Reactive collisions of electrons with molecular cations of astrophysical interest: effects to H_2^+, HD^+, H_3^+

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The dissociative recombination (DR) together with the competing reactions – rovibrational excitation/de-excitation of the hydrogen molecular ion plays a decisive role in astrophysical ionized media: stars and interstellar molecular clouds, early Universe.

Using a stepwise method based on Multichannel Quantum Defect Theory (MQDT) [1], cross sections and rate coefficients have been obtained for reactions induced on $HD^+[2]$, $H_2^+[3]$ and $D_2^+[6]$.

For H_2^+ , the full rotational computations improved considerably, the accuracy of the resulting dissociative recombination cross sections and Maxwell isotropic rate coefficients [5]. The different mechanisms taken into account for H_2^+ , i.e. direct vs indirect and rotational vs non-rotational processes are presented.

An analytic three-channel model was developed for the description of simultaneous direct and indirect DR cross sections of H_3^+ [4].

For these analyzed systems the results is in good agreement with the CRYRING (Stockholm) and TSR (Heidelberg) *magnetic* storage ring results.

References:

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