

MULTIPLE PHASE SPIRALS SUGGEST MULTIPLE ORIGINS IN GAIA DR3

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Barcelona - 7/9/23

Striking spiral patterns in z - v_z

- Phase spiral, or snail shaped features in vertical position vs. motion suggest ongoing vertical phase mixing.

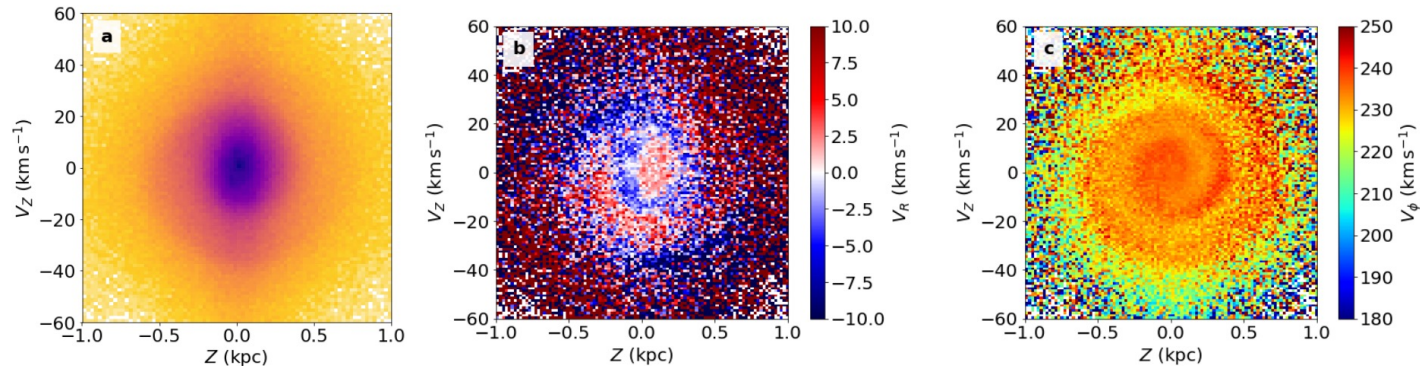
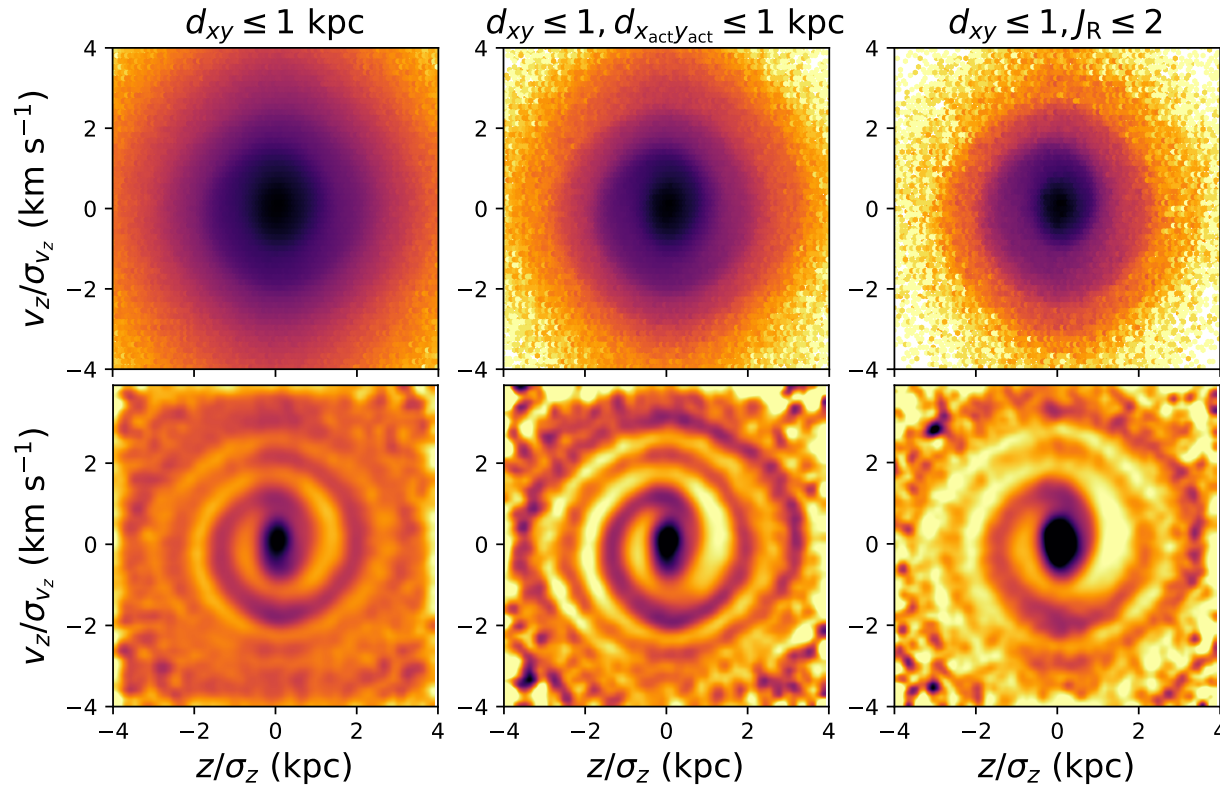
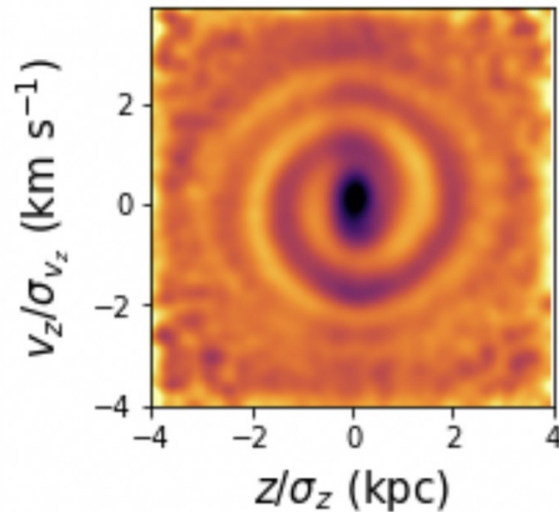
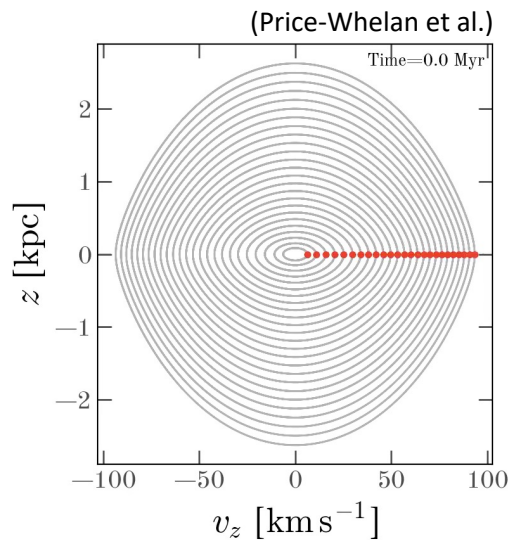


Figure 2. Distribution of stars in the vertical position-velocity plane Z - V_z for stars selected as in Fig. 1. a) Two-dimensional histogram in bins of $\Delta Z = 0.01$ kpc and $\Delta V_z = 0.1$ km s^{-1} ; b) Z - V_z plane coloured as a function of median V_R in bins of $\Delta Z = 0.02$ kpc and $\Delta V_z = 1$ km s^{-1} ; c) Same as b) but for V_ϕ .

Updated for Gaia DR3

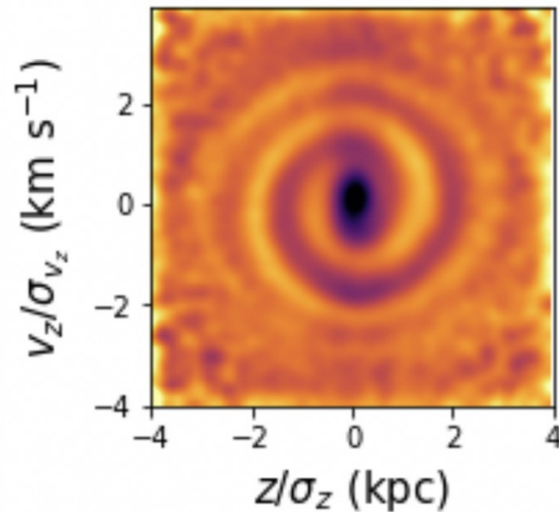
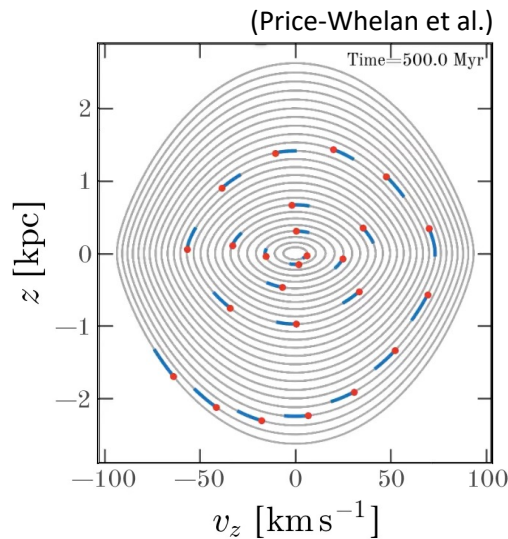


How/why do such spirals form?



Higher vertical action \rightarrow Lower vertical frequency

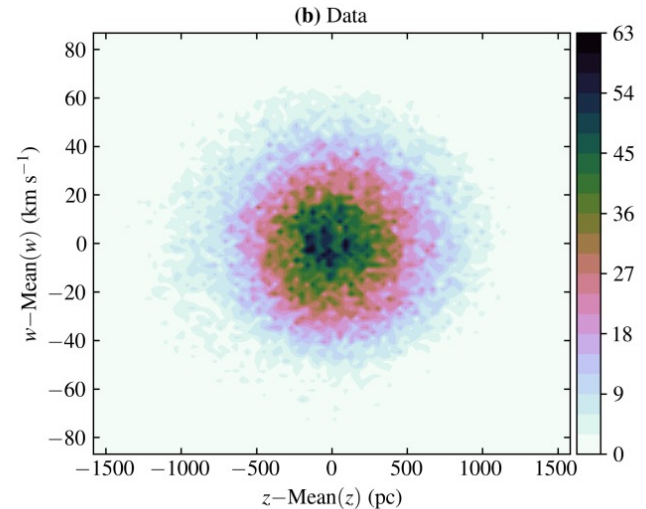
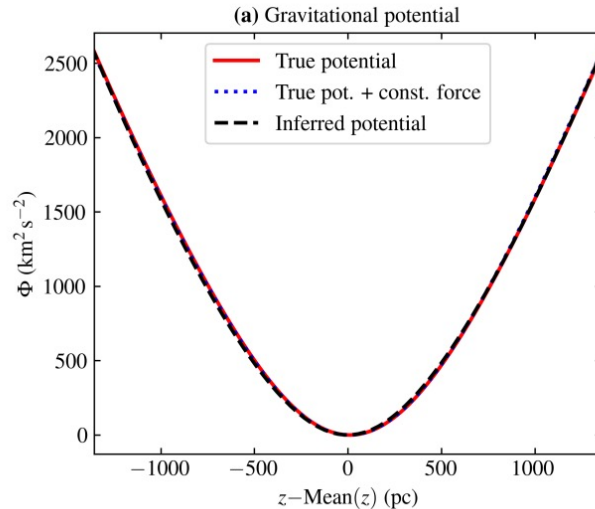
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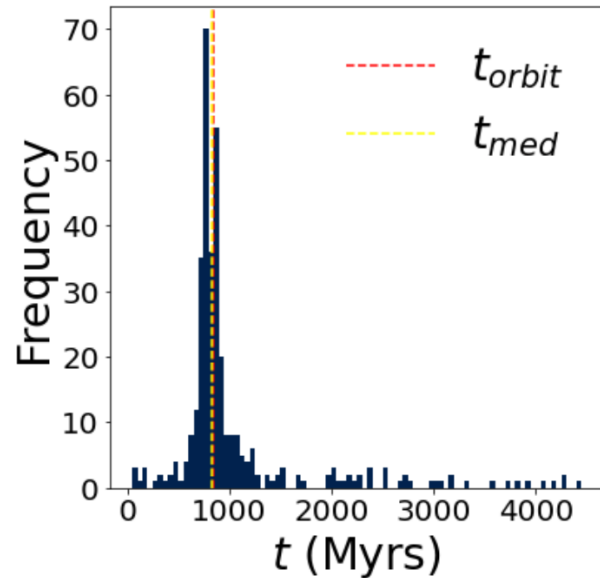
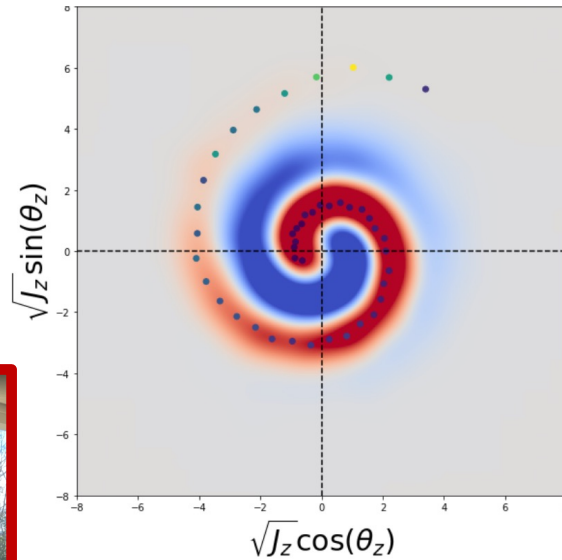
These spirals contain information about our Galaxy

- The spiral can be used to infer information about the potential



These spirals contain information about our Galaxy

- If we know the potential, we can unwind the spiral (in theory)

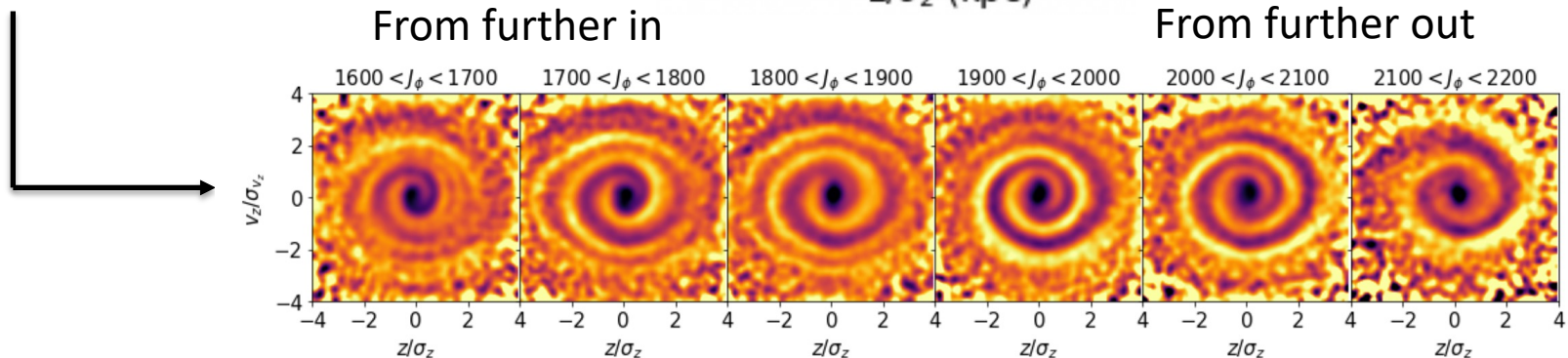
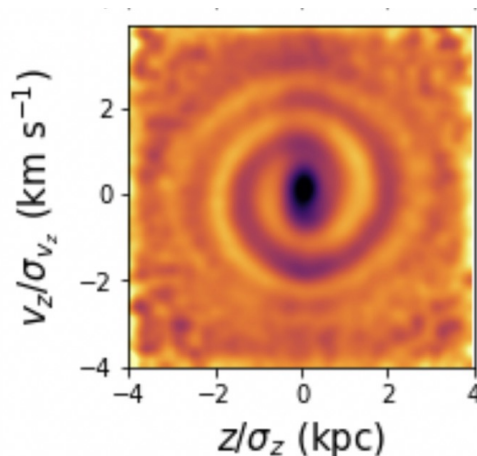


In practice, it's not quite this simple

- Self gravity of the disc (e.g. Widrow 2023)
- Older structure / previous interactions
- GMC's causing diffusion & erasing spirals (Tremaine+ 2022)
- Multiple perturbations? (Hunt+ 2022)
- Superposition of spirals (Li+ 2021, Hunt+ 2021, Gandhi+ 2022)

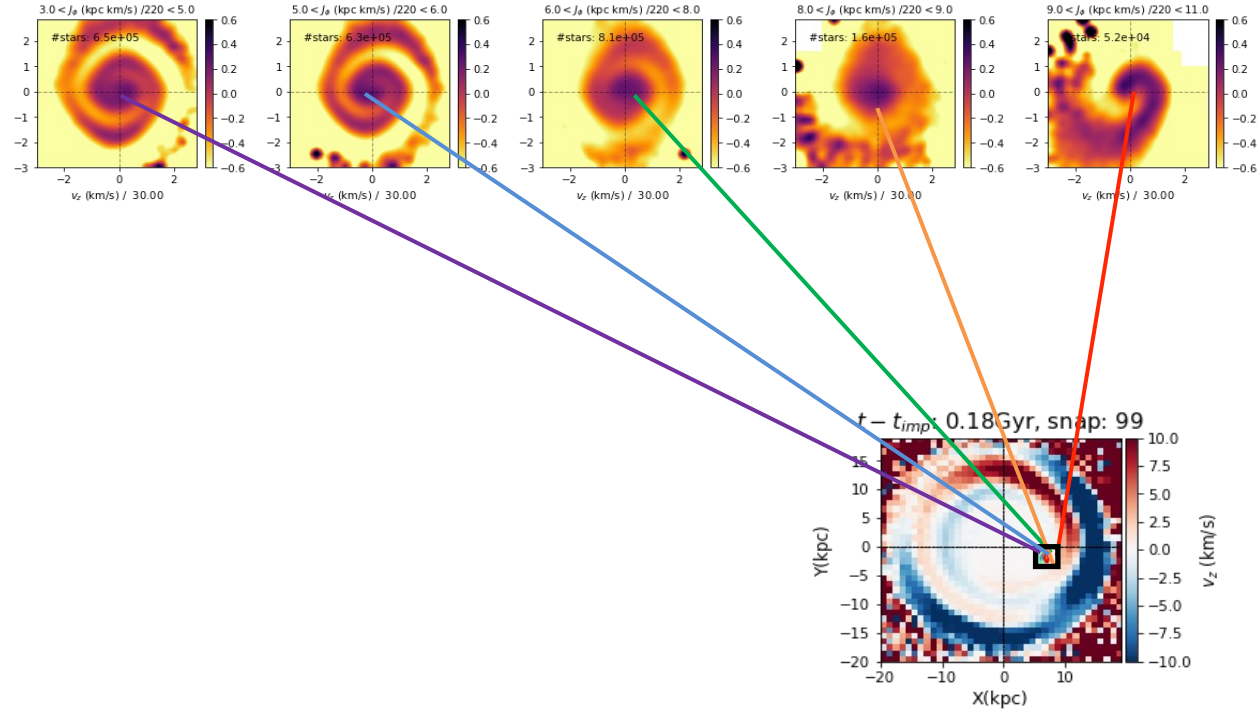
A given volume contains many spirals!

- Gaia DR3 data within 1 kpc -->
- Stars on many different orbits
- Split by angular momentum



They probe other regions of the disc

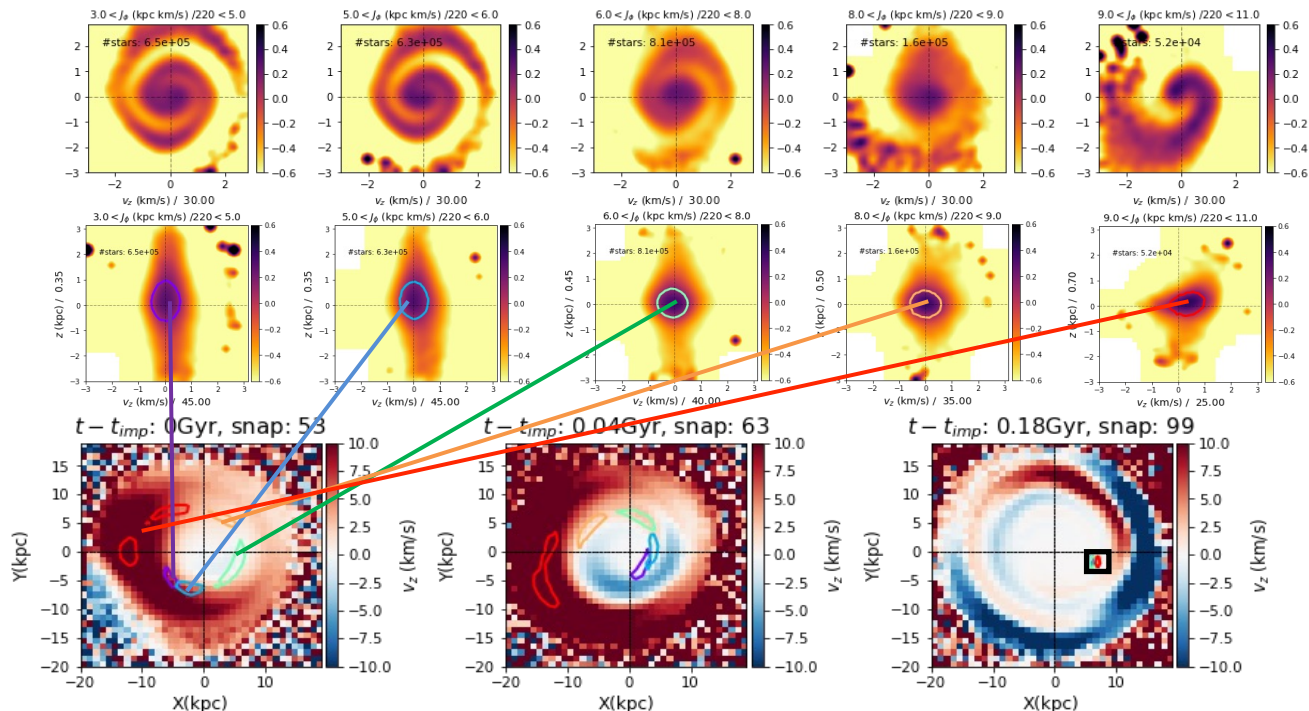
- High resolution test particle model from Gandhi et al. (2021)
- Split by angular momentum, different morphology and strength



Gandhi, Jonhston, Hunt et al. (2021)

They probe other regions of the disc

- High resolution test particle model from Gandhi et al. (2021)
- Stars near the Sun today were spread across the disc at the time of impact

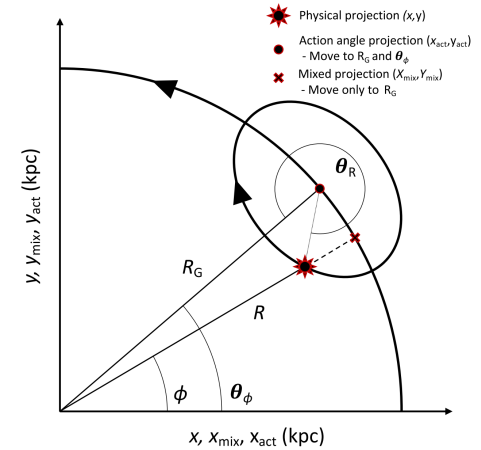
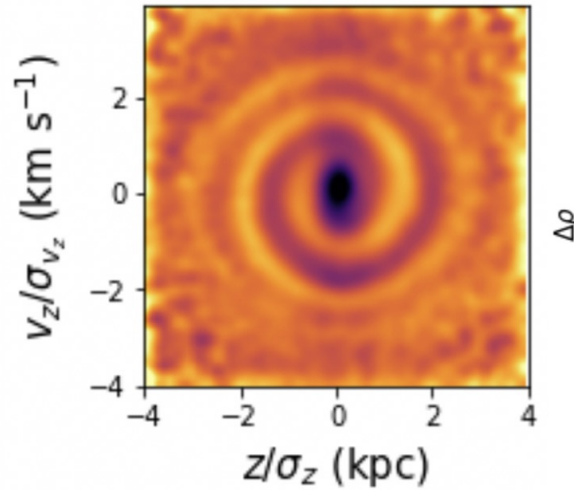


Suroor Gandhi; Graduate student at Columbia

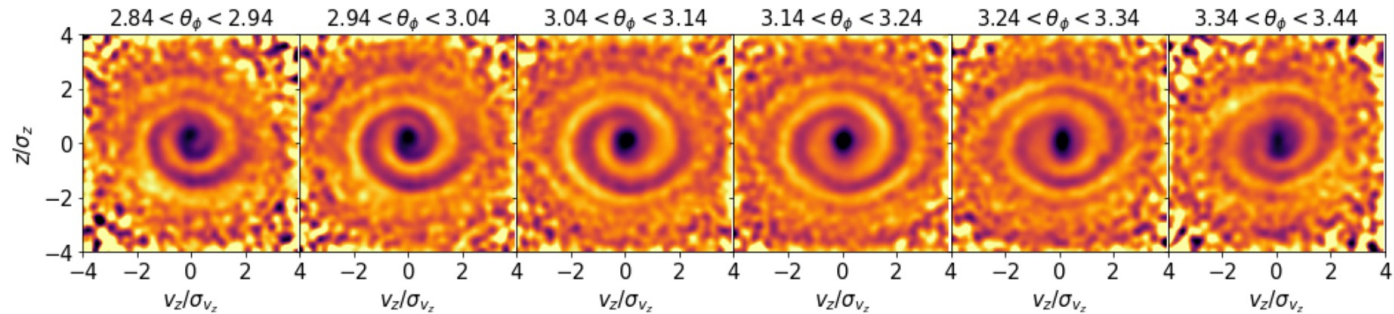
Gandhi, Jonhston, Hunt et al. (2021)

A given volume contains many spirals!

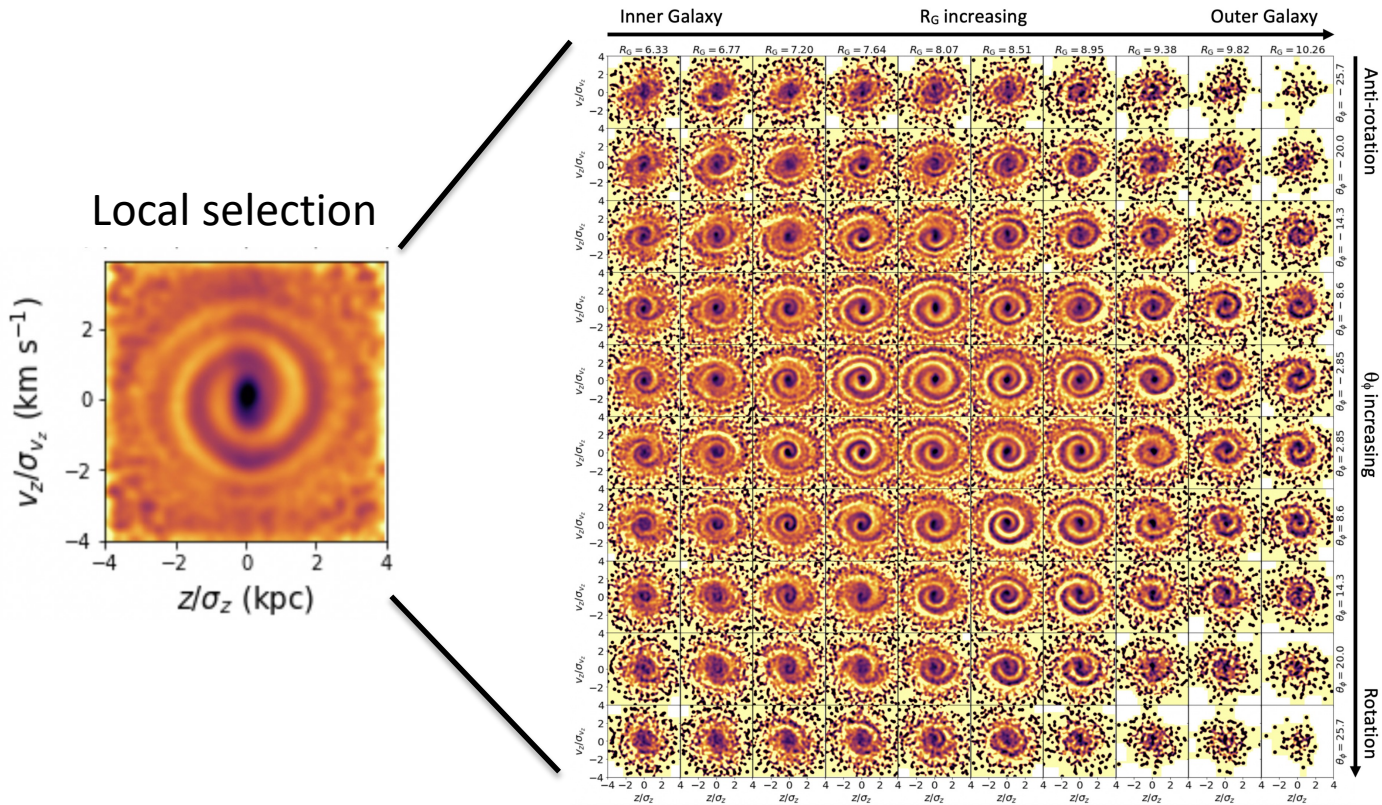
- Gaia data within 2 kpc from the Sun ->
- It contains stars on many different orbits



- Split by θ_ϕ ->



So if we split up the data in L_z and θ_ϕ

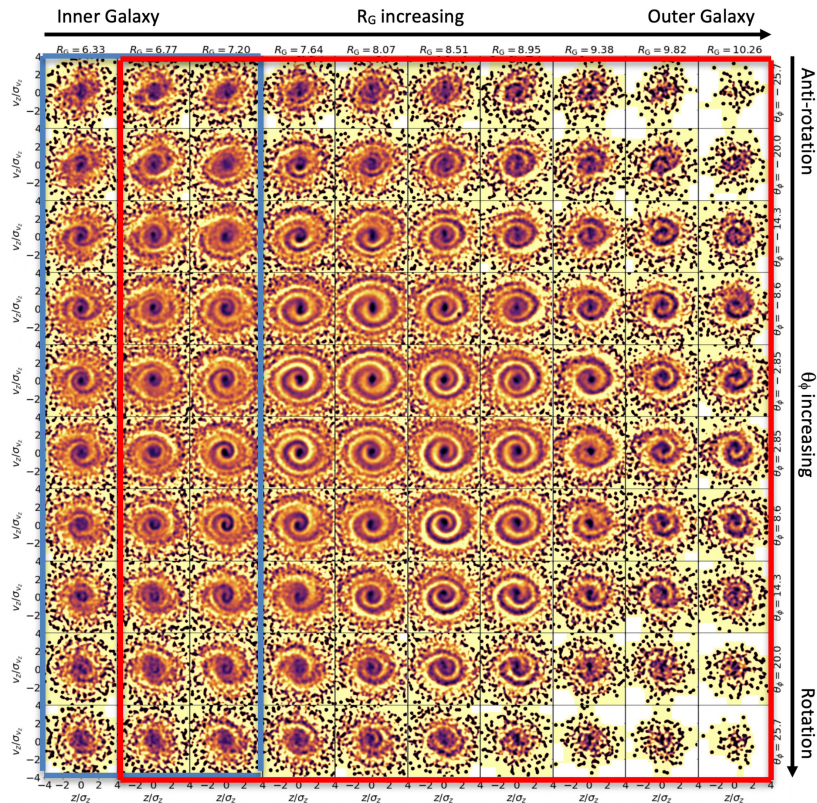


The spirals in the inner disc are 2-armed

(In the inner disc...)

Breathing mode
(inner galaxy)

(E.g. Williams et al. 2013,
Carrillo et al. 2018)

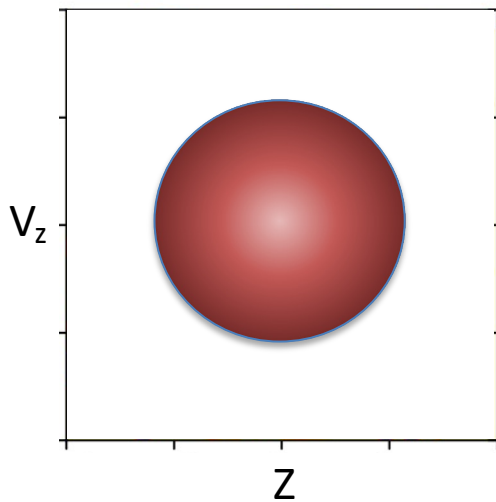


Bending mode
(Solar neighborhood
& outer Galaxy)

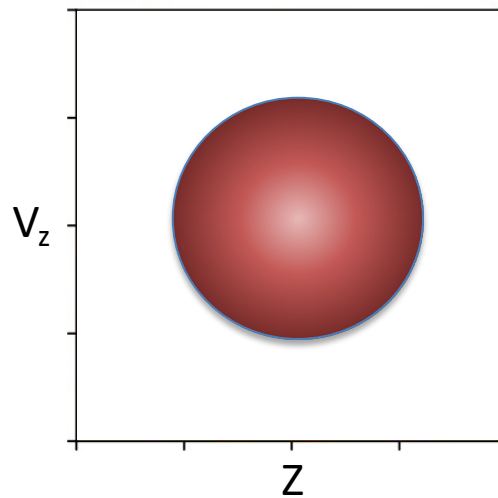
Bending vs. Breathing

Pre-interaction, everything nice and symmetric

Breathing mode



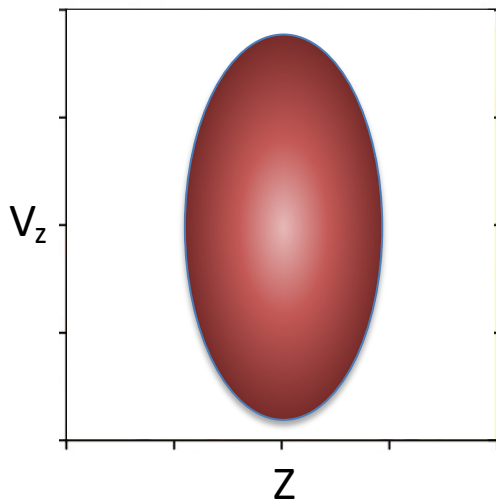
Bending mode



Bending vs. Breathing

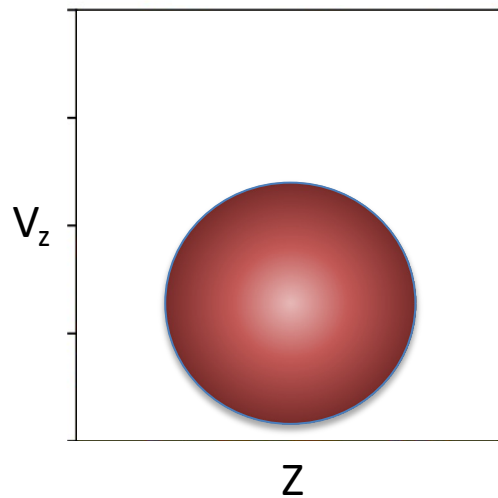
Immediately post-interaction

Breathing mode



Dynamically fast interaction

Bending mode

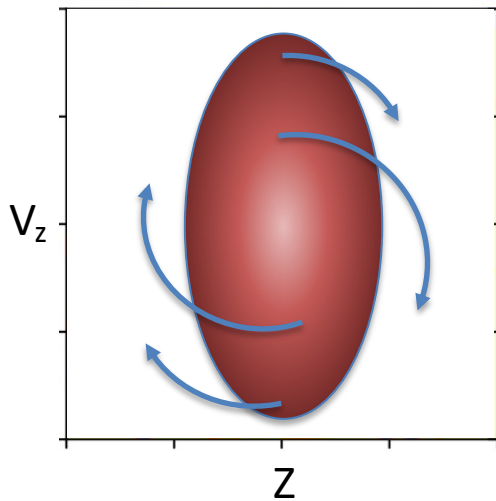


Extended interaction

Bending vs. Breathing

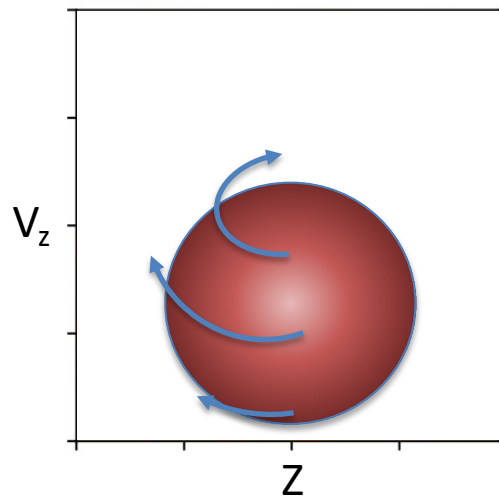
Immediately post-interaction

Breathing mode



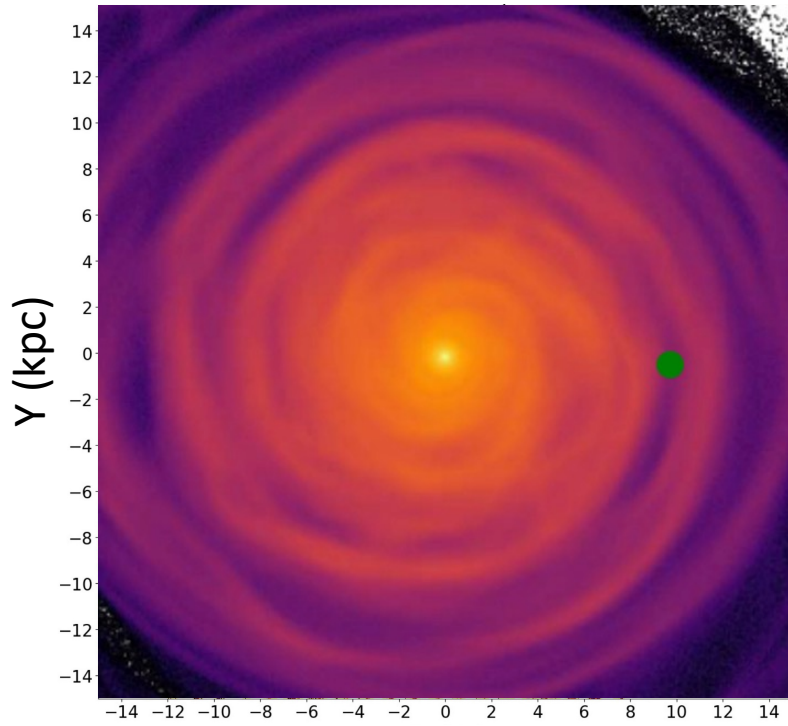
Dynamically fast interaction

Bending mode

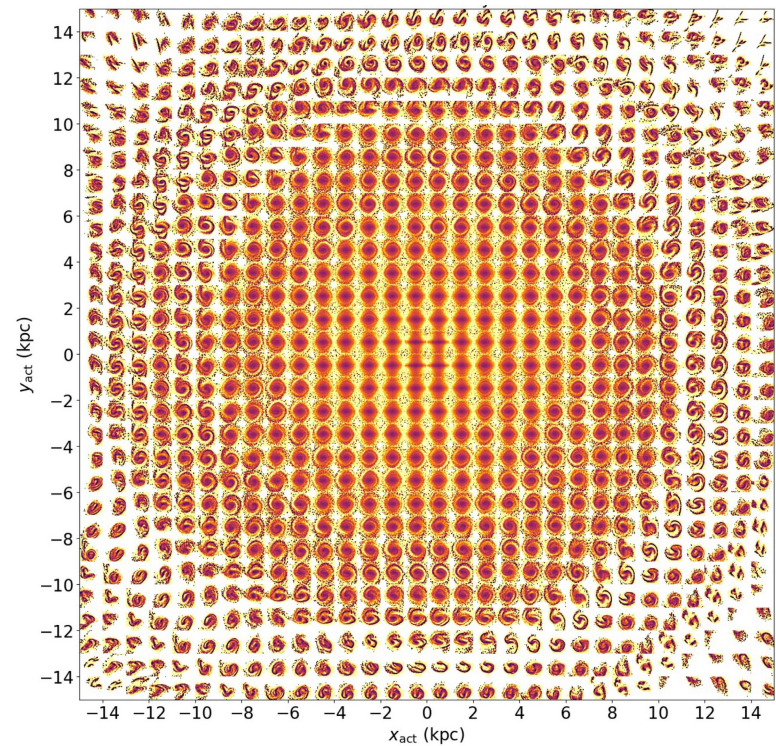


Extended interaction

Simulations of satellite mergers make phase spirals



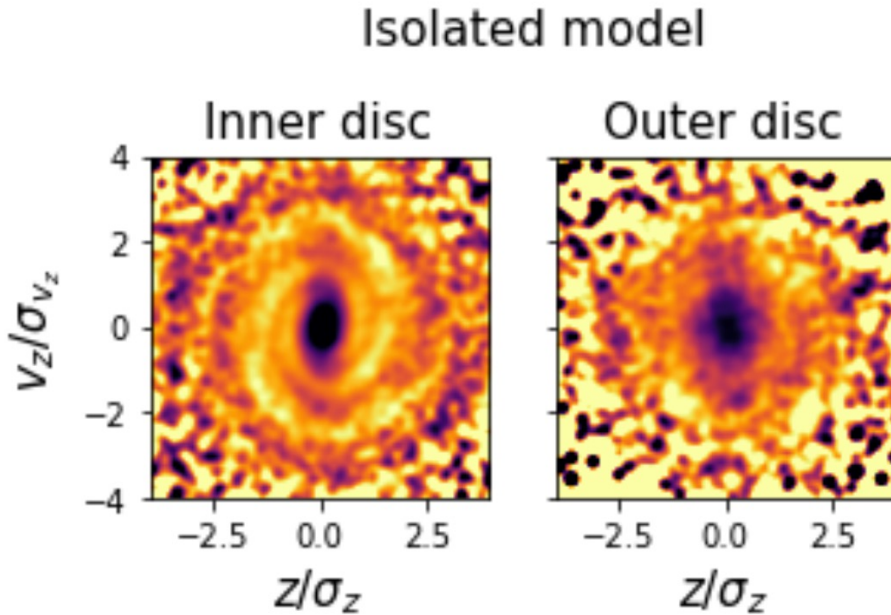
X (kpc)



Hunt et al. (2021)

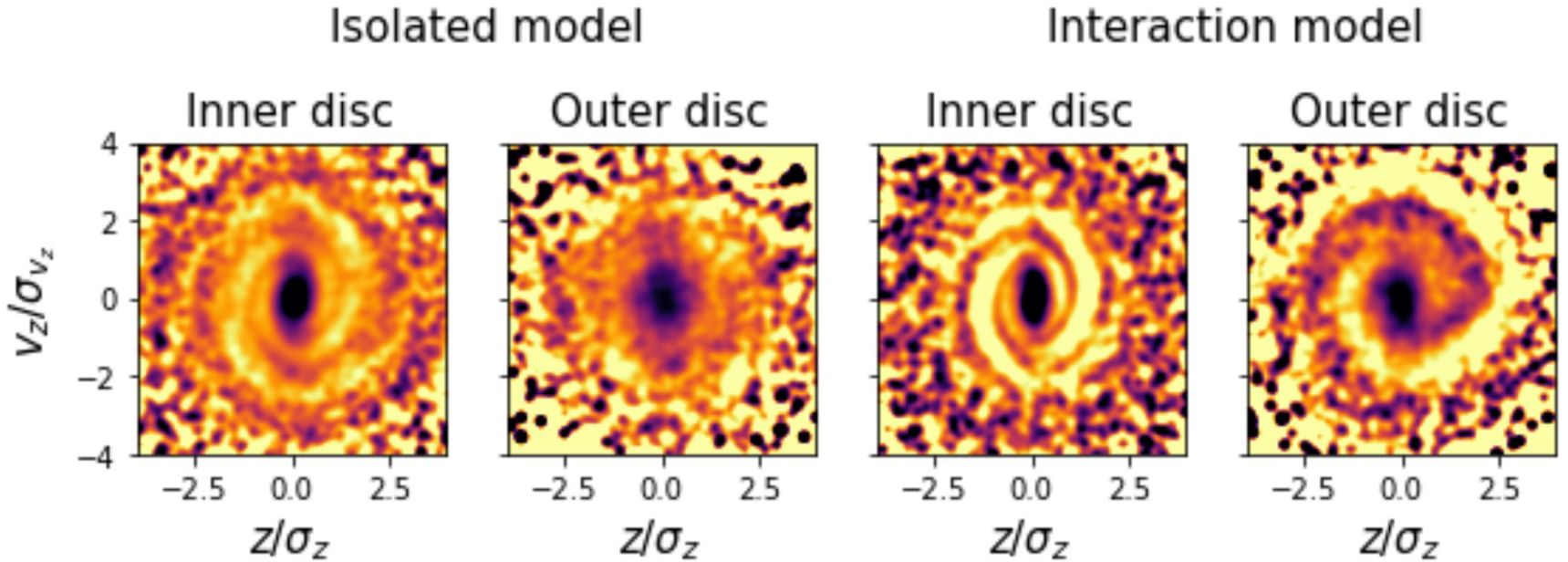
See also many other groups; Antoja, Laporte, Bland-Hawthorn, Li, Grand and more...

(Some) Isolated models make 2-armed spirals



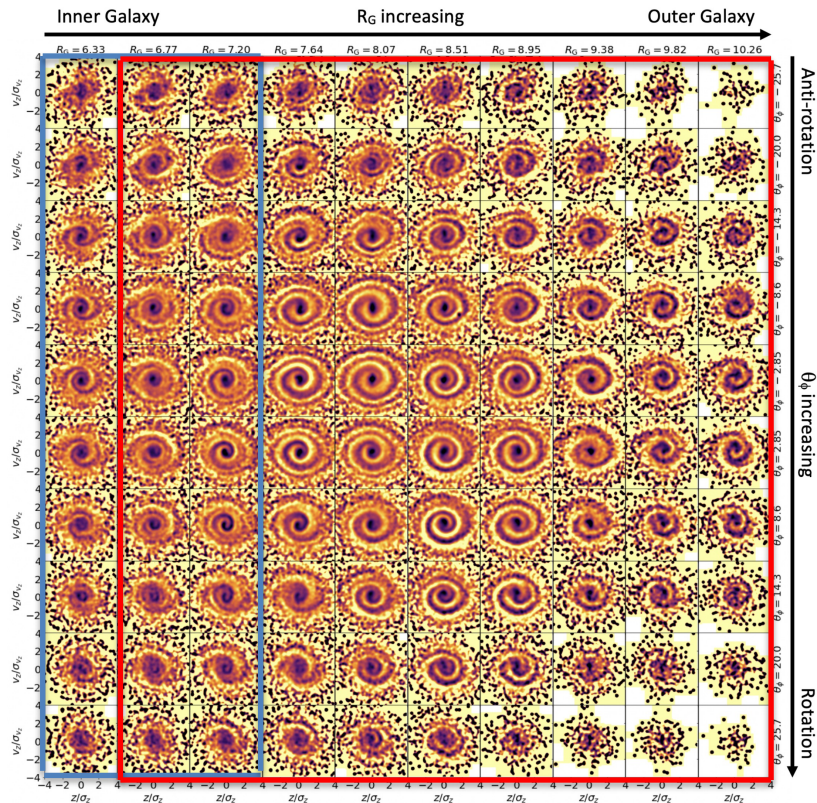
- Barred spiral N-body model which generates a breathing mode in the inner disc
- (The bar does not buckle)
- They grow and decay periodically

(Some) Isolated models make 2-armed spirals



Transition from internally to externally induced

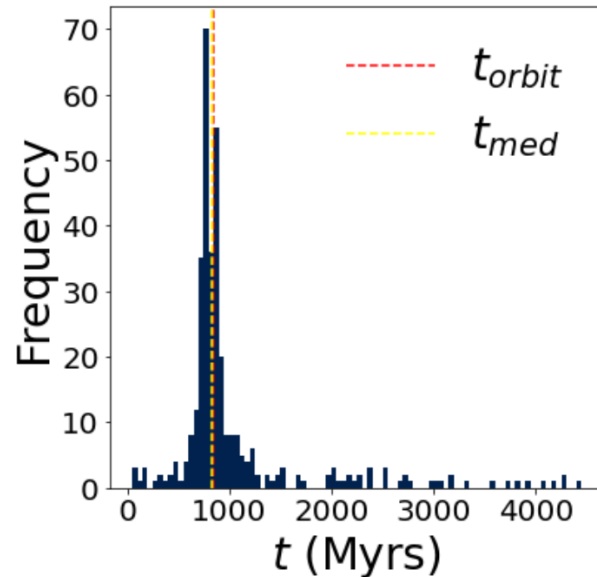
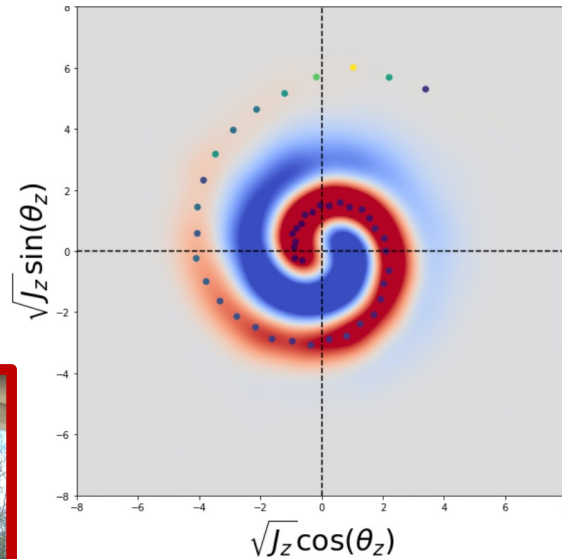
Bar/spiral arm induced?
Breathing mode
(inner galaxy)



Satellite induced?
Bending mode
(Solar neighborhood
& outer Galaxy)

These spirals contain information about our Galaxy

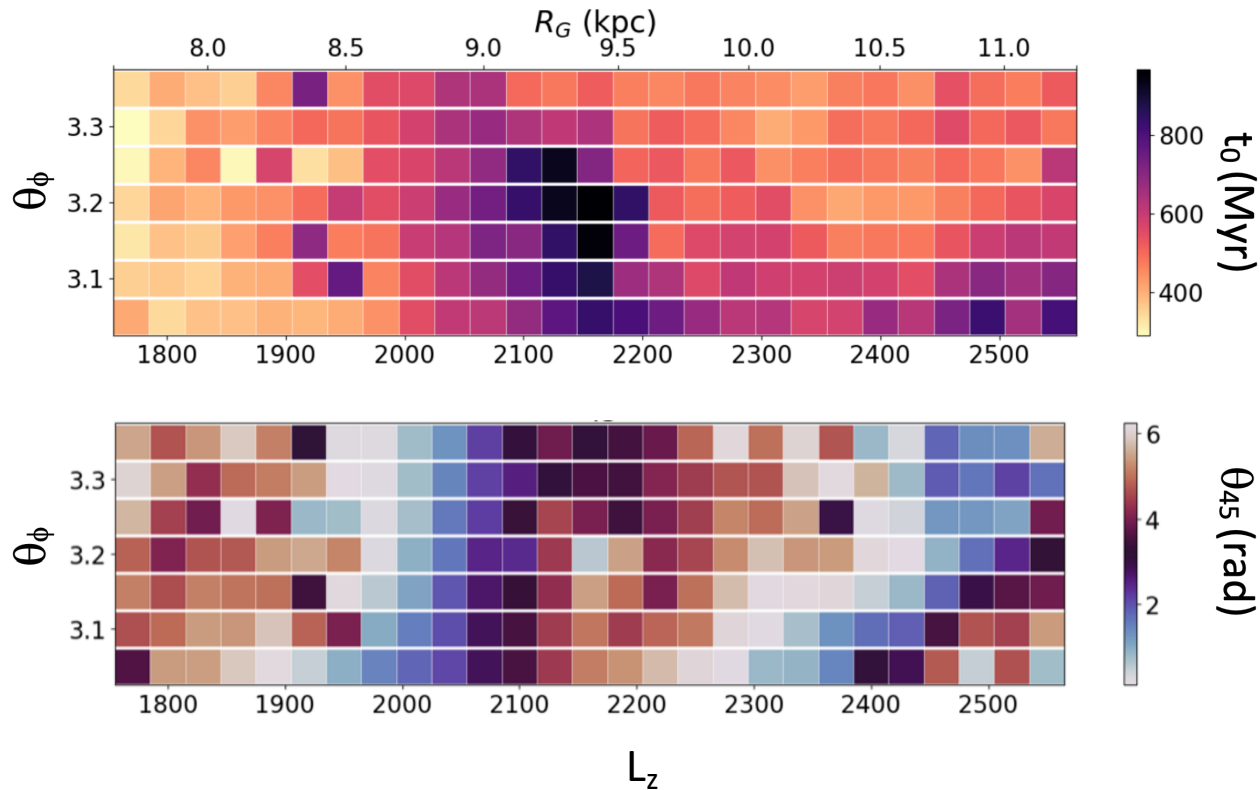
- If we know the potential, we can unwind the spiral (in theory)



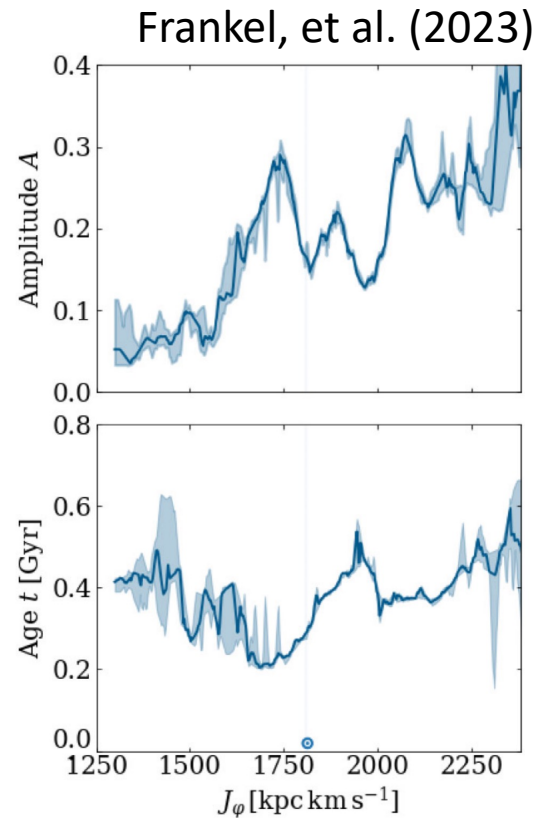
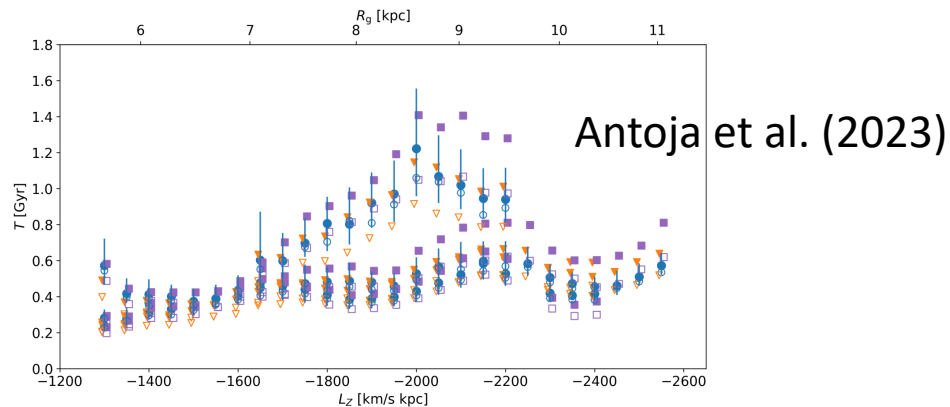
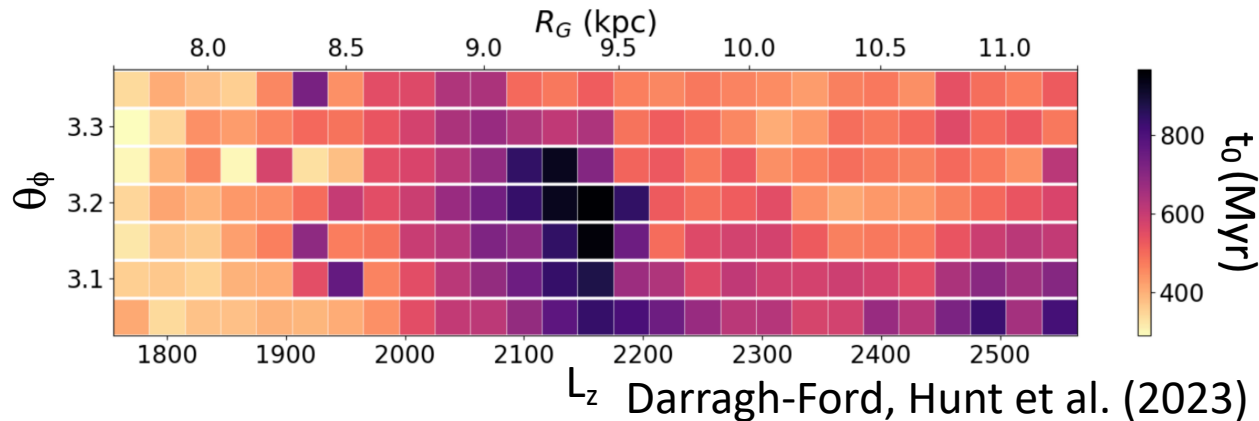
Phase spirals in Gaia DR3 show coherent patterns

(in the one-armed spirals)

- Range of impact times
- Multiple events, or a complex response



Independent estimates find similar patterns

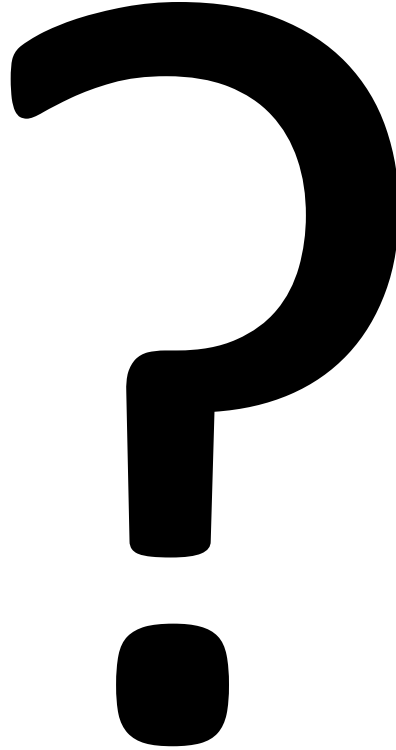


So, what's making the spirals?

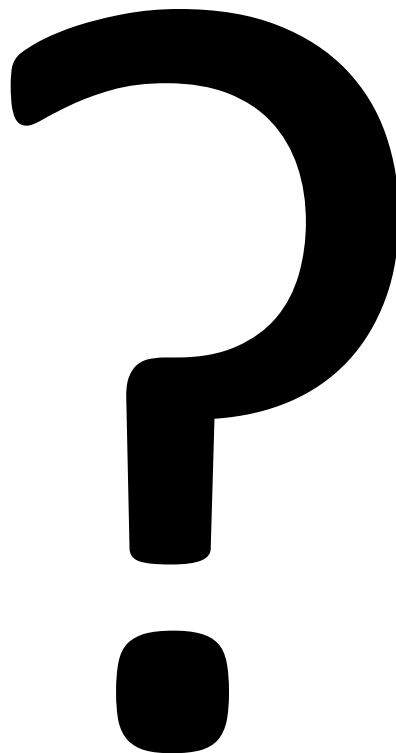
- The Sagittarius dwarf? - E.g. Hunt, Antoja, Laporte, Bland-Hawthorn +more
- Any other recent/current dwarfs? - Banik et al. (2022) says no...
- The Galactic bar?
 - Buckling? - Khoperskov et al. (2020)
 - Resonant growth? - Hunt et al. (2022)
- Spiral structure? - Hunt et al. (2022), Li et al. (2023)
- Dark Matter subhalo population - Tremaine et al. (2022)
- MW Dark Matter halo response - Grand et al. (2022)

But also

- GMC's can erase them through J_z diffusion - Tremaine et al. (2022)



Ask me again after DR4!



Elise Darragh-Ford, Stanford



Thanks to my collaborators!

(whose work is relevant to this talk!)

Uddipan Banik, Yale

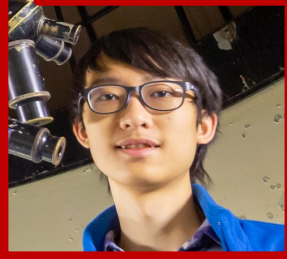


Kathryn Johnston, Columbia

Mathew Bub, Perimeter



Jo Bovy, Toronto



Henry Leung,
Toronto



Christina Eilers, MIT



Suroor Gandhi, NYU



Morgan Bennett,
Toronto



Wilma Trick,
MPIA



David Hogg, NYU



Adrian Price-Whelan, CCA



Ioana Stelea, Columbia



Daisuke Kawata, UCL



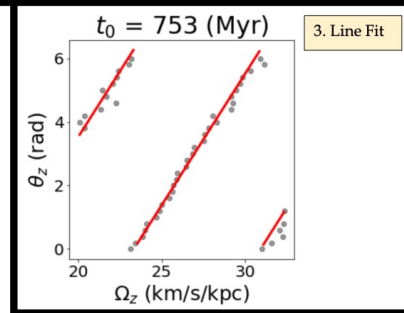
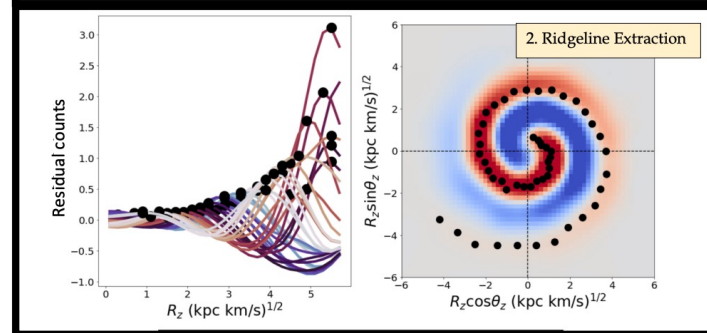
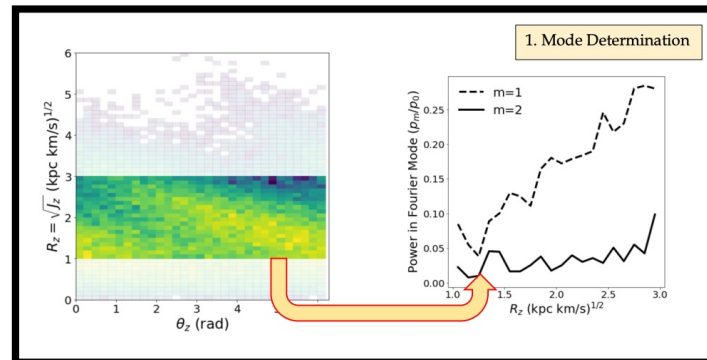
Chervin Laporte,
Barcelona

Summary

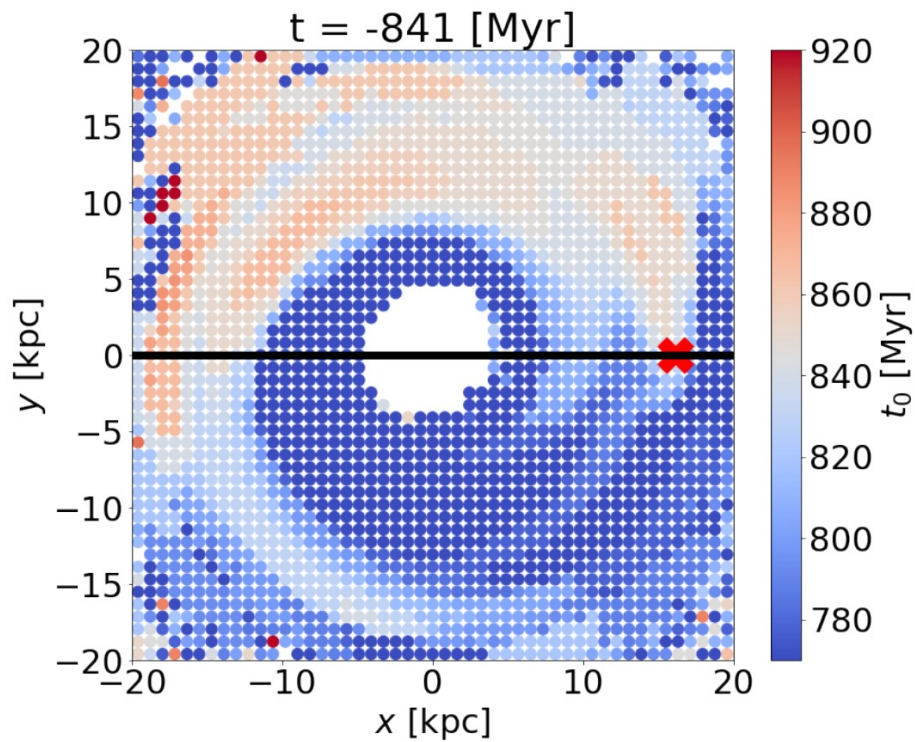
- Gaia has revealed a Galaxy in disequilibrium
- Features explainable by, the Bar, spiral arms, or mergers
- High resolution models let us examine local phase space structures without selection effects
- Gaia DR3 reveals a transition from ‘breathing spirals’ to ‘bending spirals’ (Hunt et al. 2022)
- Tracking the spirals should inform on the nature of the interaction *and* the structure of the disc



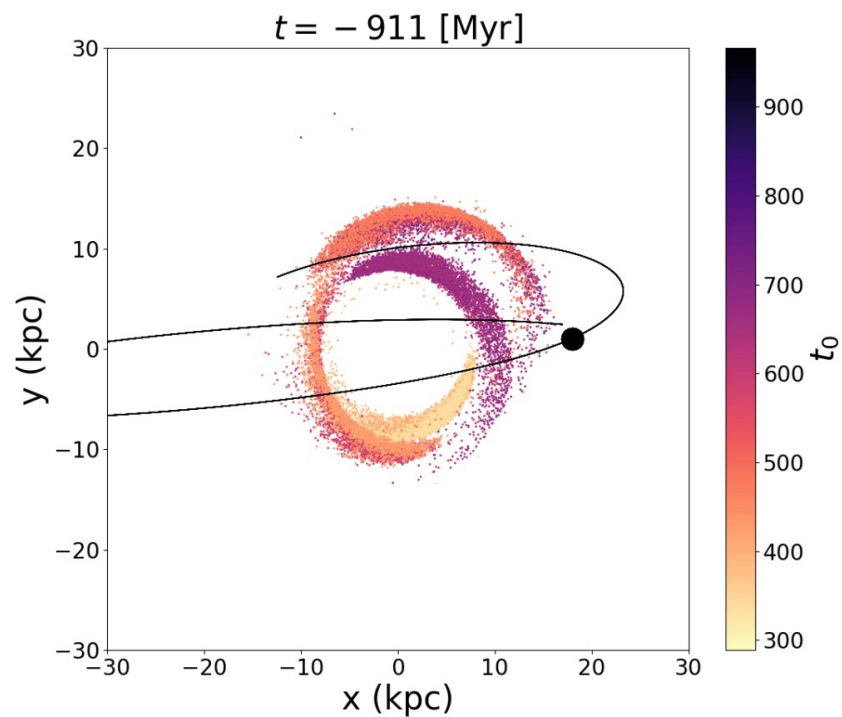
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Simulation



Gaia data





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