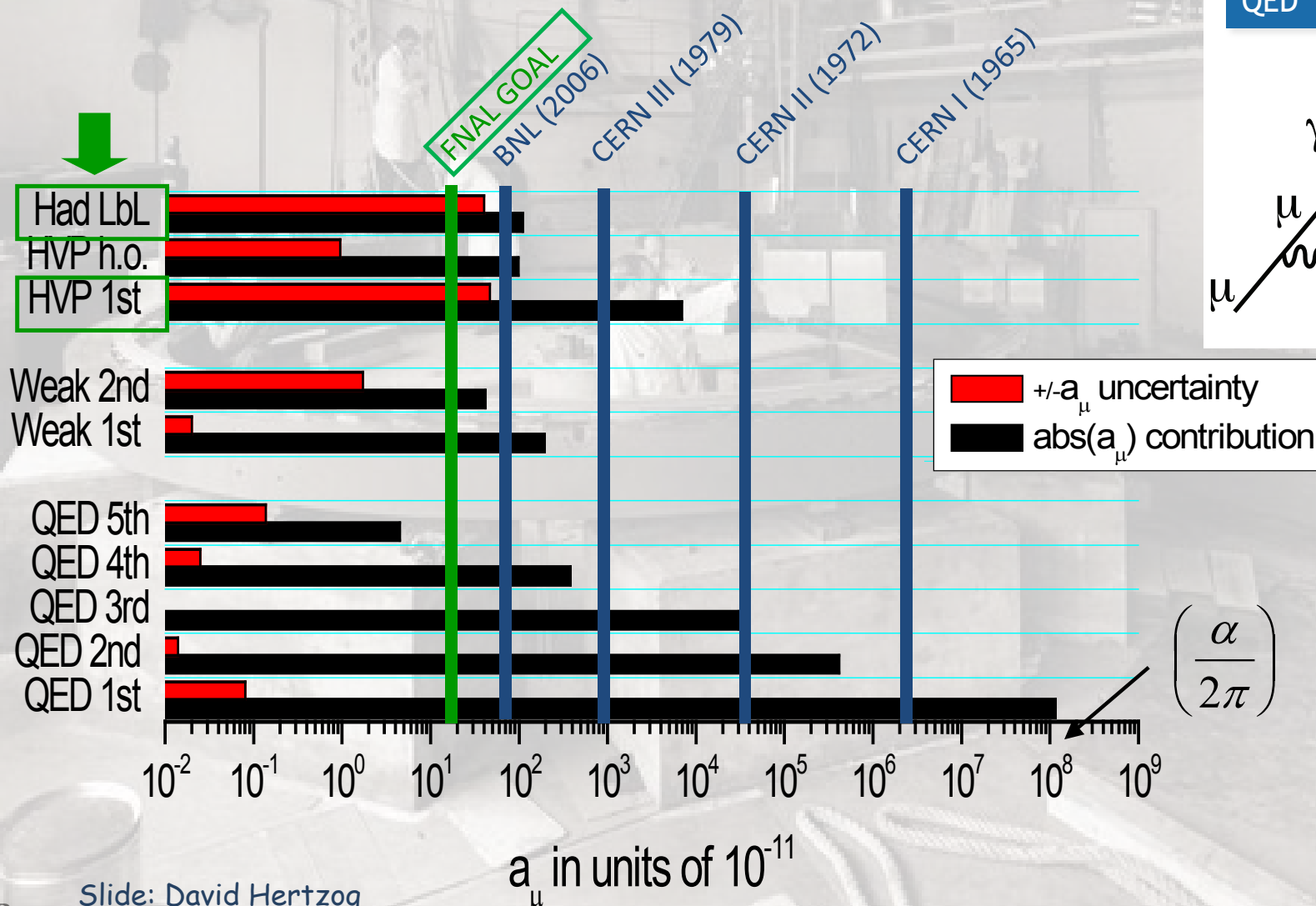
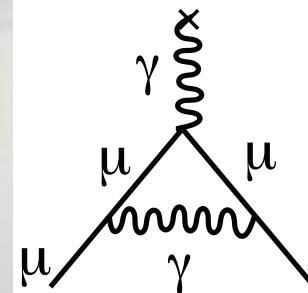
Need to compute the SM prediction with high precision!  *Not so easy!*

2.4 On the importance of hadronic contributions

From D. Hertzog

Muon $g-2$ measurements sensitivity

QED

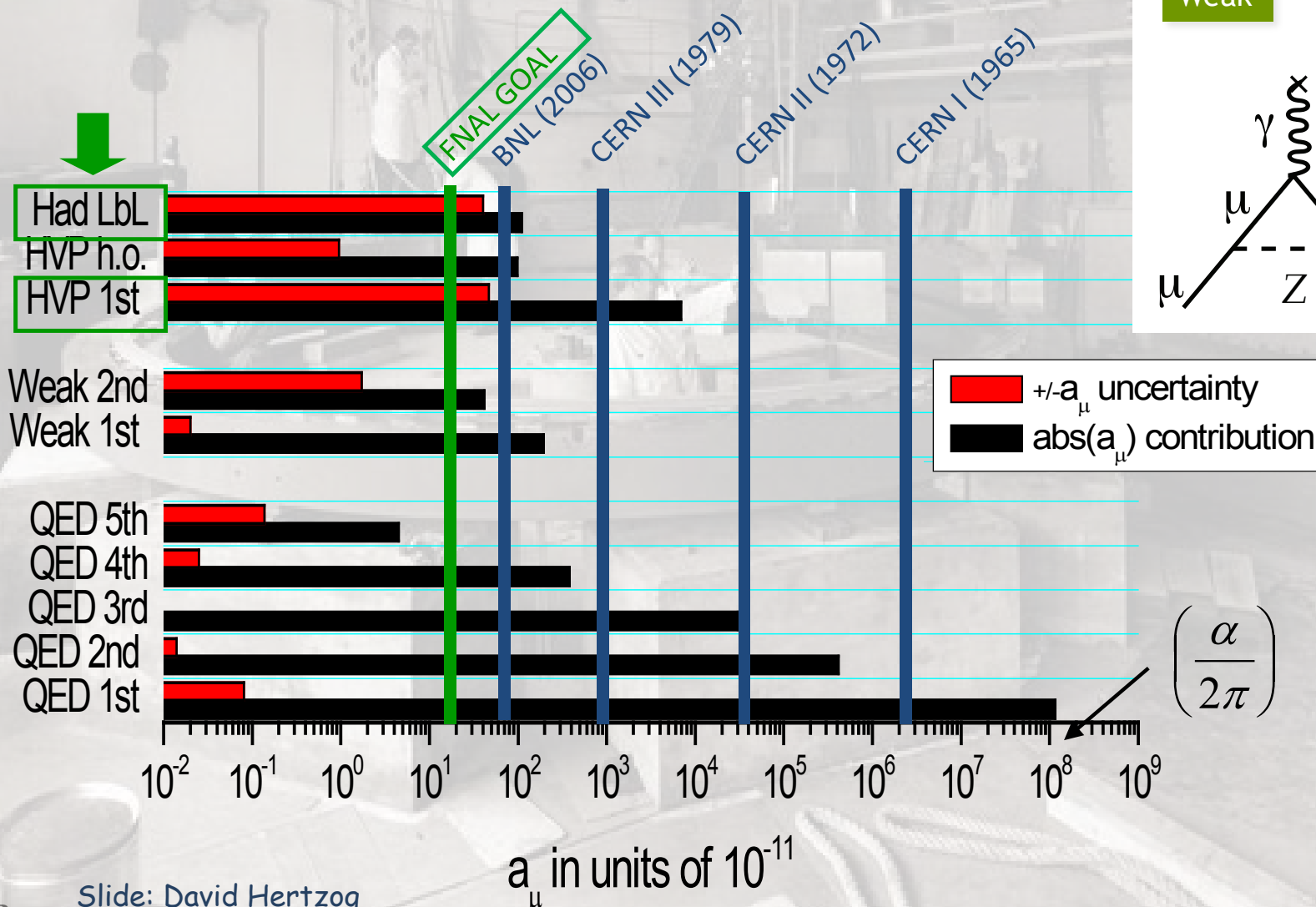


Slide: David Hertzog

2.4 On the importance of hadronic contributions

From D. Hertzog

Muon $g-2$ measurements sensitivity



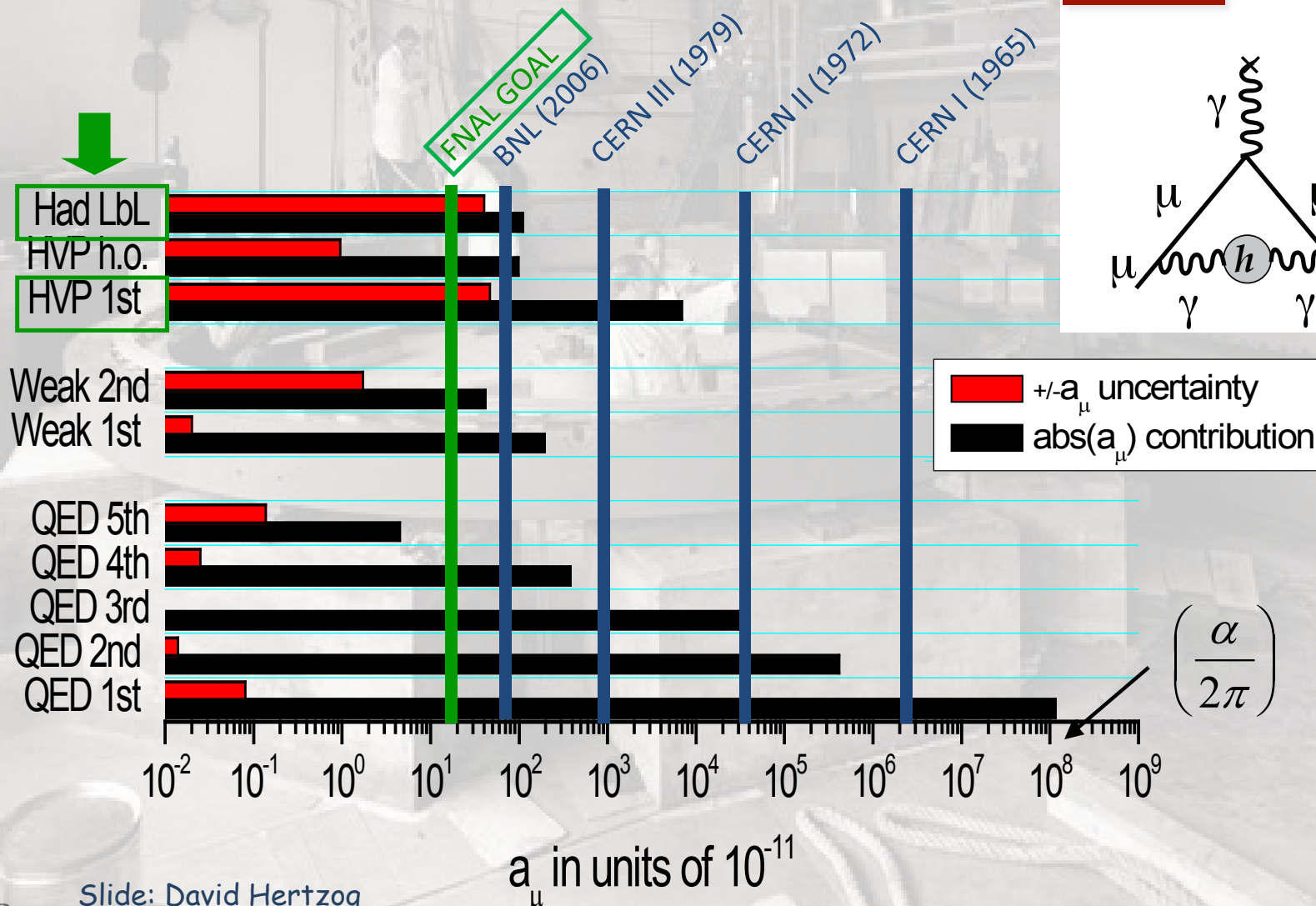
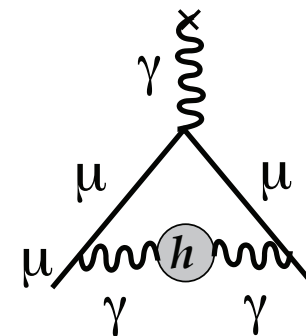
Slide: David Hertzog

2.4 On the importance of hadronic contributions

From D. Hertzog

Muon $g-2$ measurements sensitivity

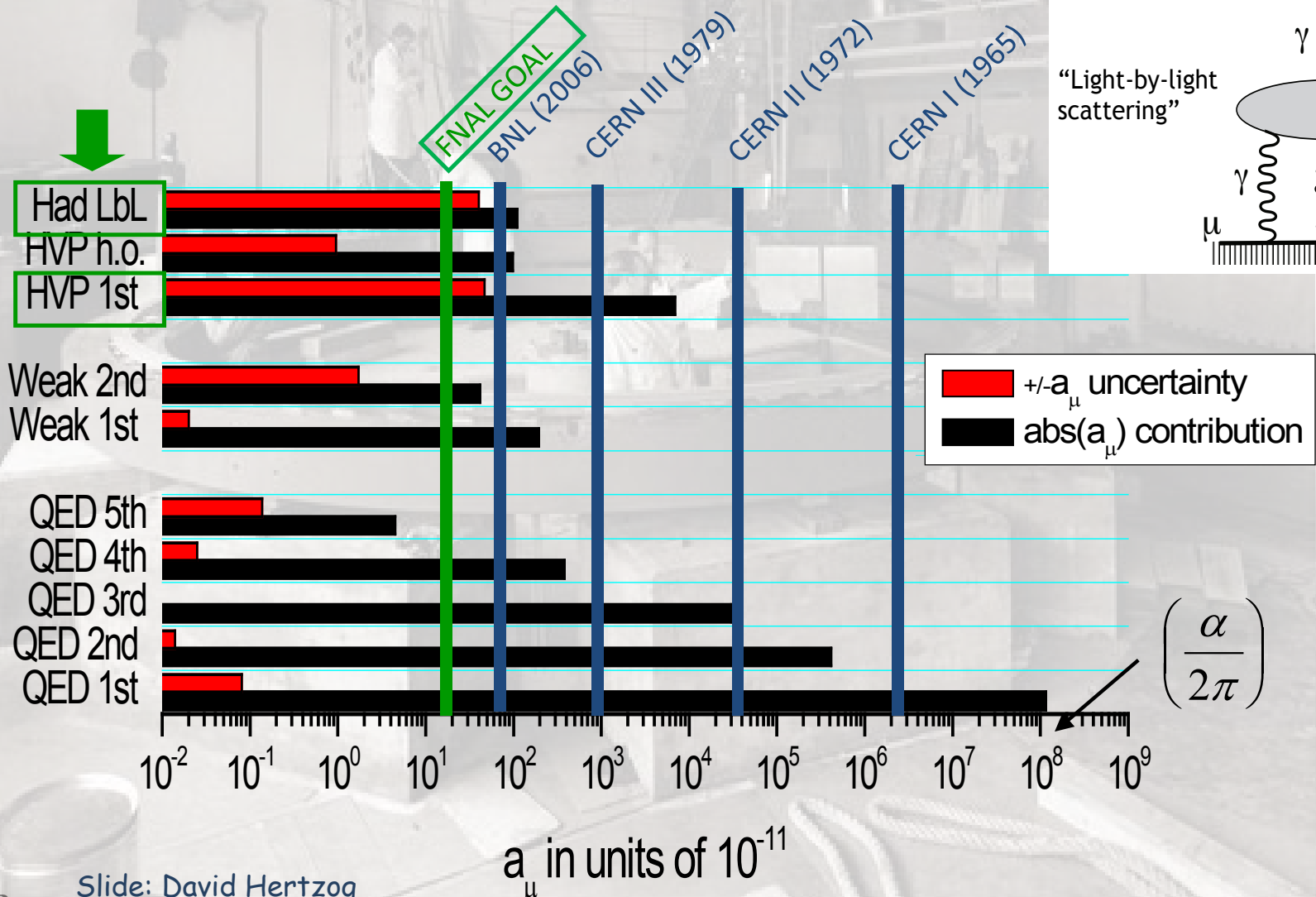
Hadronic



Slide: David Hertzog

2.4 On the importance of hadronic contributions

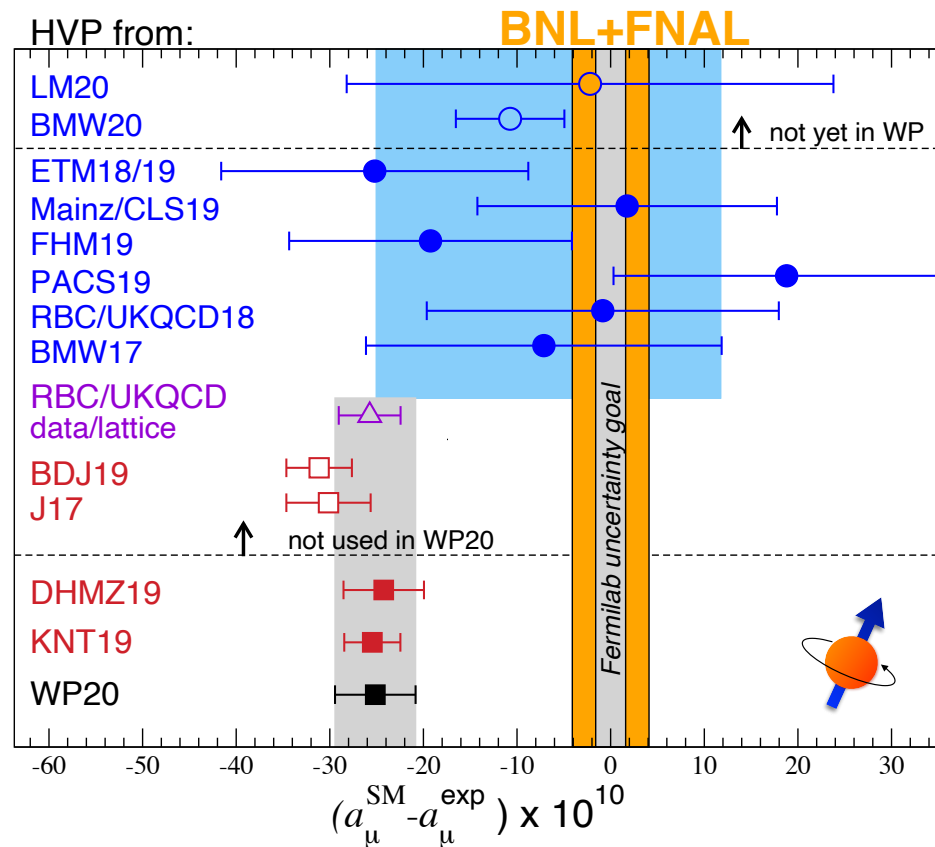
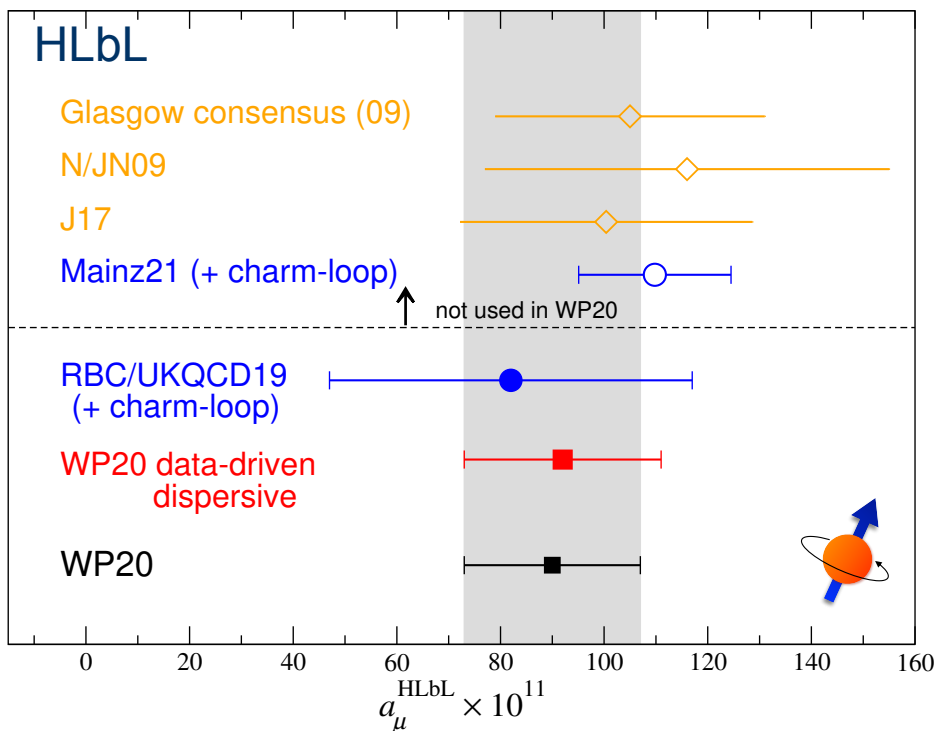
Muon $g-2$ measurements sensitivity



Slide: David Hertzog

2.5 Recent Developments

Colangelo et al.
Snowmass 2022



Comparison of the Standard Model Prediction using Lattice QCD, the data driven dispersive approaches and Models

➡ Tension on HVP between the lattice result from BMW and the data driven app.

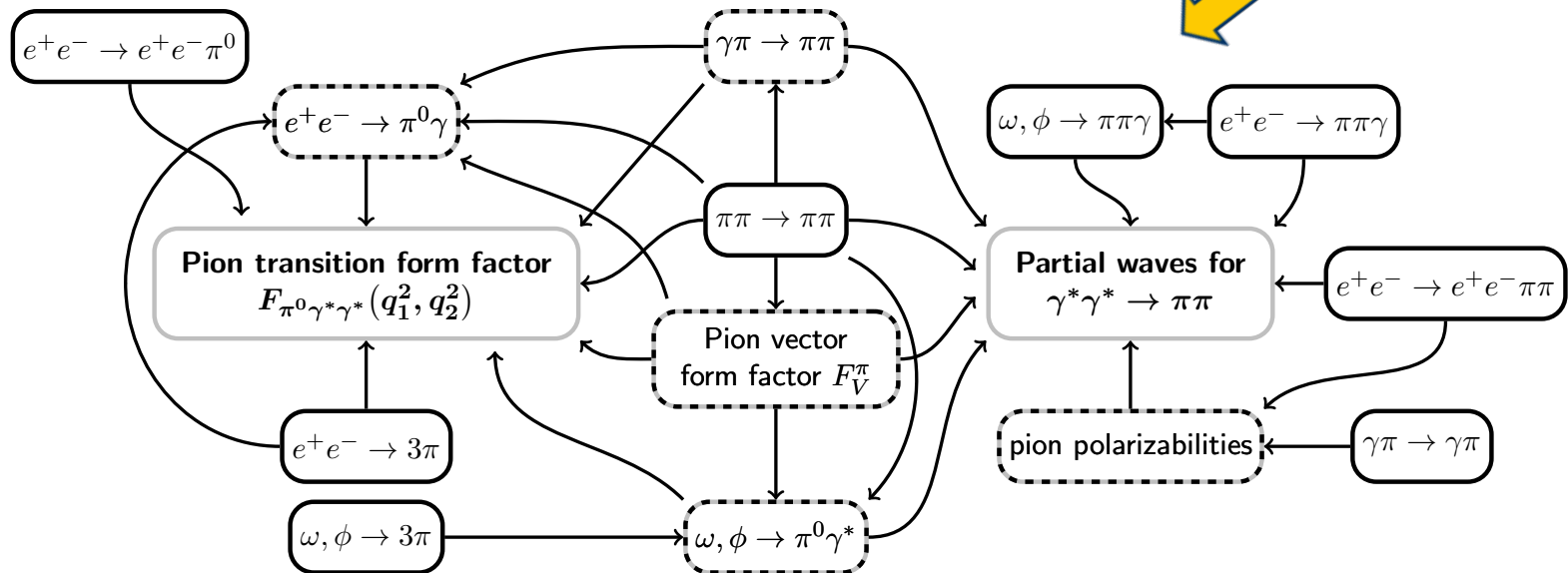
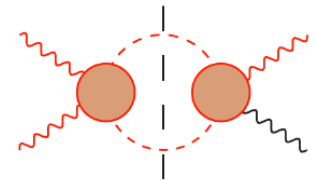
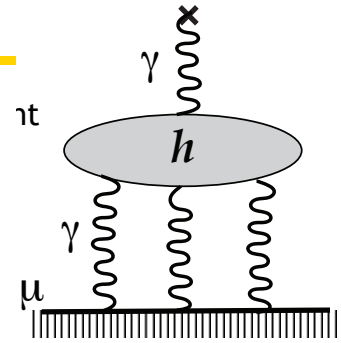
Model independent determination of LBL: experimental inputs

- For light-by-light scattering: until recently it was believed that dispersion relation approach not possible (4-point function)

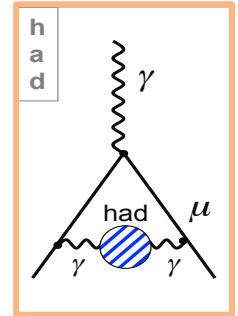
➔ only model dependent estimates

- But recent progress from Bern group: *Colangelo, Hoferichter, Procura, Stoffer'14*

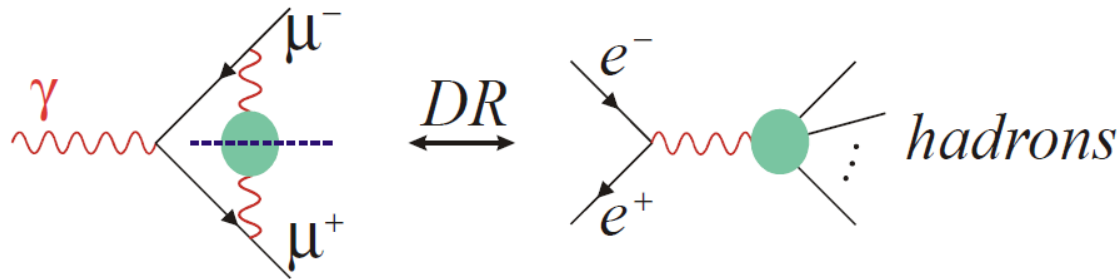
➔ *Data driven* estimate possible using *dispersion relations!*



Model independent determination of HVP



- Hadronic contribution cannot be computed from first principles due to low-energy hadronic effects
- Use analyticity + unitarity \Rightarrow real part of photon polarisation function from *dispersion relation* over *total hadronic cross section data*



$$R_V(s) = \frac{\sigma(e^+e^- \rightarrow \text{hadrons})}{\sigma(e^+e^- \rightarrow \mu^+\mu^-)}$$

- Leading order hadronic vacuum polarization :

$$a_\mu^{\text{had,LO}} = \frac{\alpha^2 m_\mu^2}{(3\pi)^2} \int_{4m_\pi^2}^{\infty} ds \frac{K(s)}{s^2} R_V(s)$$

- Low energy contribution dominates : $\sim 75\%$ comes from $s < (1 \text{ GeV})^2$
 \Rightarrow $\pi\pi$ contribution extracted from data

2.5 Recent Developments

*Ignatov et al., CMD-3,
2302.08834 [hep-ex]*

- New result from CMD3 in Novosibirsk

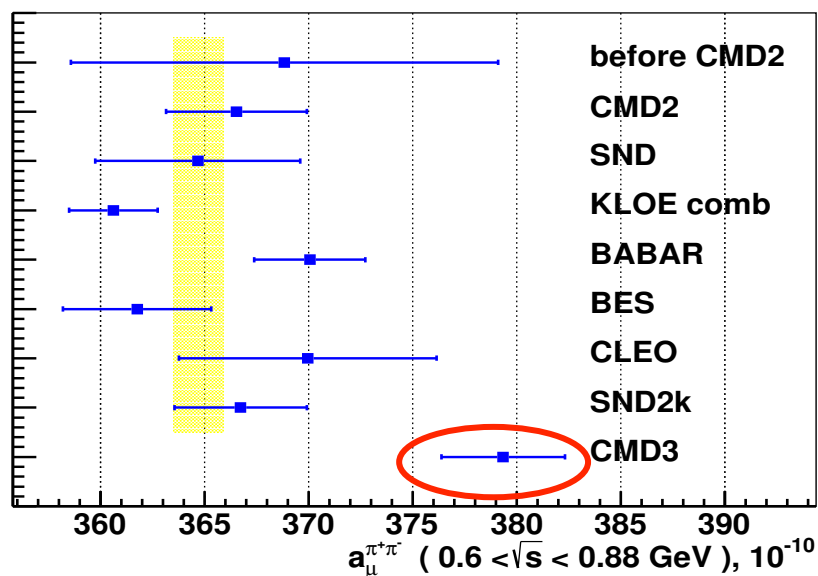


Figure 36: The $\pi^+\pi^-(\gamma)$ contribution to $a_\mu^{had,LO}$ from energy range $0.6 < \sqrt{s} < 0.88$ GeV obtained from this and other experiments.

Experiment	$a_\mu^{\pi^+\pi^-,LO}, 10^{-10}$
before CMD2	368.8 ± 10.3
CMD2	366.5 ± 3.4
SND	364.7 ± 4.9
KLOE	360.6 ± 2.1
BABAR	370.1 ± 2.7
BES	361.8 ± 3.6
CLEO	370.0 ± 6.2
SND2k	366.7 ± 3.2
CMD3	379.3 ± 3.0

Table 4: The $\pi^+\pi^-(\gamma)$ contribution to $a_\mu^{had,LO}$ from energy range $0.6 < \sqrt{s} < 0.88$ GeV obtained from this and other experiments.

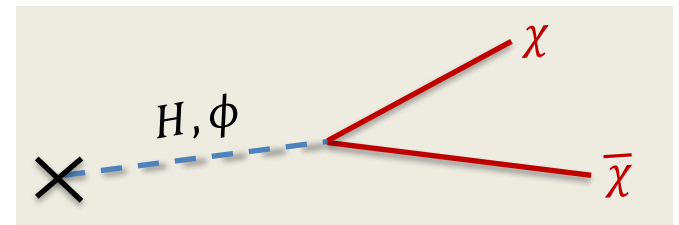
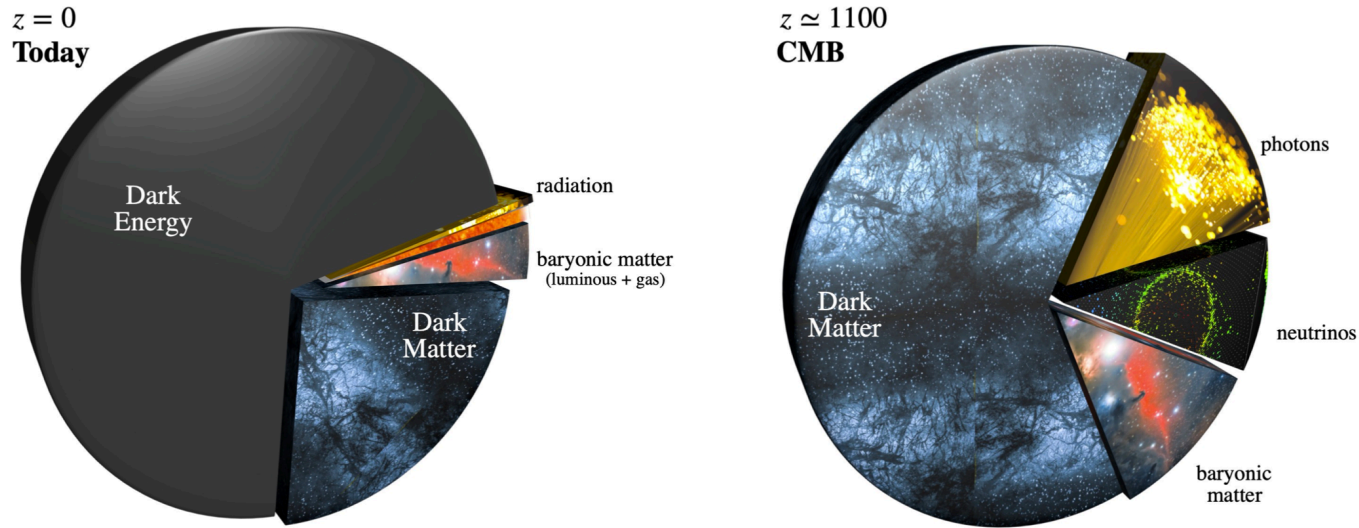
3. Looking for dark matter with a Higgs-mixed Scalar

In collaboration with *P. Blackstone (Indiana University),*
J. Tarrus Castella (Barcelona)
J. Zupan (Cincinnati)

to appear

3.1 Scalar Portal to Dark Matter

- Evidence of Dark Matter from gravitational interactions

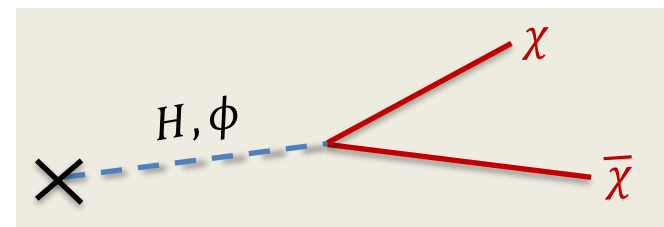


3.1 Scalar Portal to Dark Matter

- Evidence of Dark Matter from gravitational interactions
- It could couple to visible matter
- Different portals:

Portal	Interactions
Dark Photon, A'_μ	$-\epsilon F'_{\mu\nu} B^{\mu\nu}$
Dark Higgs, S	$(\mu S + \lambda S^2) H^\dagger H$
Heavy Neutral Lepton, N	$y_N L H N$
Axion-like pseudo scalar, a	$a F \tilde{F} / f_a, a G \tilde{G} / f_a, (\bar{\psi} \gamma^\mu \gamma_5 \psi) \partial_\mu a / f_a$

See e.g. Cirelli, Strumia & Zupan'24



3.2 Higgs-mixed scalar portal

$$\mathcal{L}_{\text{scalar}} = \mathcal{L}_{\text{SM}} + \mathcal{L}_{\text{DS}} - (\mu S + \lambda_{SH} S^2) H^\dagger H$$



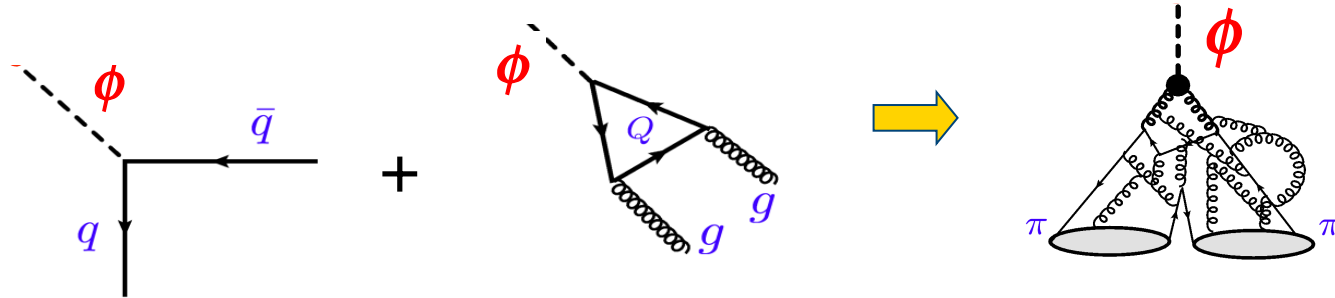
$$\mathcal{L}_{\text{eff}} = - \sum_q c_q \frac{m_q}{v_W} \bar{q} q \phi - \sum_\ell c_\ell \frac{m_\ell}{v_W} \bar{\ell} \ell \phi + c_g \frac{\alpha_s}{12\pi v_W} \phi G_{\mu\nu}^a G^{a\mu\nu} + c_\gamma \frac{\alpha}{\pi v_W} \phi F_{\mu\nu} F^{\mu\nu}$$

- After EW symmetry breaking S - H mixing \Rightarrow mass eigenstate which is predominantly S is called ϕ
- **Higgs-mixed scalar scenario:** $c_q = c_\ell = c_g = \sin \theta_h = s_\theta$

just two parameters: $\sin \theta_h$ and m_ϕ

3.3 Constraints below 2 GeV

- Decay dominated at low energies by 2 pions and 2 kaons



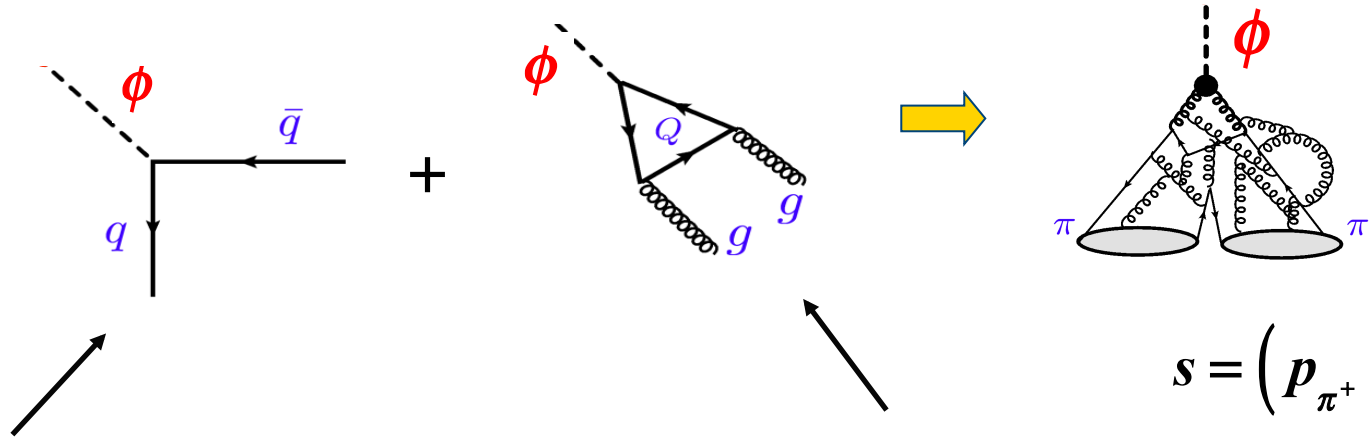
- Problem : Have the hadronic part under control, ChPT not valid at these energies!

➔ Use *form factors* determined with *dispersion relations* matched at low energy to *CHPT*
Donoghue, Gasser and Leutwyler'90, Truong & Wiley'89, Raby & West'88, Voloshin'86
Celis, Cirigliano, E.P.'14, Monin, Boyarsky & Ruchayskiy'18, Winkler'19

- Dispersion relations: based on **unitarity**, **analyticity** and **crossing symmetry**

➔ Take *all rescattering* effects into account
 $\pi\pi$ final state interactions important

3.4 ϕ decays



$$\langle \pi^+ \pi^- | m_u \bar{u}u + m_d \bar{d}d | 0 \rangle \equiv \Gamma_\pi(s)$$

$$\langle \pi^+ \pi^- | \theta_\mu^\mu | 0 \rangle \equiv \theta_\pi(s)$$

$$\langle \pi^+ \pi^- | m_s \bar{s}s | 0 \rangle \equiv \Delta_\pi(s)$$

$$\theta_\mu^\mu = -9 \frac{\alpha_s}{8\pi} G_{\mu\nu}^a G_a^{\mu\nu} + \sum_{q=u,d,s} m_q \bar{q}q$$

$$\Gamma_{\phi \rightarrow PP} = \frac{A_P}{16\pi m_\phi} \sigma_P(m_\phi^2) |G_P(m_\phi^2)|^2$$

$A_P = \#$ final states

$$G_P(s) = \left\langle PP \left| c_g \frac{\alpha_s}{12\pi v_W} G_{\mu\nu}^a G^{a\mu\nu} - \sum_q c_q \frac{m_q}{v_W} \bar{q}q \right| 0 \right\rangle$$

3.5 Unitarity

- Coupled channel analysis** up to $\sqrt{s} \sim 2$ GeV: *Mushkelishvili-Omnès* approach

Inputs: $l=0$, S-wave $\pi\pi$ and KK data

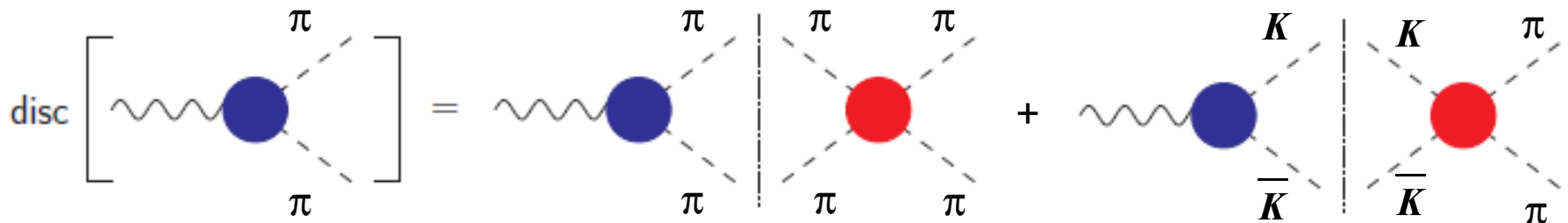
Donoghue, Gasser, Leutwyler'90

Moussallam'99

Daub, Dreiner, Hanart, Kubis, Meissner'13

Celis, Cirigliano, E.P.'14

- Unitarity \Rightarrow the discontinuity of the form factor is known



$$\text{Im}F_n(s) = \sum_{m=1}^2 T_{nm}^*(s)\sigma_m(s)F_m(s)$$

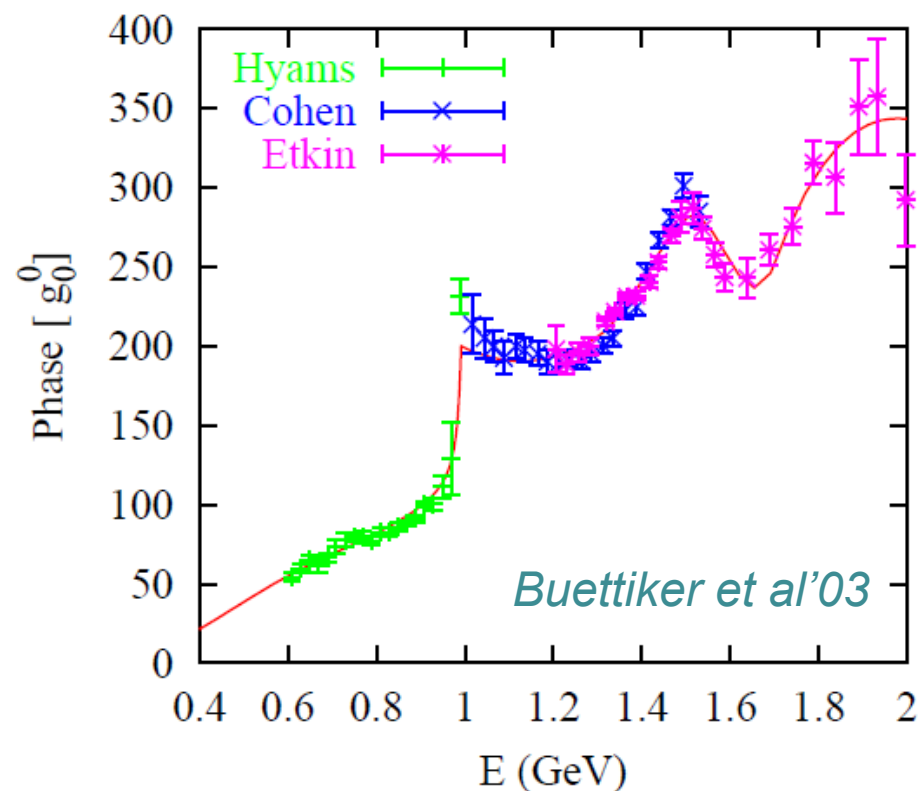
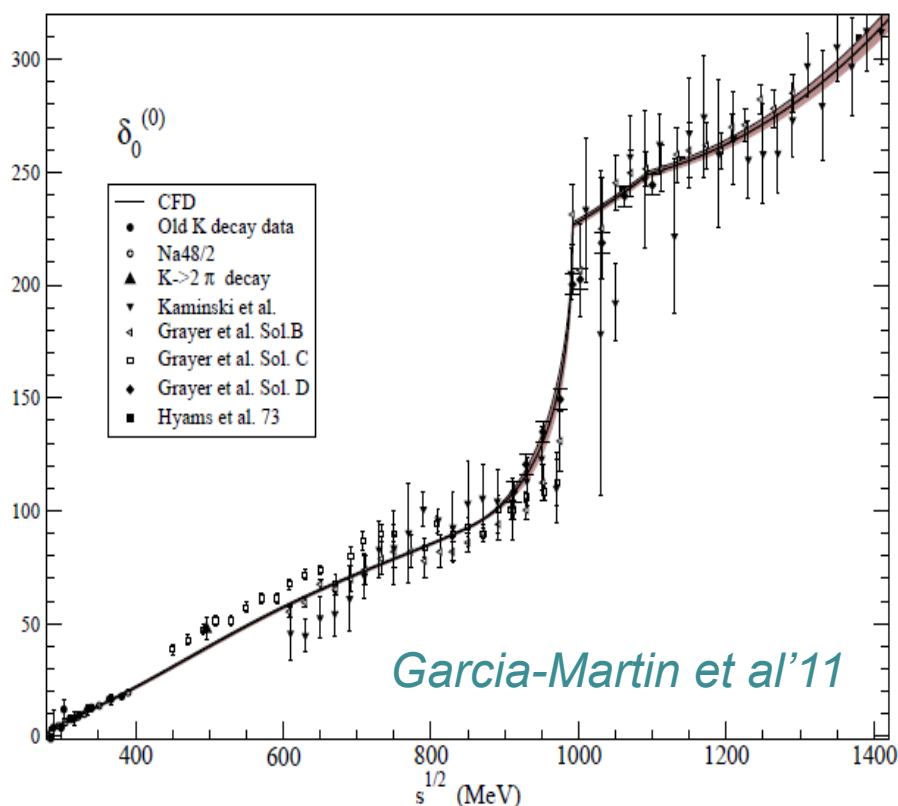
$n = \pi\pi, K\bar{K}$

Scattering matrix:

$$\begin{pmatrix} \pi\pi \rightarrow \pi\pi, & \pi\pi \rightarrow K\bar{K} \\ K\bar{K} \rightarrow \pi\pi, & K\bar{K} \rightarrow K\bar{K} \end{pmatrix}$$

3.6 Inputs for the coupled channel analysis

- Inputs : $\pi\pi \rightarrow \pi\pi, K\bar{K}$



- A large number of theoretical analyses *Descotes-Genon et al'01*, *Kaminsky et al'01*, *Buettiker et al'03*, *Garcia-Martin et al'11*, *Colangelo et al.'11*, *Pelaez & Rodas'22* and all agree
- 3 inputs: $\delta_\pi(s)$, $\delta_K(s)$, η from *Pelaez et al.* \rightarrow reconstruct *T matrix*

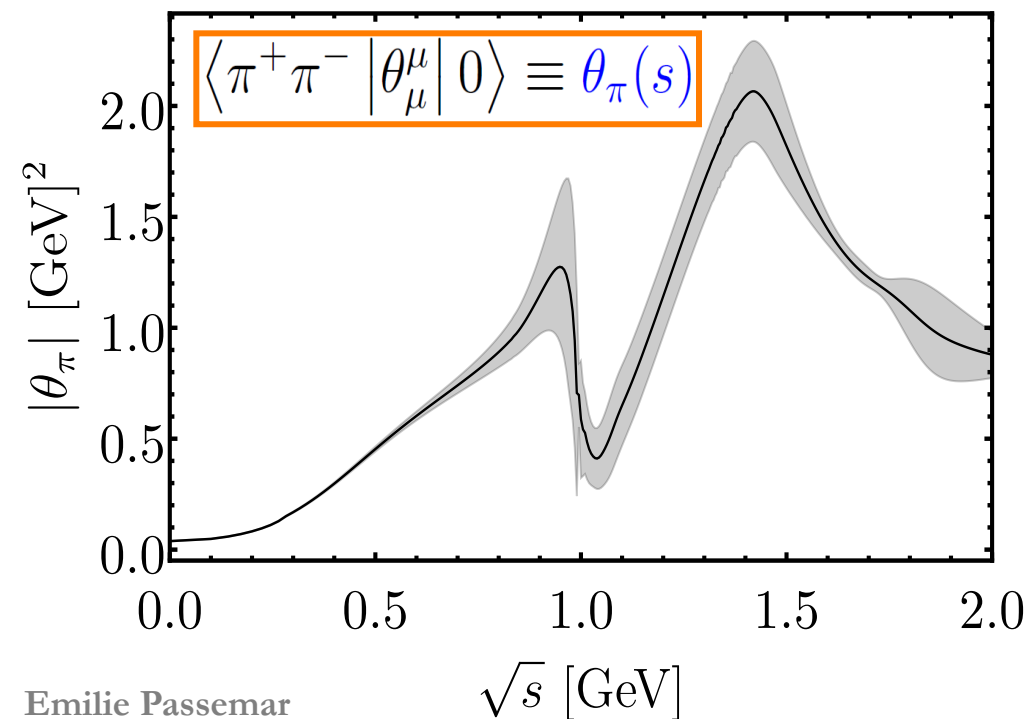
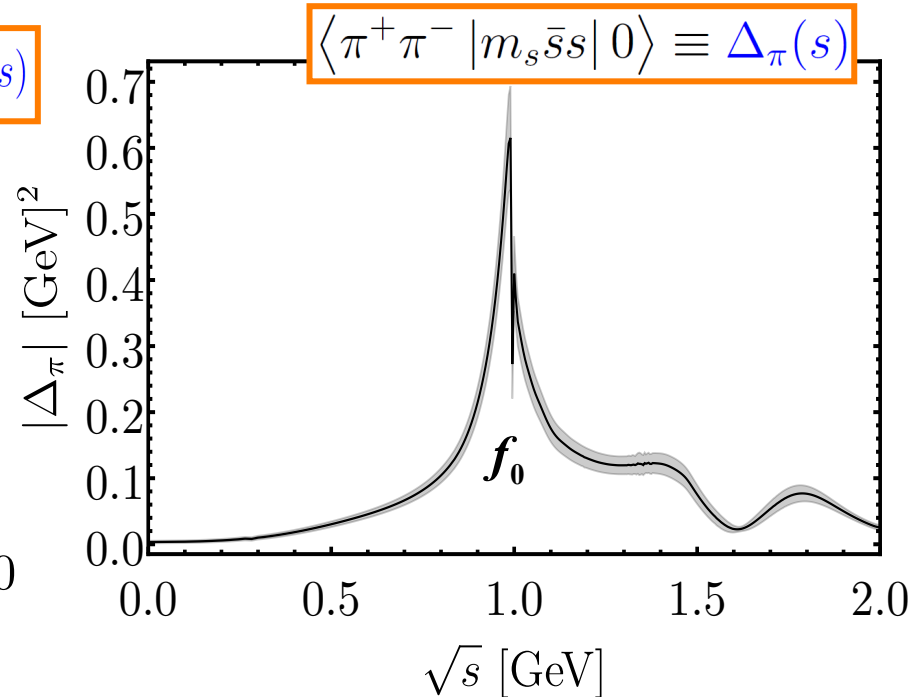
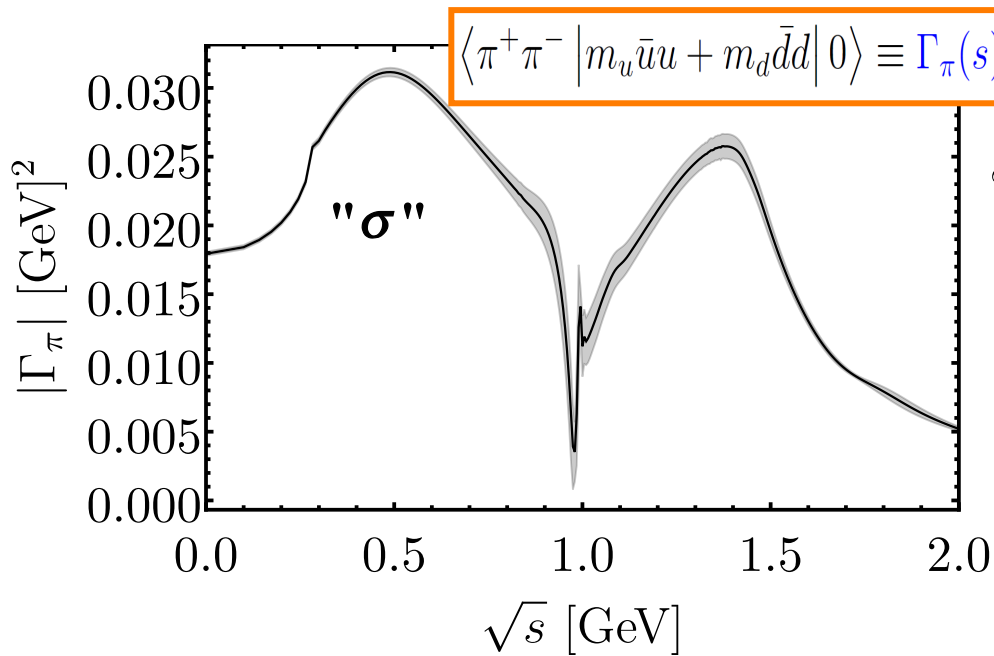
3.7 Dispersion relations

- General solution to *Mushkhelishvili-Omnès* problem:

$$\begin{pmatrix} F_\pi(s) \\ \frac{2}{\sqrt{3}}F_K(s) \end{pmatrix} = \begin{pmatrix} C_1(s) & D_1(s) \\ C_2(s) & D_2(s) \end{pmatrix} \begin{pmatrix} P_F(s) \\ Q_F(s) \end{pmatrix}$$

Canonical solution falling as $1/s$ for large s (obey unsubtracted dispersion relations)

Polynomial determined from a matching to ChPT + lattice

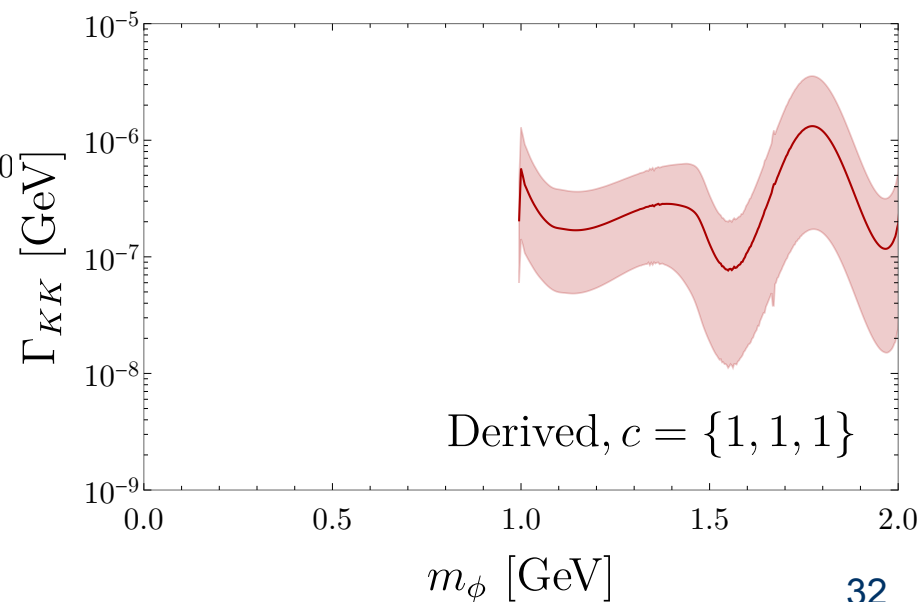
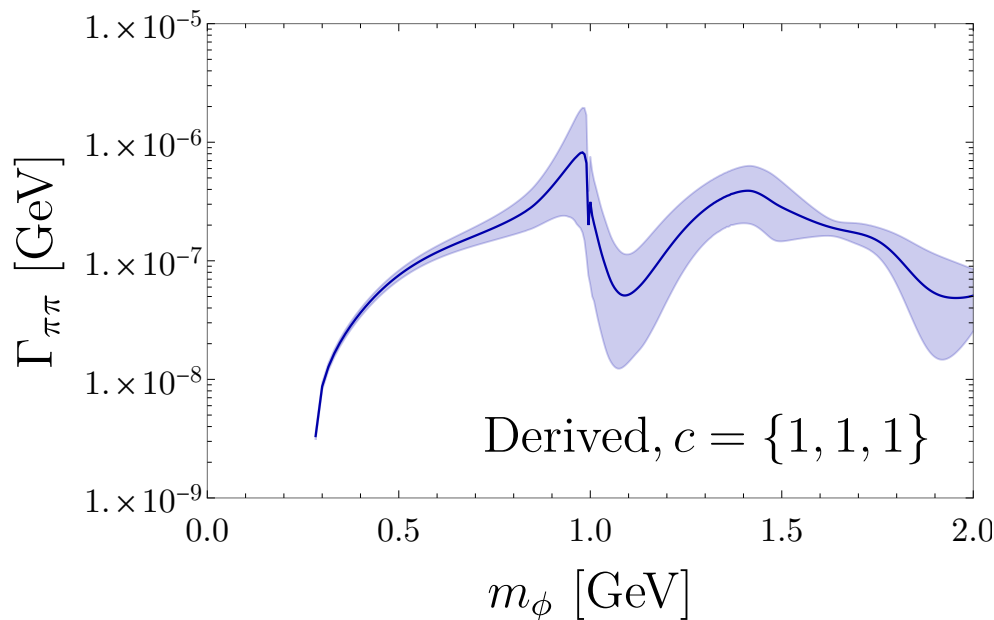


- Uncertainties:
 - Varying the matching conditions
 - T matrix inputs

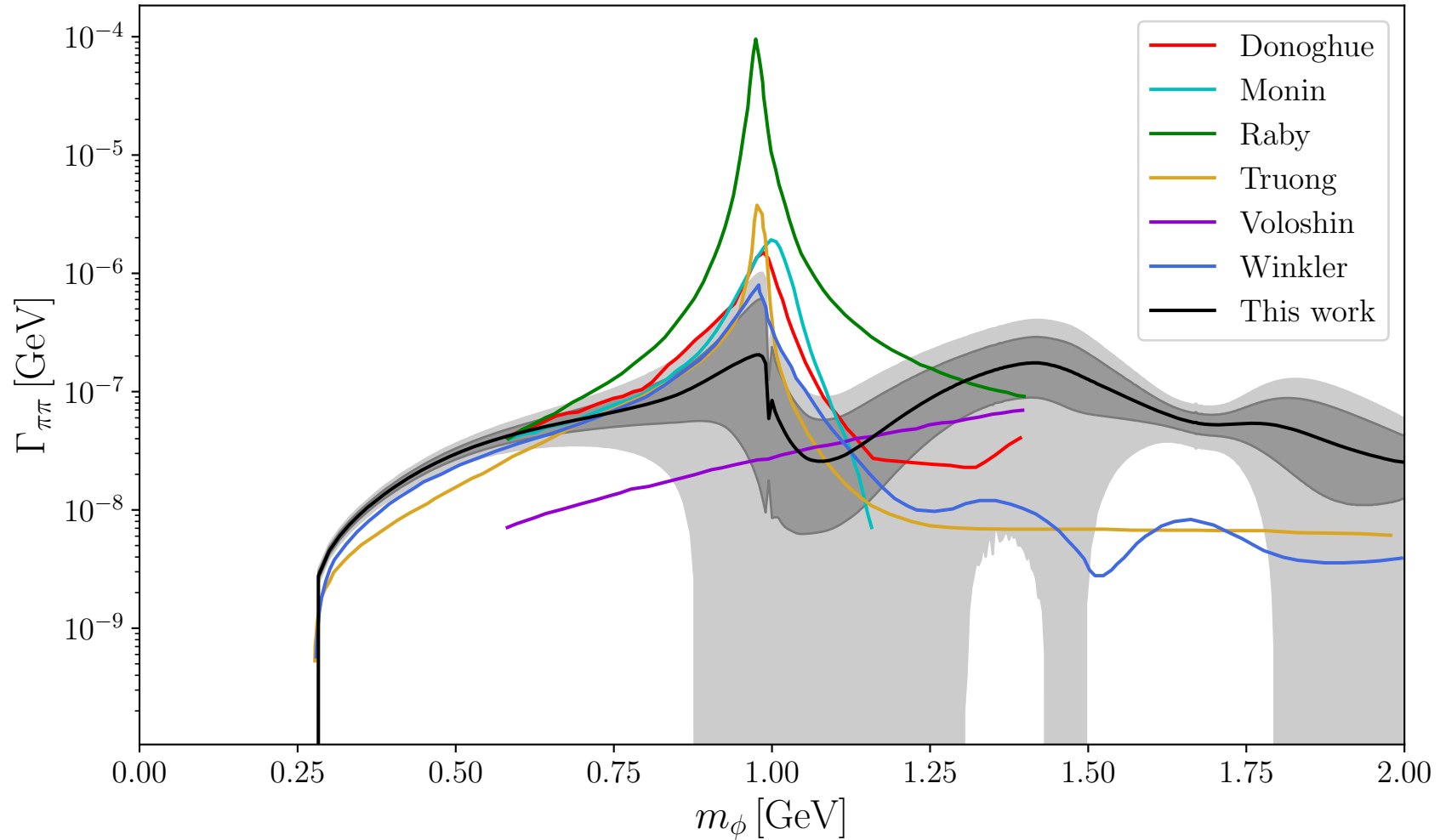
3.8 Decay widths

$$\Gamma_{\phi \rightarrow PP} = \frac{A_P}{16\pi m_\phi} \sigma_P(m_\phi^2) |G_P(s)|^2$$

$$A_\pi = 3$$
$$A_K = 4$$



3.8 Decay widths: comparison



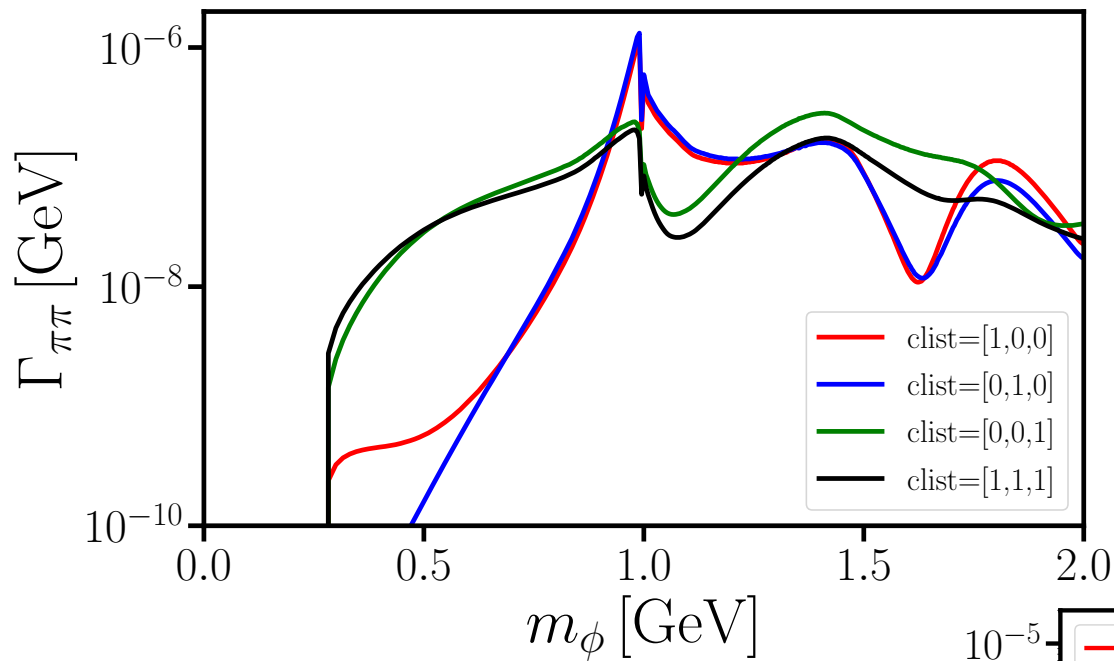
3.9 General Coupling Structure: hipsofcobra

github.com/blackstonep/hipsofcobra

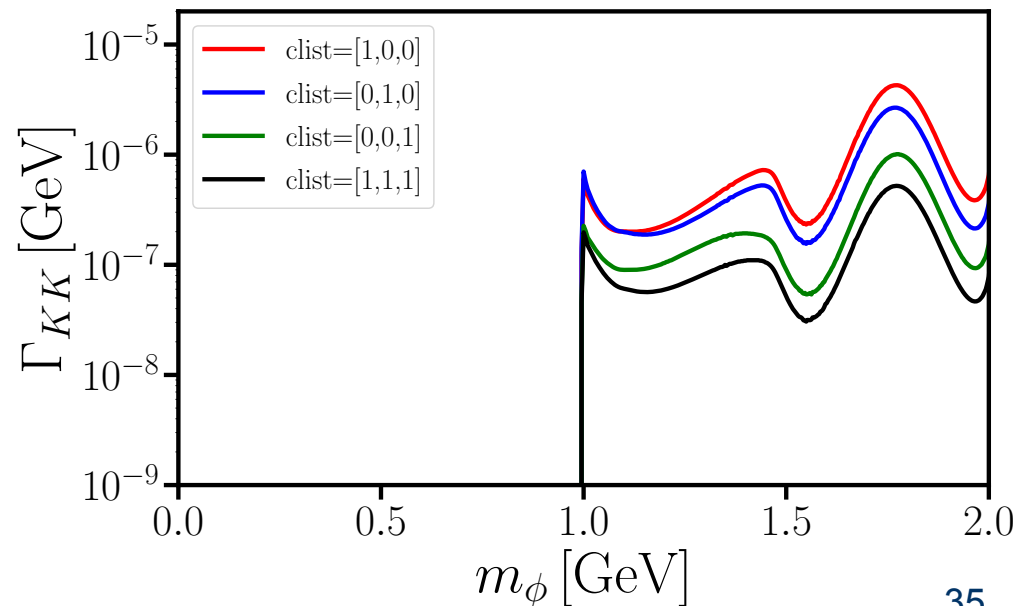
Higgs
Portal
Scalar
Off-
Flavor
Couplings
Branching
Ratio

```
1 from hipsofcobra import HipsofCobra
2 import numpy as np
3 import matplotlib.pyplot as plt
4
5 hips = HipsofCobra( clist=[1,1,1], Pname='pi', method='derived' )
6
7 hips.write_widths() # Write widths csv file.
8
9 hips.plot_G_contours(
10     color='k', xlim=[0,4], ylim=None, PrintQ=True, ShowQ=True
11 ) # Produce pdf plot of |G|.
12
13 hips.plot_width_contours(
14     color='k', xlim=[0,2], ylim=[1e-9,1e-5], PrintQ=True, ShowQ=True
15 ) # Produce pdf plot of width phi->PP.
16
17 hips.plot_sl(
18     xlim=[0,4], ylim=None, PrintQ=True, ShowQ=True
19 ) # Produce plot of |G|s from all iterations.
```

3.9 General Coupling Structure: hipsofcobra




coupl. to u,d
coupl. to s
coupl. to g
coupl. to u,d,s,g



4. Conclusion and Outlook

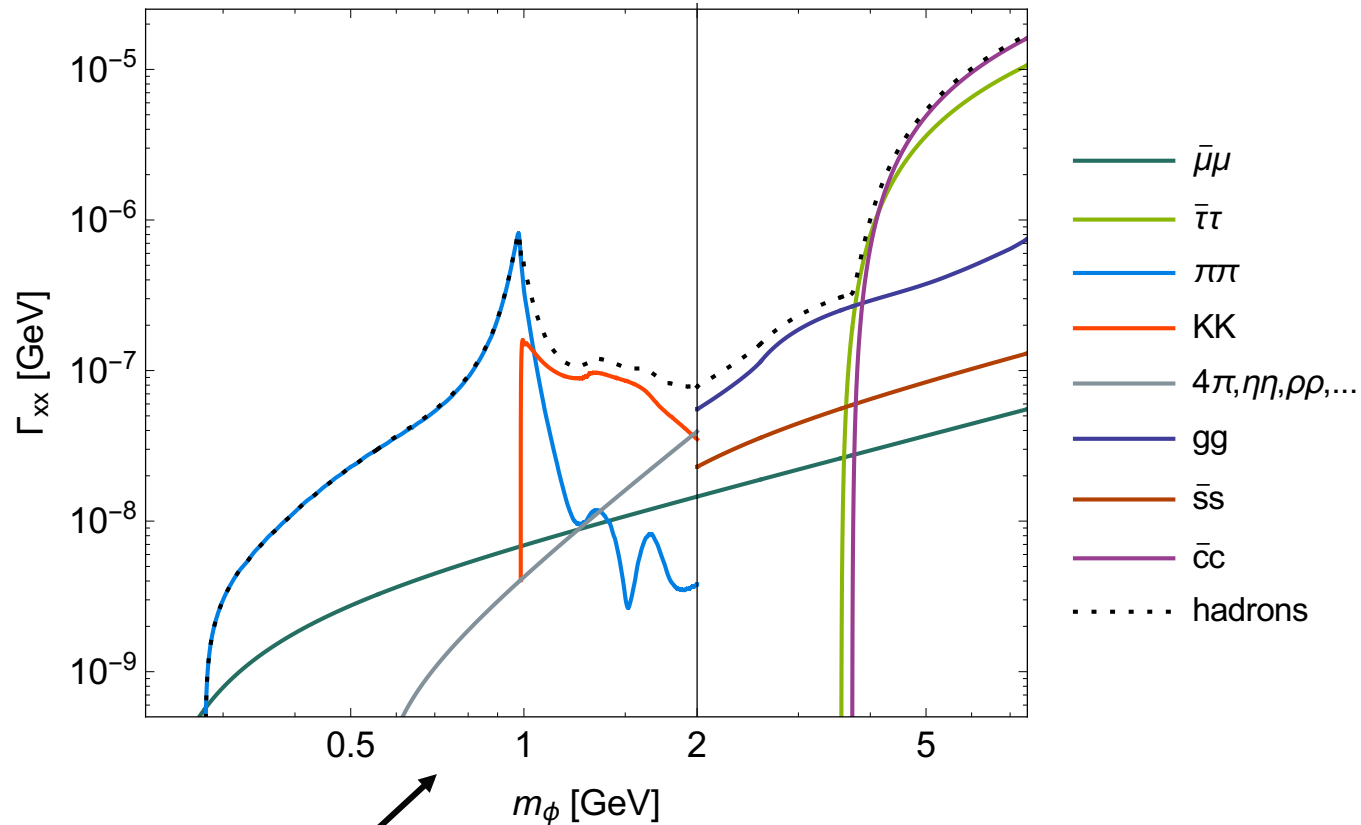
Conclusion and Outlook

- To look for physics beyond the Standard Model rare processes and precision measurements are very useful and powerful
- Many experiments are giving very precise results in the flavour sector, e.g, *Belle-II, BESIII, LHCb, NA62, JLab*
- Matching theoretically the level of precision is crucial  For some modes hadronic uncertainties are the limiting factor
- Theoretical tools: EFTs such as ChPT, Dispersion Relations, Lattice QCD...
- I gave 2 examples:
 - Anomalous magnetic moment of the muon
 - Constraints on Higgs-mixed Scalar DM scenarios
- There have been many more this afternoon and through the whole

5. Back-up

3.4 Prospects

From Winkler'19



Improvement up to 2 GeV

- Can we also improve at higher energies?

3.4 Prospects

From Winkler'19

