



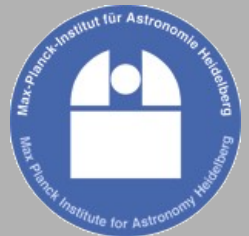
ZENTRUM FÜR
ASTRONOMIE

DFG Deutsche
Forschungsgemeinschaft
German Research Foundation

Beyond Gaia DR3: tracing the $[\alpha/M] - [M/H]$ bimodality from the Inner to the outer Milky Way disc with *Gaia* RVS and Convolutional Neural-Networks

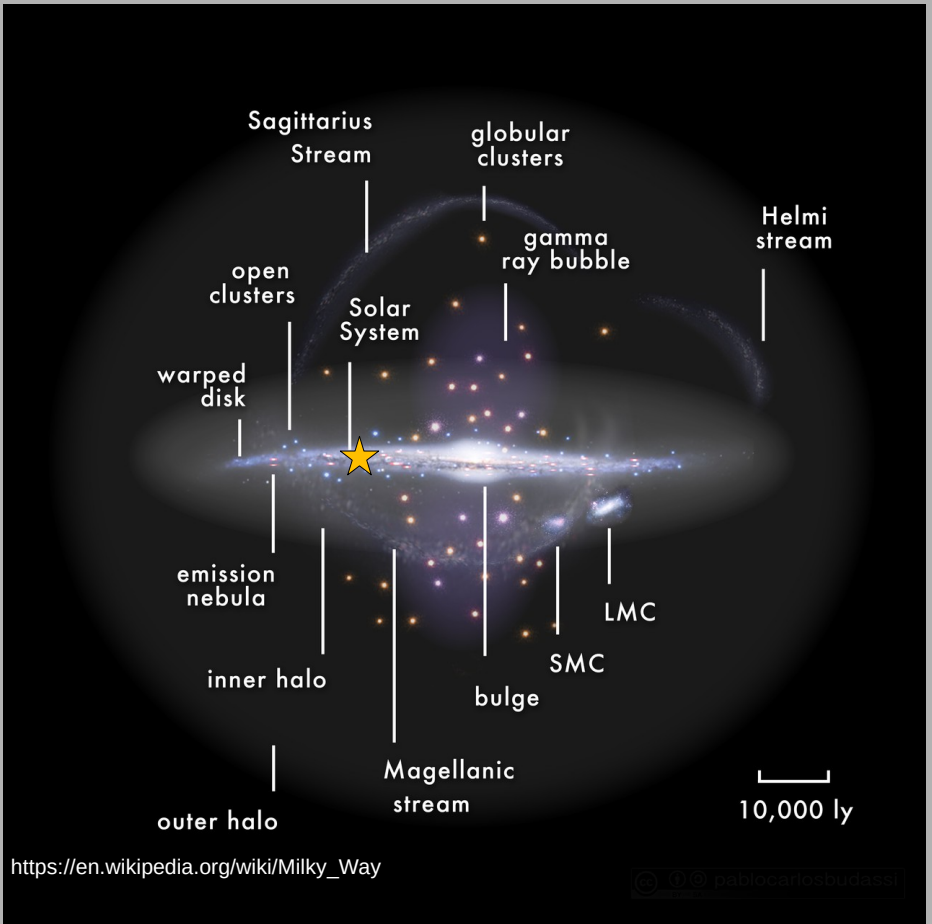
Guillaume Guiglion (GG; LSW / MPIA, guiglion@mpia.de)
& Samir Nepal (AIP)

The Milky Way Revealed by Gaia:
The Next Frontier (Sept. 5-7, 2023)



Galactic Archaeology

- studying the formation and evolution of the Milky Way and it's local volume
- **need for stellar chemistry, kinematics & ages**

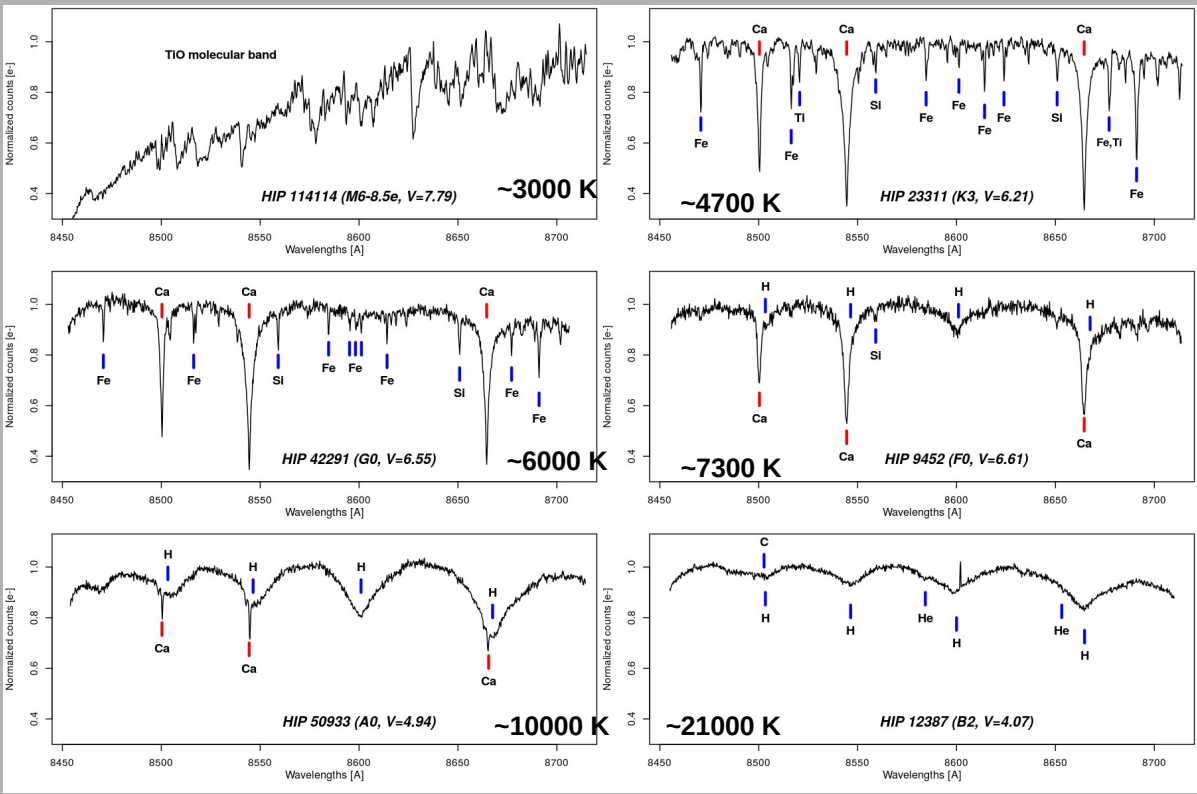


A collage of logos and images for various astronomical surveys and observatories. At the top left is the RAVE (RAy AStrometric and VELOCITY) Survey logo. Next to it is the Gaia ESO logo, featuring a satellite dish. To the right is a photograph of the LAMOST (Large Area Multi-Object Spectroscopic Telescope Facility) observatory. Below these are the APOGEE (Asteroseismic and Planet Oscillations and Gyrochronology) logo and the GALAH (Galactic Archaeology LAMOST) logo. A large blue arrow points downwards from GALAH to the SDSS (Sloan Digital Sky Survey) logo, which includes a target icon. To the right of SDSS is the Gaia logo. Below SDSS is the MOONS (Multi-Object Near-Infrared Spectrograph) logo, featuring images of moons. To the left of MOONS is a large 3D number '4' with 'MOST' written below it. At the bottom right is the Maunakea Spectroscopic Explorer logo.

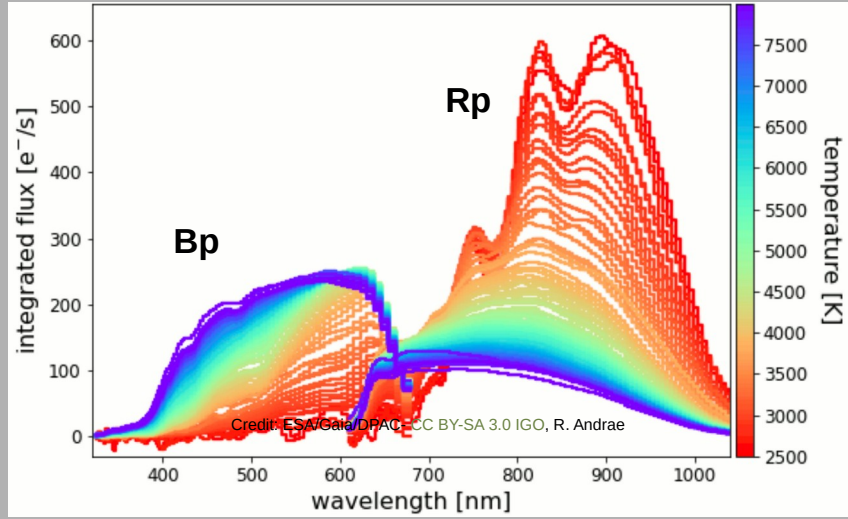
Huge data analysis challenge !!

What Gaia DR3 gave us:

→ 10^6 RVS spectra, $R \sim 11500$ (Katz et al. 2022)



→ 220 millions BP&RP spectra
 $R \sim 30-100$ (De Angeli et al. 2022)



- 1.5×10^6 parallaxes (Lindegren et al. 2021)
- 1.8×10^6 G mags
- 1.5×10^6 BP & RB mags

→ See Recio-Blanco et al. 2023 + talk for standard spectroscopic analysis of RVS spectra

**Can we exploit in a homogeneous way
Gaia spectra (RVS + BP/RP)
magnitudes (G, Bp, Rp)
and parallaxes
for supercharged stellar
parametrization ?**

Analysis of the 1 million *Gaia* RVS-spectra with CNNs

Beyond *Gaia* DR3:
tracing the $[\alpha/M] - [M/H]$ bimodality
from the Inner to the outer Milky Way disc
with *Gaia* RVS and Convolutional Neural-Networks

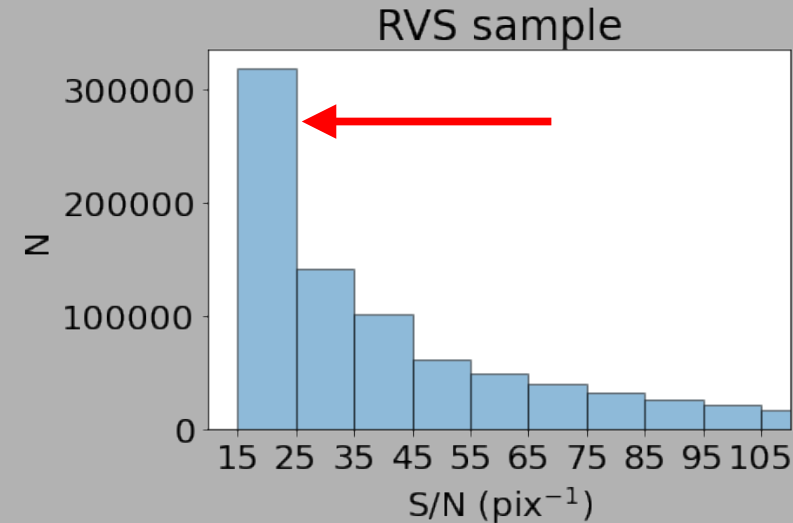
G. Guiglion¹, S. Nepal^{2,3}, C. Chiappini², S. Khoperskov², G. Traven⁴, A. B. A. Queiroz², M. Steinmetz²,
M. Valentini², Y. Fournier², A. Vallenari⁵, K. Youakim⁶, M. Bergemann¹, S. Mészáros^{7,8}, S. Lucatello^{9,10},
R. Sordo⁵, S. Fabbro¹¹, I. Minchev², G. Tautvaišienė¹², Š. Mikolaitis¹², J. Montalbán¹³



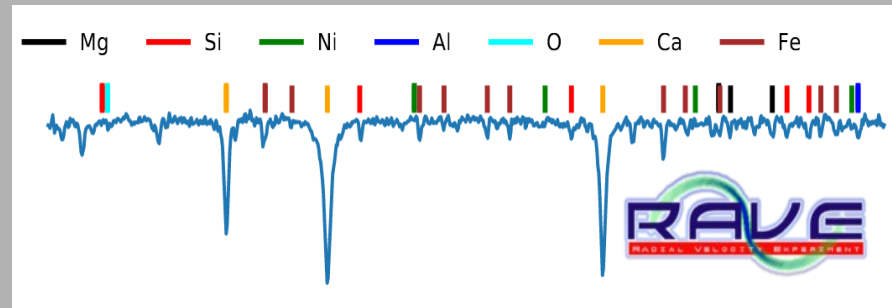
→ **Under revision :)**

Motivations and goals:

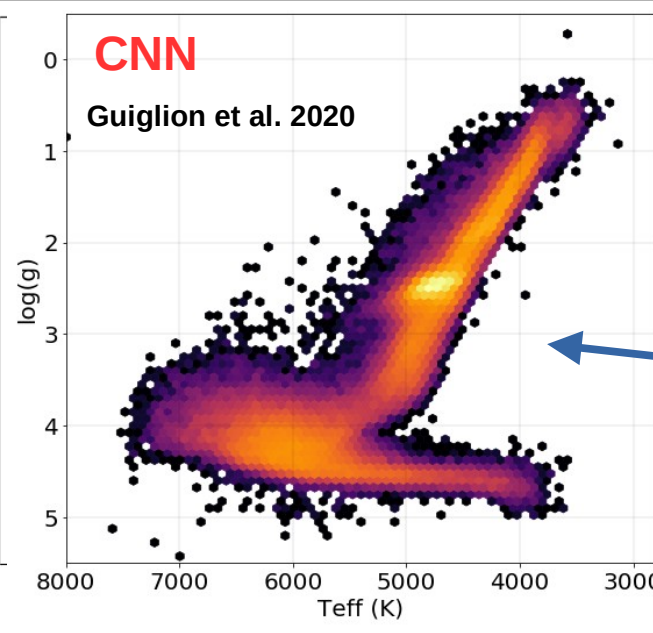
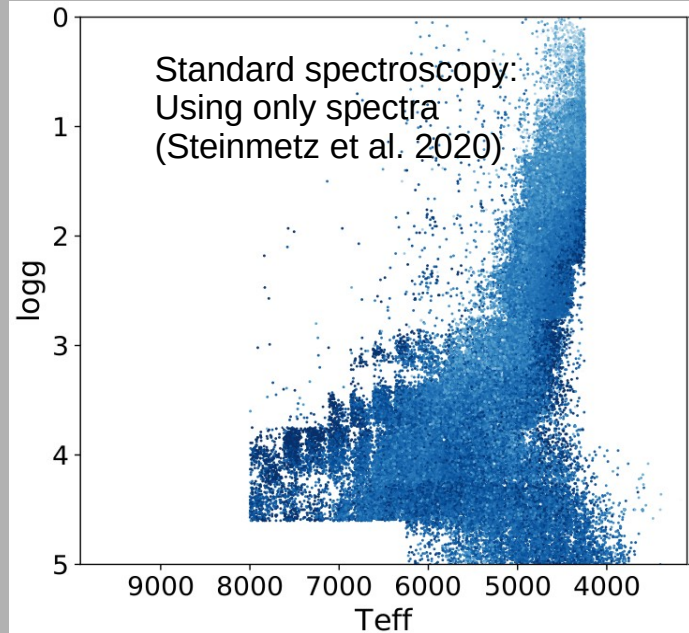
- Use homogeneously the full *Gaia* data product
- Leverage the low-S/N RVS sample
No GSP-Spec labels with 13 “good” flags within $15 < S/N < 25$
- Set the machine-learning path for *Gaia* data analysis
(DR4 in 2025, DR5 in 2027)



Our experience with CNNs and *Gaia*-like spectra



→ GG et al. 2020:
1st application of CNNs combining RAVE spectra,
Gaia magnitudes, and parallaxes



Breaking the spectral
degeneracies

→ Recent CNN developments in Nepal, GG et al. 2023 and Ambrosch, GG et al. 2023

Analysis of the 1 million *Gaia* RVS-spectra with CNNs

Training sample



R~22000

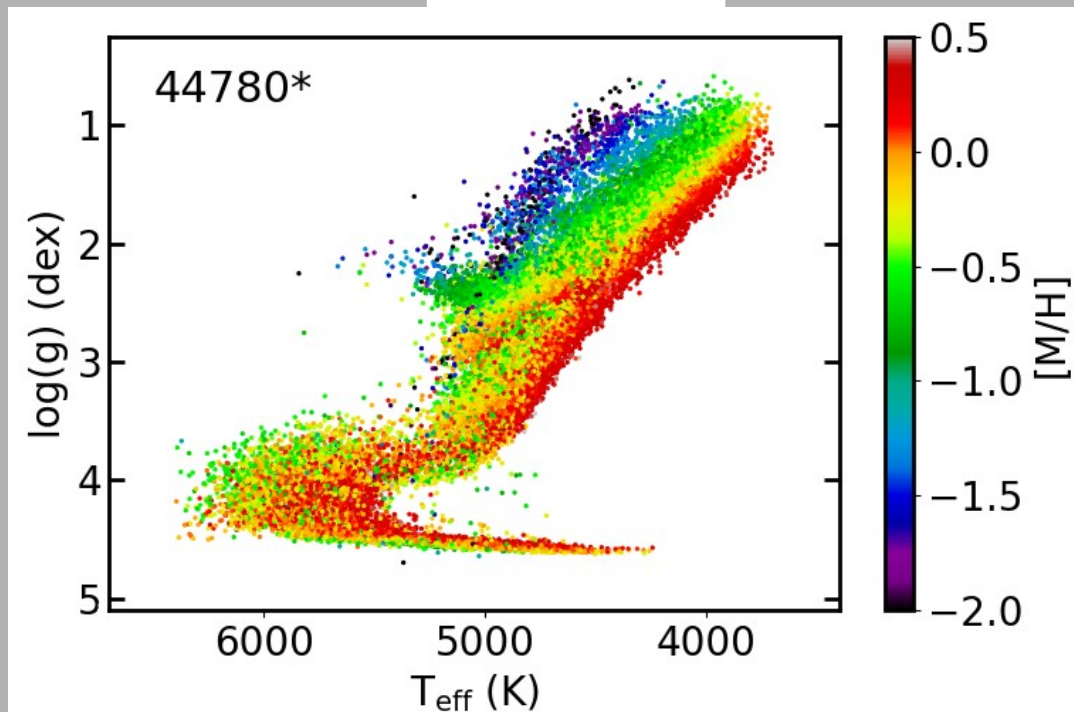


Knowledge transfer
from high-quality
high-res APOGEE labels
 T_{eff} , $\log(g)$, $[M/H]$, $[\alpha/M]$, $[Fe/H]$
to intermediate-res RVS



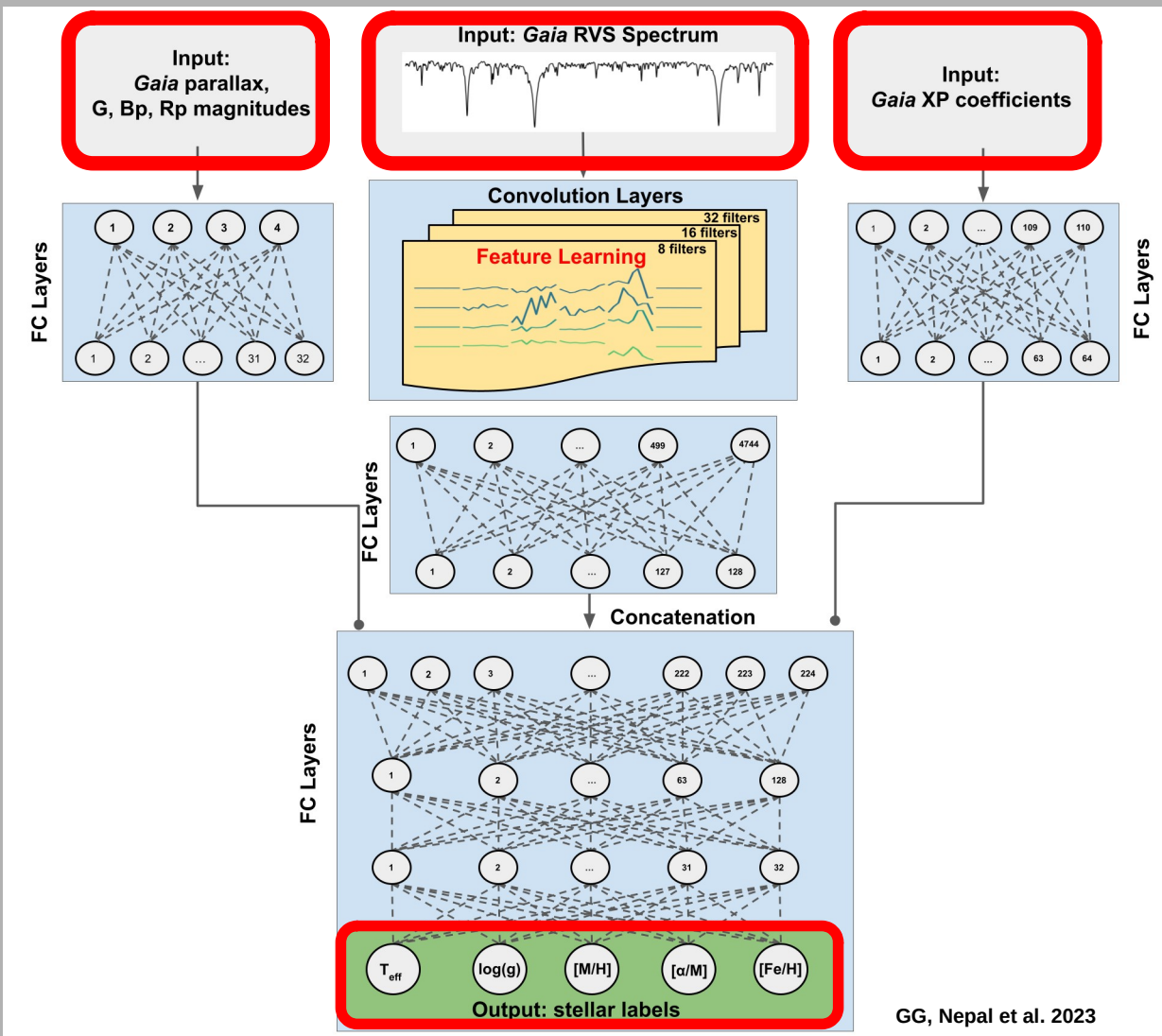
gaia R~11400

Training sample



GG, Nepal et al. 2023

A hybrid Convolutional Neural-Network for *Gaia*-RVS analysis

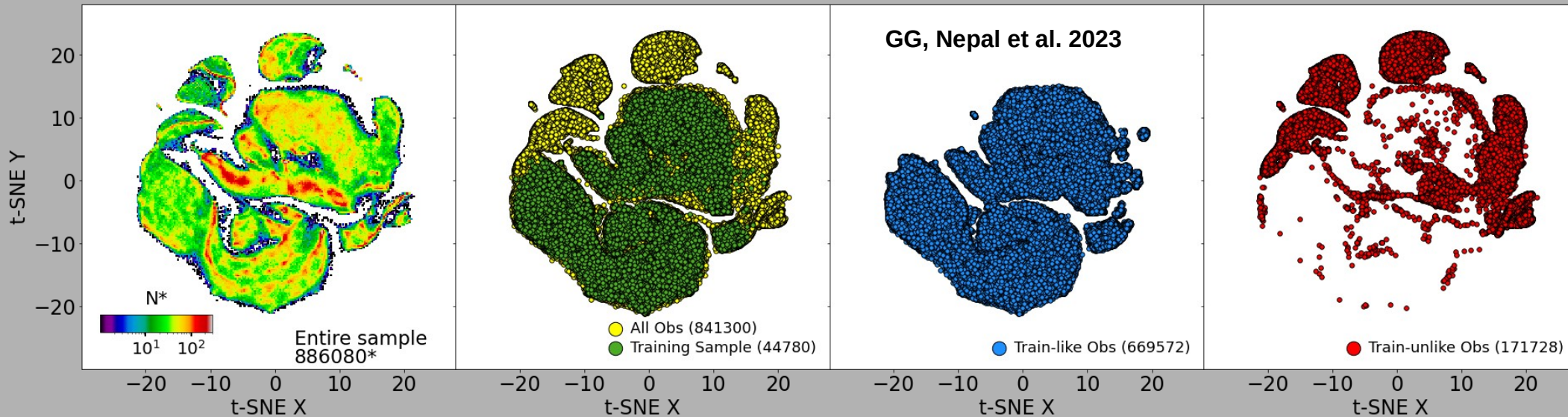


→ Observed sample:
841300 RVS stars

→ Prediction time
3300 stars / second

How to ensure that a label falls within the training sample limits ?

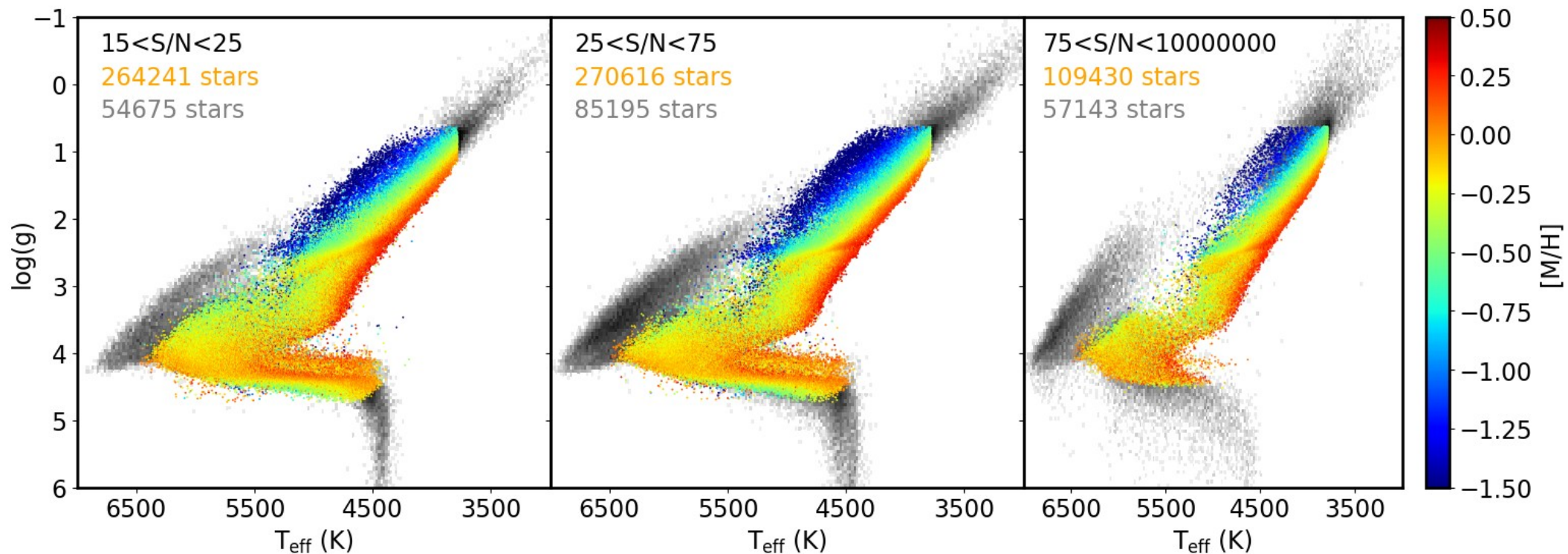
- Labels within T_{eff} , $\log(g)$, $[M/H]$, $[\alpha/M]$, $[Fe/H]$, G , and parallax limits of training sample.
- t-SNE classification of RVS spectra



→ **644287 RVS stars within TS**

Robust estimates of T_{eff} , $\log(g)$, $[M/H]$ for 690000 *Gaia* stars

GG, Nepal et al. 2023



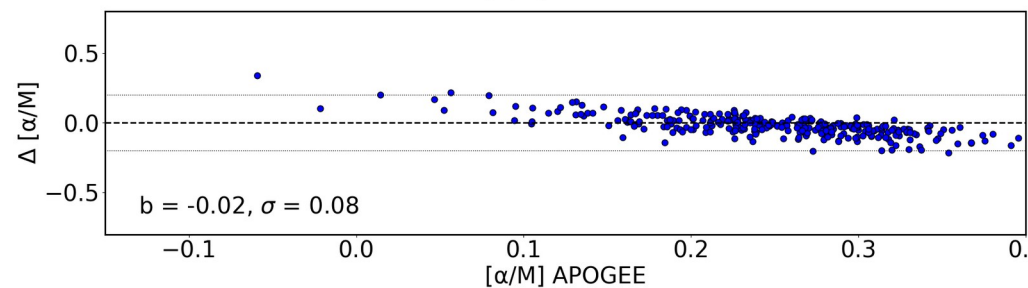
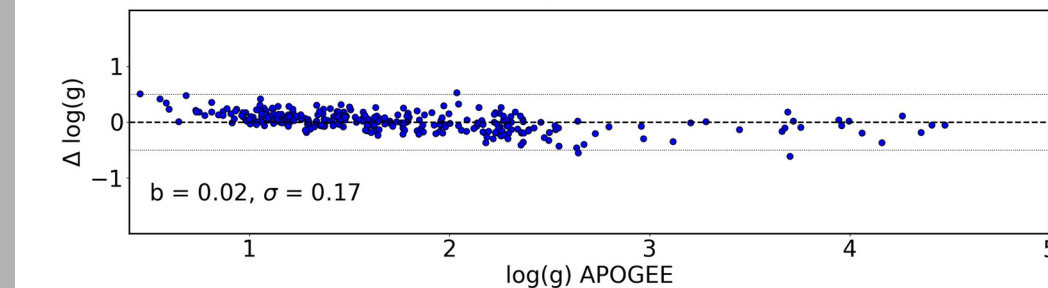
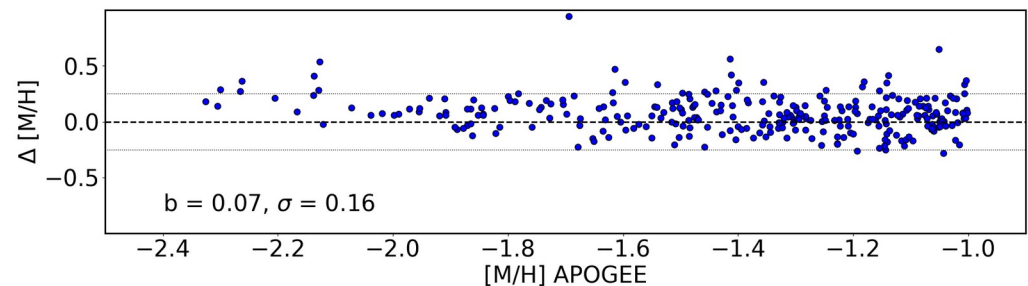
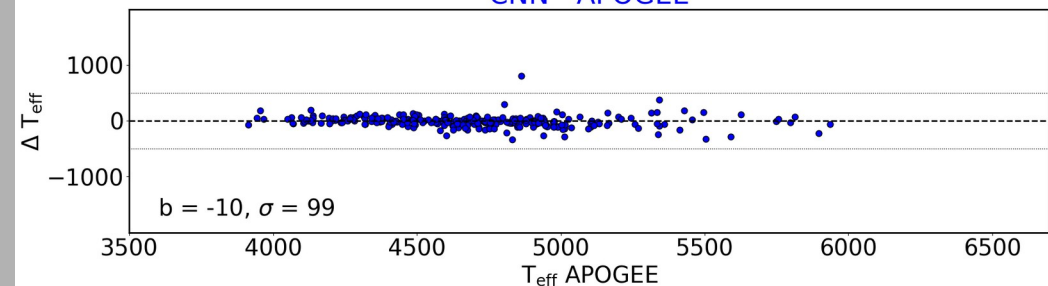
→ By adding magnitudes, parallaxes and XP data, CNN is able to break spectral degeneracies in *Gaia* RVS spectra.

→ CNN results are as good as the training set can be.

CNN performances for halo stars

→ $15 < S/N < 25$

CNN - APOGEE

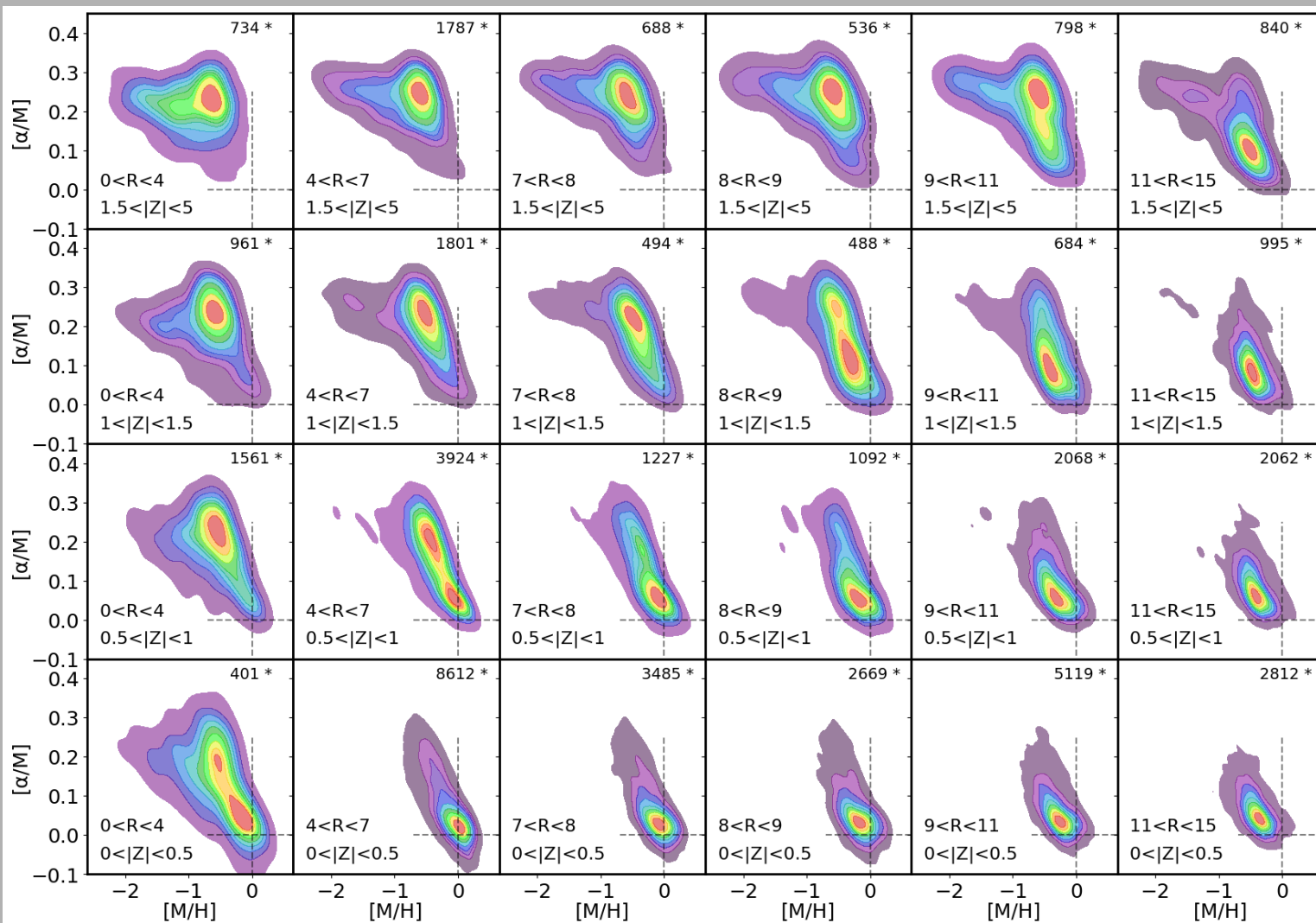


GG, Nepal et al. 2023

→ CNN provides precise and accurate labels down to $[M/H] = -2.4$ dex

Chemical cartography of the Milky Way, for Inner to Outer regions with *Gaia* and CNN

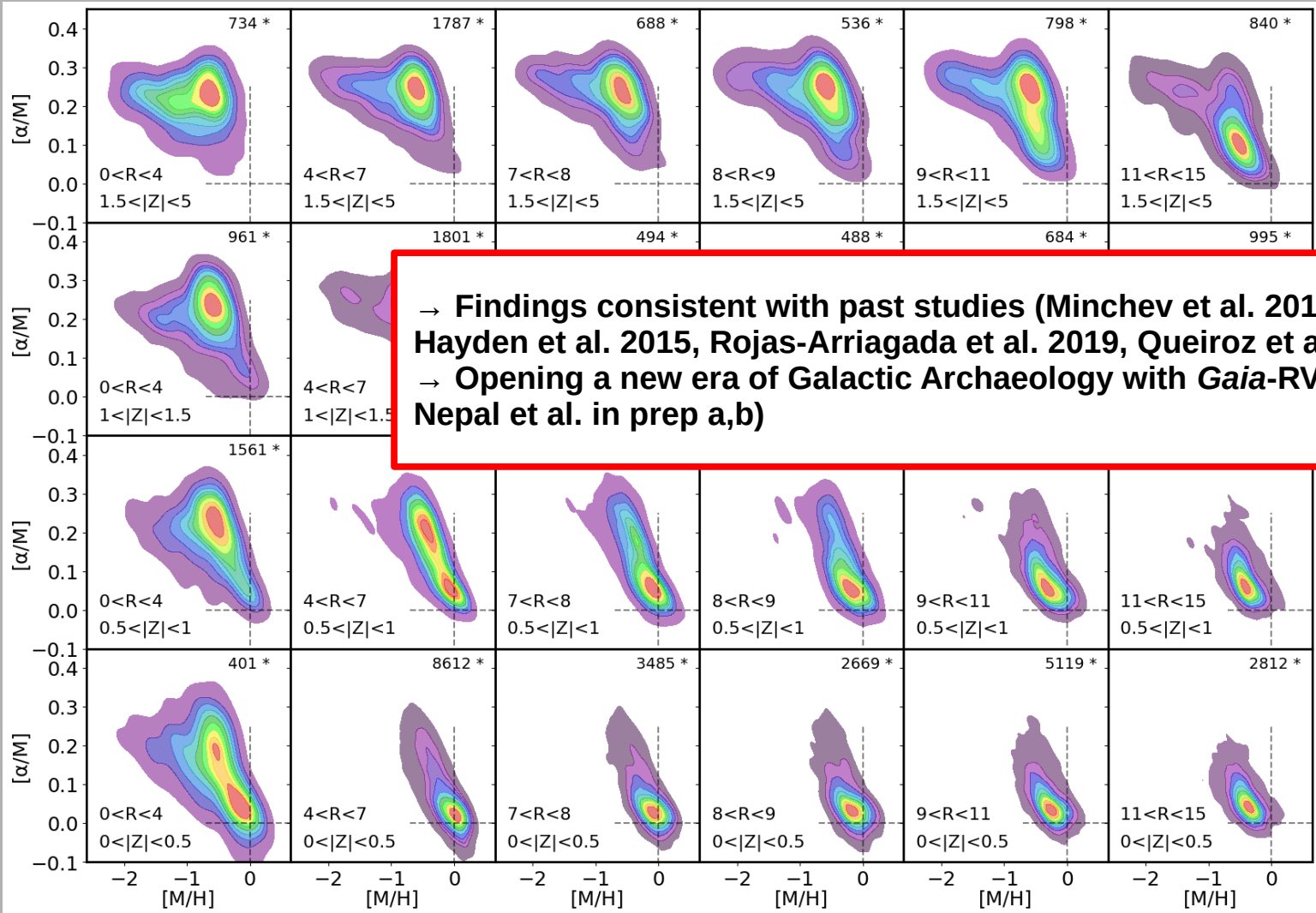
45838 stars
 $15 < S/N < 25$



GG, Nepal et al. 2023

Chemical cartography of the Milky Way, for Inner to Outer regions with *Gaia* and CNN

45838 stars
 $15 < S/N < 25$

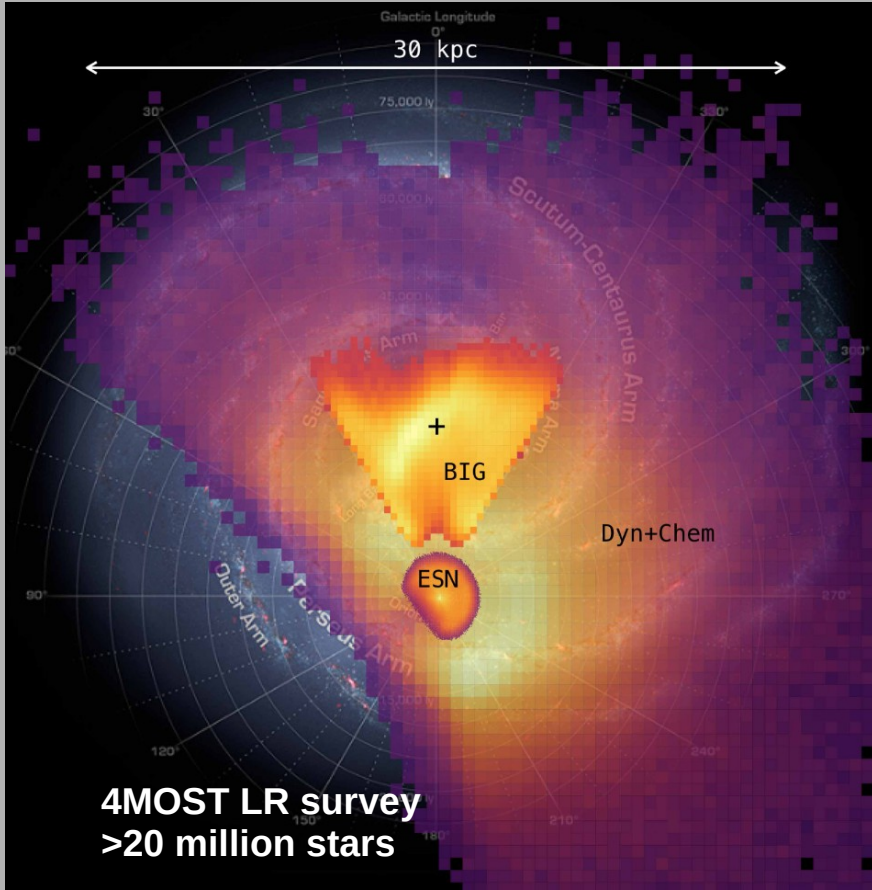


→ Findings consistent with past studies (Minchev et al. 2015, Anders et al. 2014, Hayden et al. 2015, Rojas-Arriagada et al. 2019, Queiroz et al. 2020, 2021).
→ Opening a new era of Galactic Archaeology with *Gaia*-RVS (Guiglion et al. In prep, Nepal et al. in prep a,b)

GG, Nepal et al. 2023

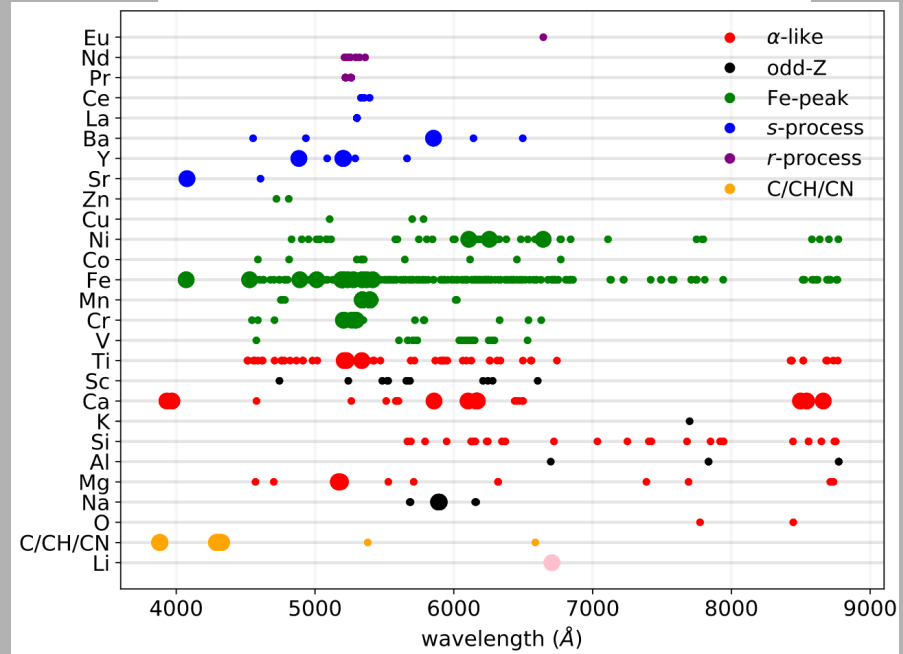
Why using CNN on low-res spectra ?

→ 4MIDABLE-LR Disc and Bulge surveys (Chiappini et al. 2019)



4MIDABLE-LR ESO proposal 2020

>20 elements to be measured at R=5000



4MIDABLE-LR ESO proposal 2020

→ Developing **CNN** for 4MIDABLE-LR D1(>) spectral analysis.

Summary:

- Hybrid CNN is an optimal method for combining full Gaia data product
 - Leveraging the large set of low S/N RVS spectra
- CNN parametrization is fast and robust (several 10^3 stars per second)

Insights:

- Future spectroscopic surveys will strongly benefit from CNNs
- Standard spec. and ML methods complement each other
- CNN parametrization mainly reliable within the training sample limits
 - The training sample should be built in a pro-active way



guiglion@mpia.de

