

# International Symposium on Nuclei in the Cosmos XVIII

Girona, June 17, 2025

## The Role of Carbon-Oxygen Shell Interactions in the Nucleosynthesis and Final Fate of Massive Stars

**Lorenzo Roberti**, INFN – Laboratori Nazionali del Sud, Catania, Italy

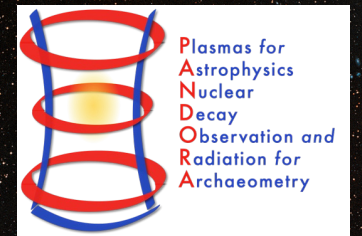
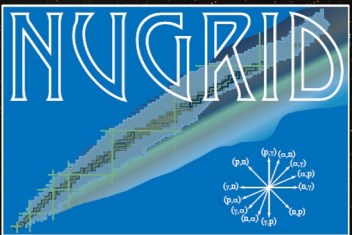
**Main collaborations:**

AsFiN + Pandora groups

Budapest group: M. Pignatari, M. Lugaro

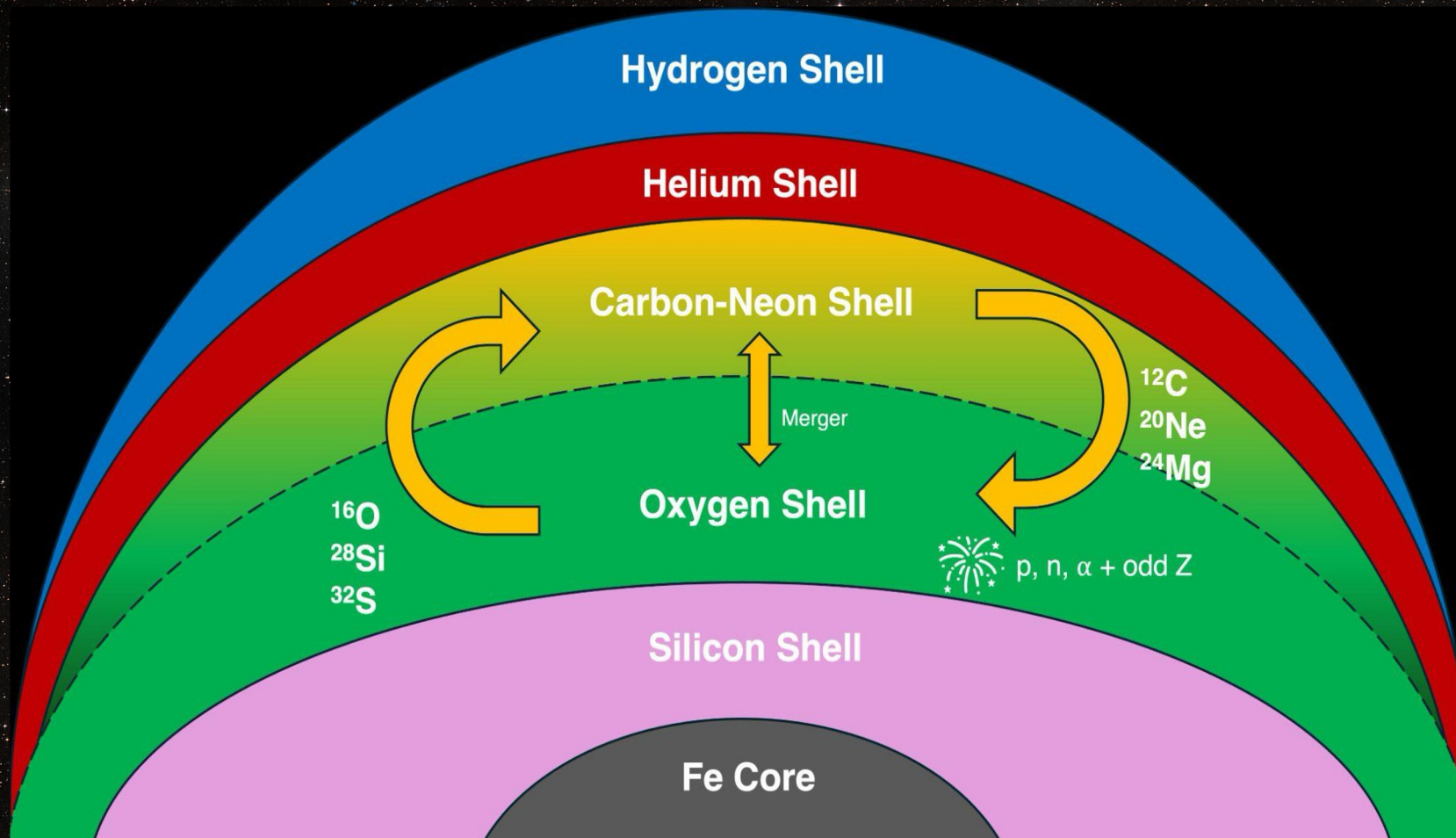
Rome + Berkeley group: M. Limongi, A. Chieffi, A. Falla, L. Boccioli

NuGrid collaboration





# Carbon-oxygen shell mergers in massive stars



- Ingestion of some C (and Ne) in the O burning shell during late stages of the evolution;
- Convective-reactive event  $\rightarrow$  Formation of an extended mixed convective zone;
- Peculiar **nucleosynthesis** and impact on the **explodability**;
- Found often in 1D stellar models with  $M_{\text{ini}} \leq 25 M_{\odot}$ ;
- Confirmed by 3D simulations (e.g., Rizzuti+24).



# Carbon-oxygen shell mergers in massive stars

SCAN ME









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**ACCEPTED!**

**DOI: 10.1051/0004-6361/202554461**

## The occurrence and impact of carbon-oxygen shell mergers in massive stars

L. Roberti<sup>1,2,3,4,5,\*</sup>, M. Pignatari<sup>2,3,6,5</sup>, H. E. Brinkman<sup>7,5</sup>, S. K. Jeena<sup>8</sup>, A. Sieverding<sup>9</sup>, A. Falla<sup>10,4</sup>,  
M. Limongi<sup>4,11,12</sup>, A. Chieffi<sup>13,14,12</sup>, and M. Lugaro<sup>2,3,15,14</sup>



# Carbon-oxygen shell mergers in massive stars

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# Nucleosynthesis in a C-O shell merger

- Enhanced O-burning products and production of odd-Z elements (P, Cl, K, Sc);
- Efficient photodisintegration of heavy elements: production of the weak s-process and  $\gamma$ -process nuclei.

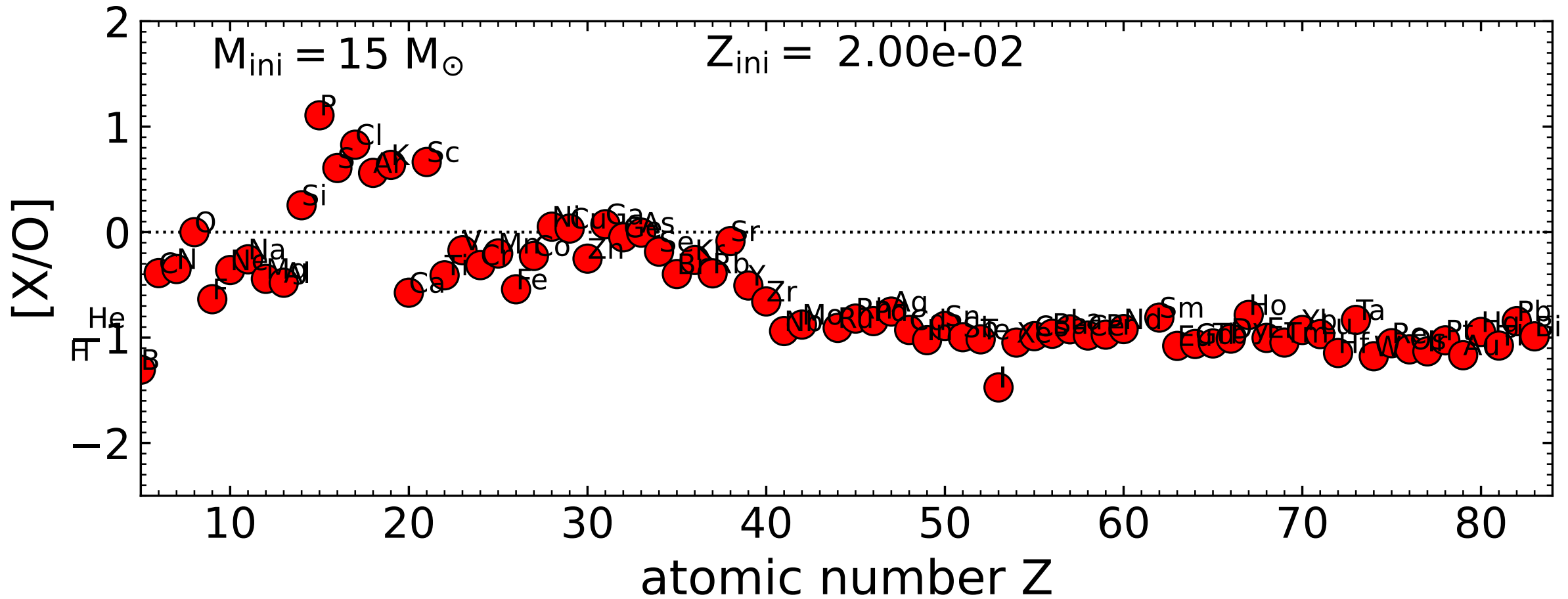


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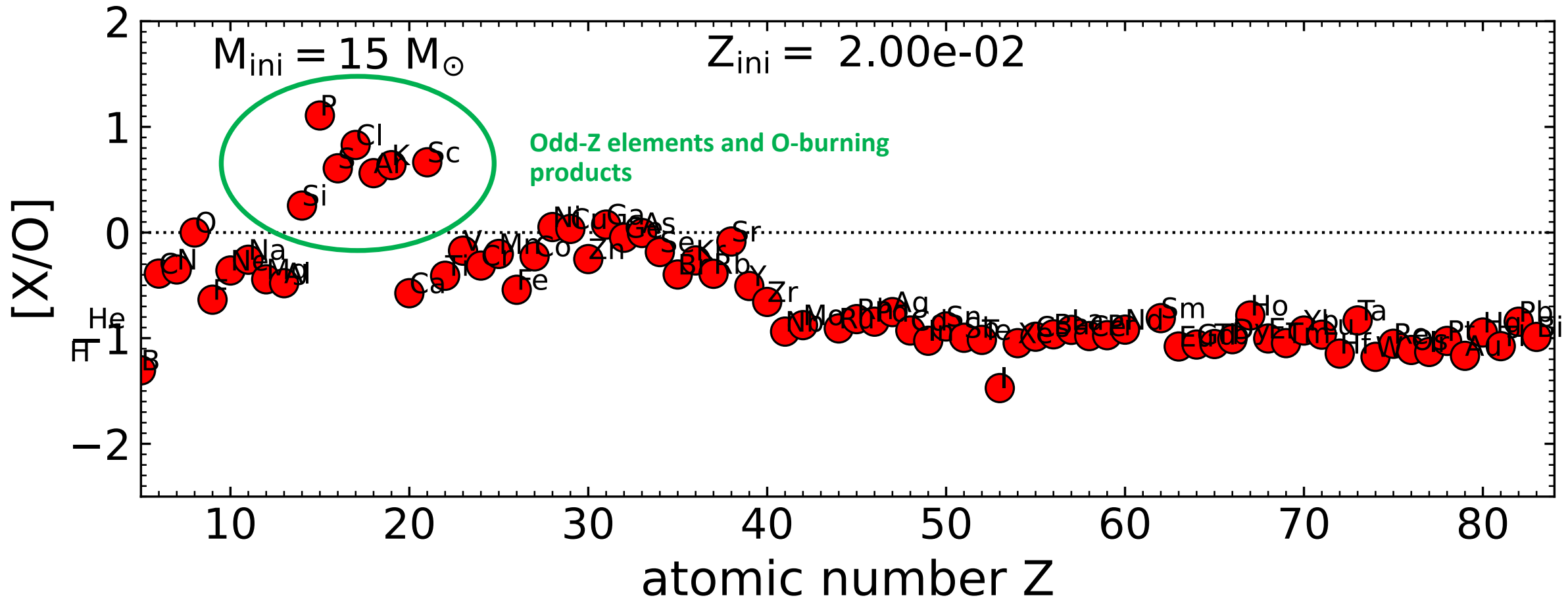


# Nucleosynthesis: odd-Z element production



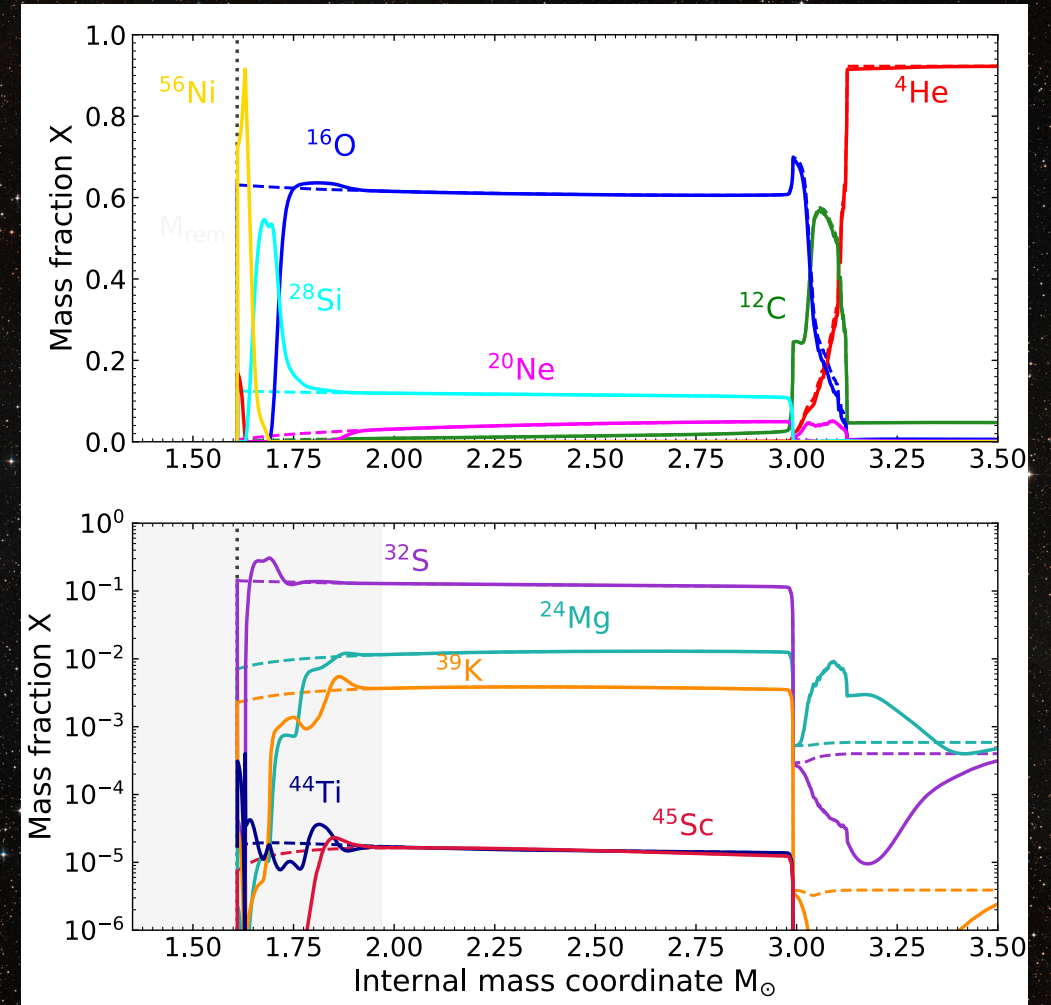
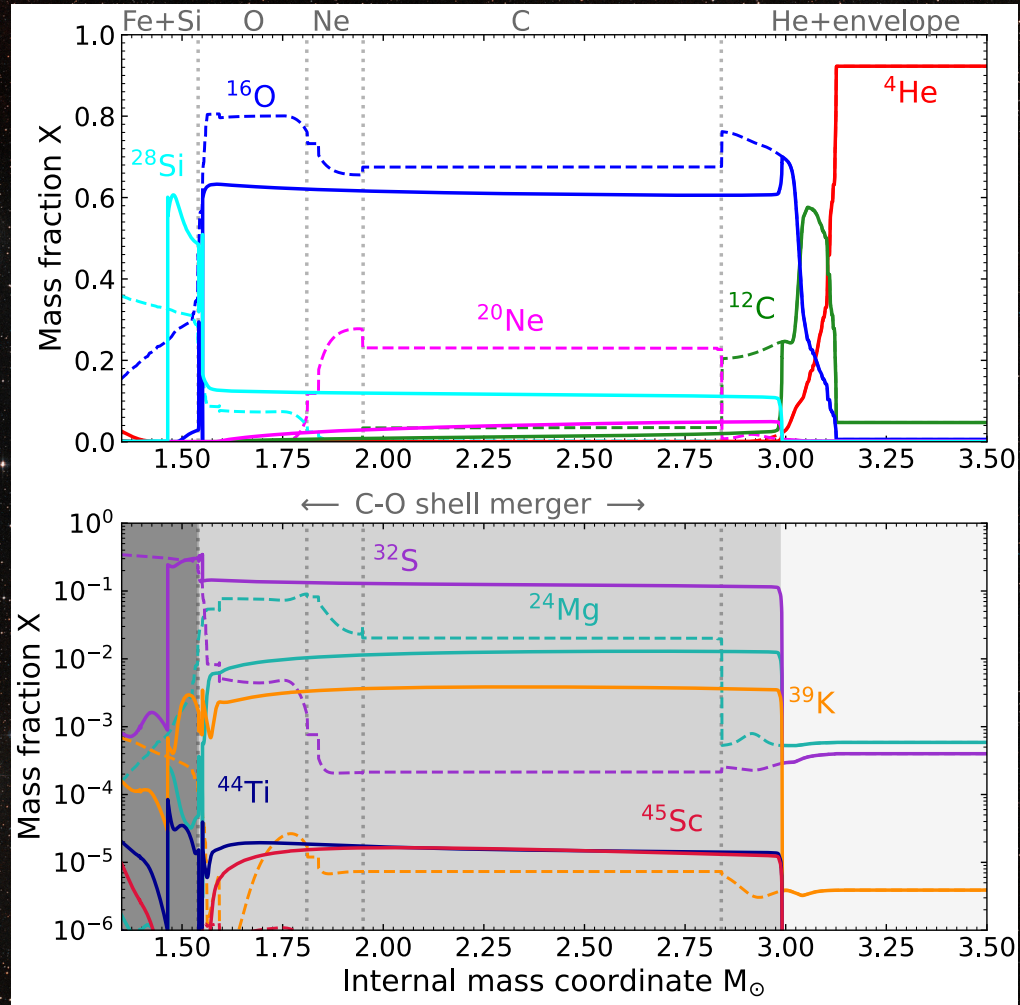


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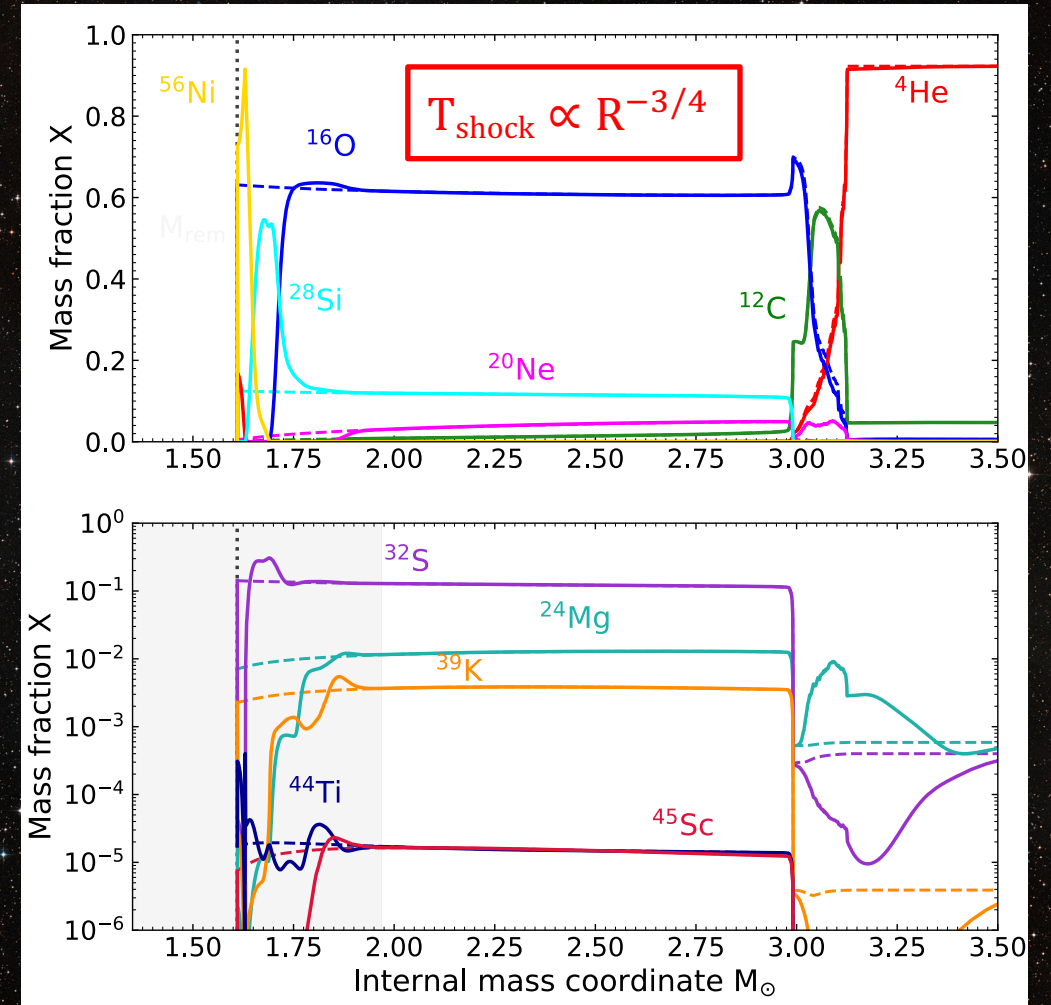
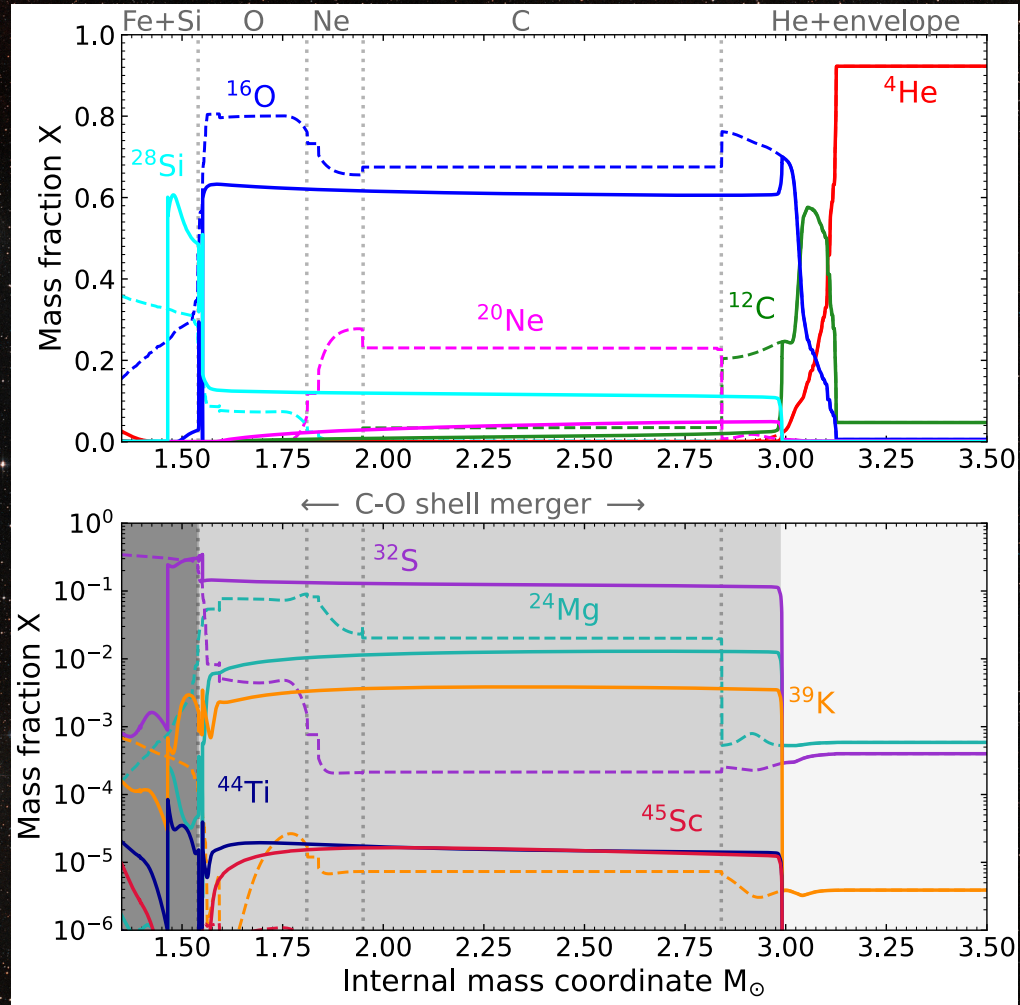


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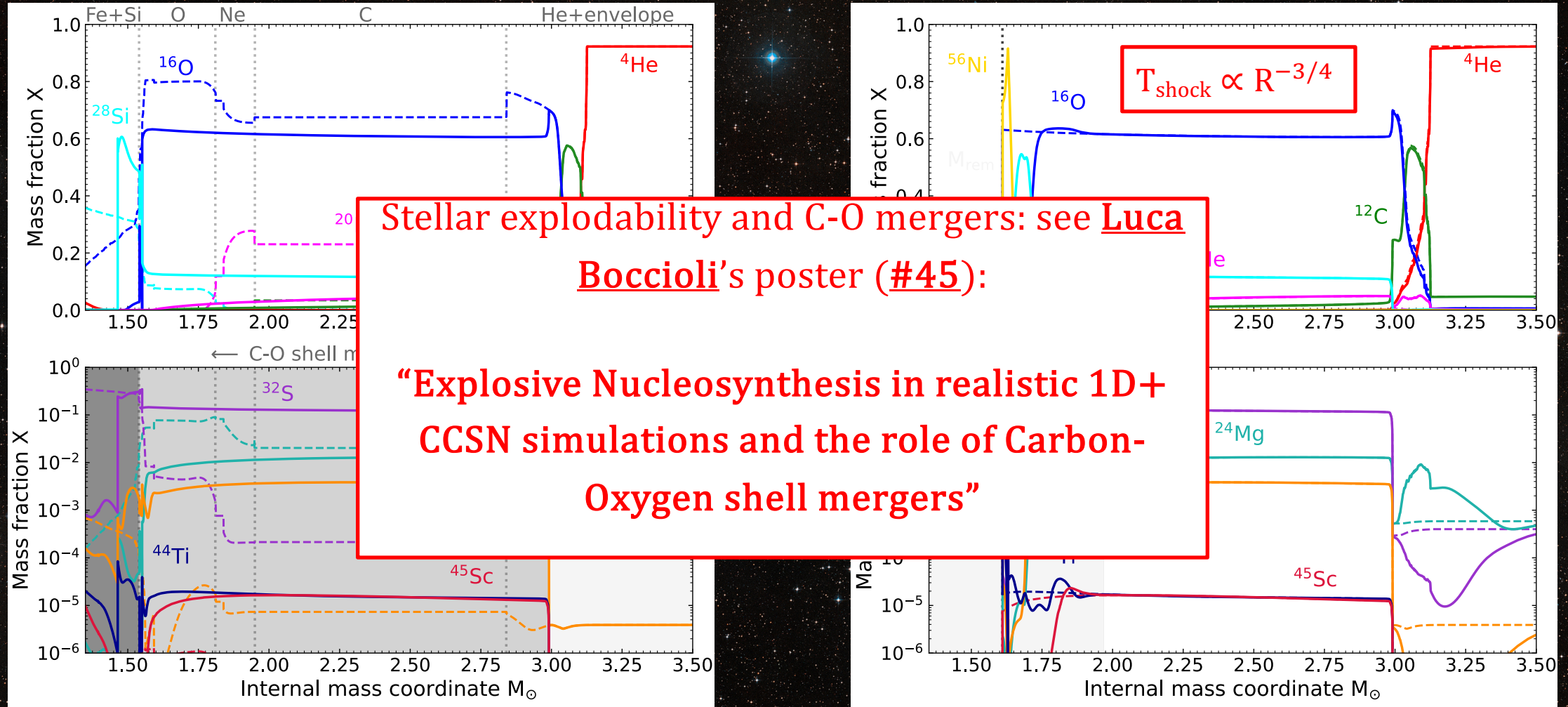


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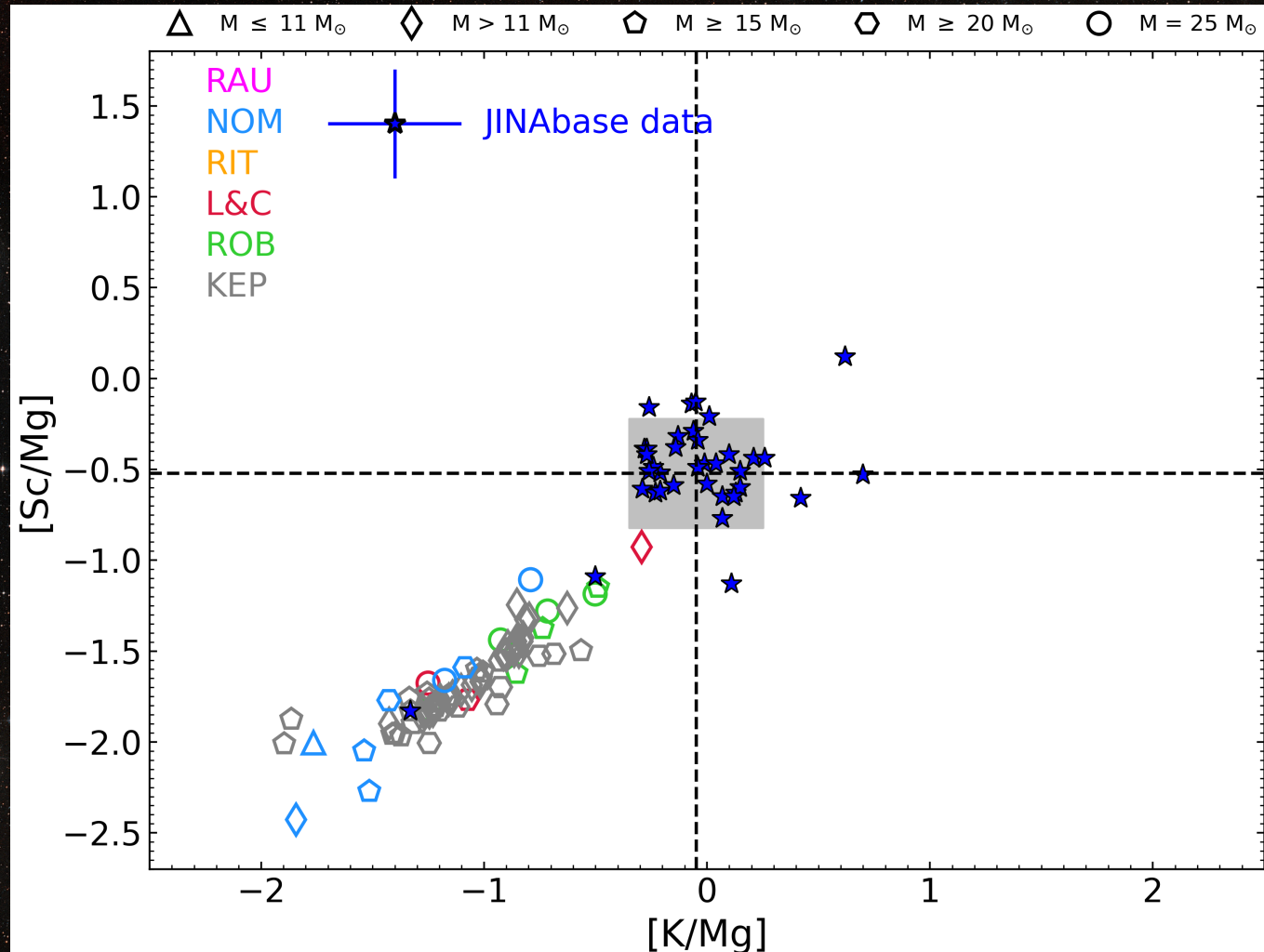


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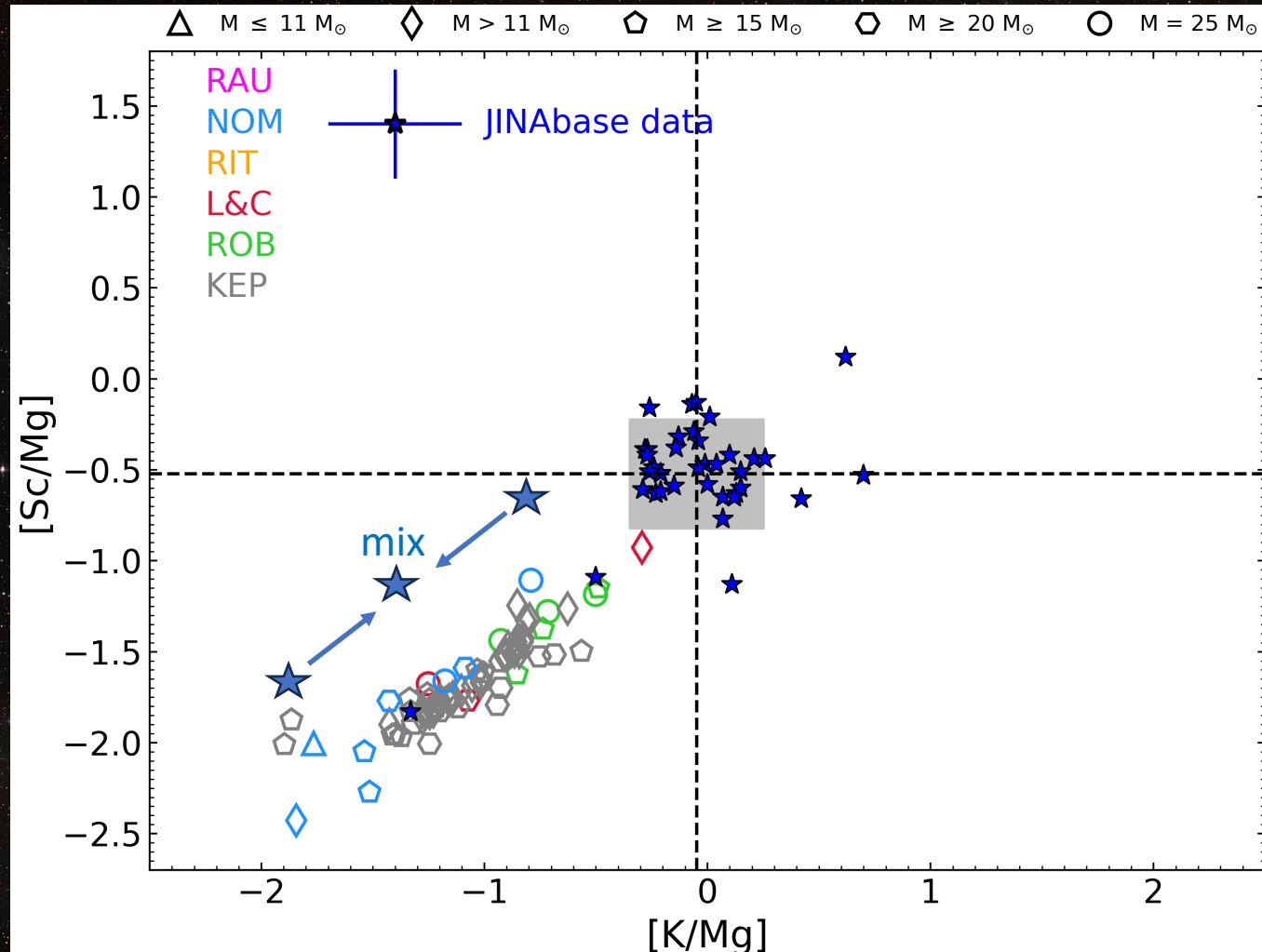
# Stellar Archaeology: odd-Z elements



- Comparisons between Pop III models and observations with  $[Fe/H] < -3 \rightarrow$  one or few enrichment episodes;



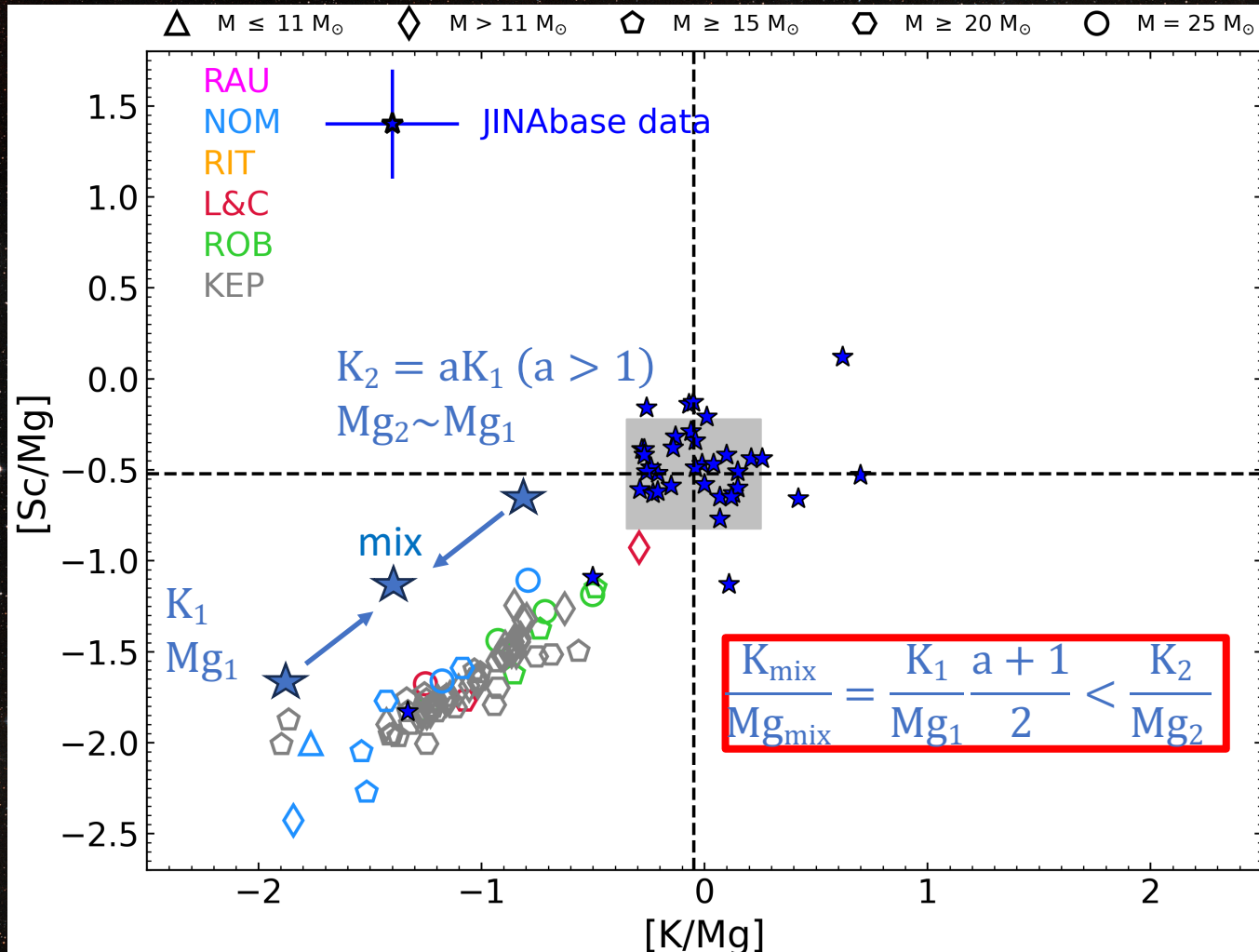
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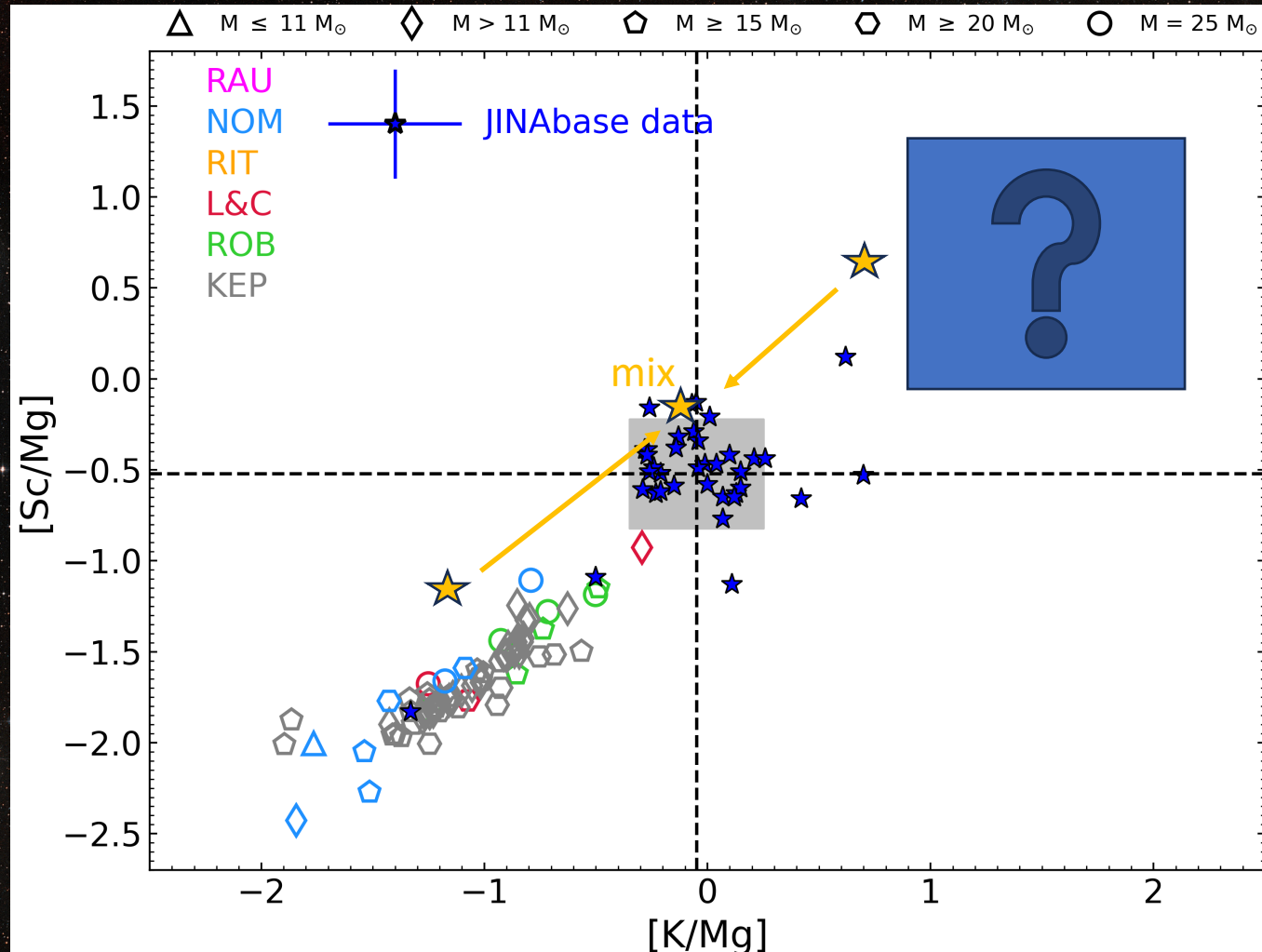
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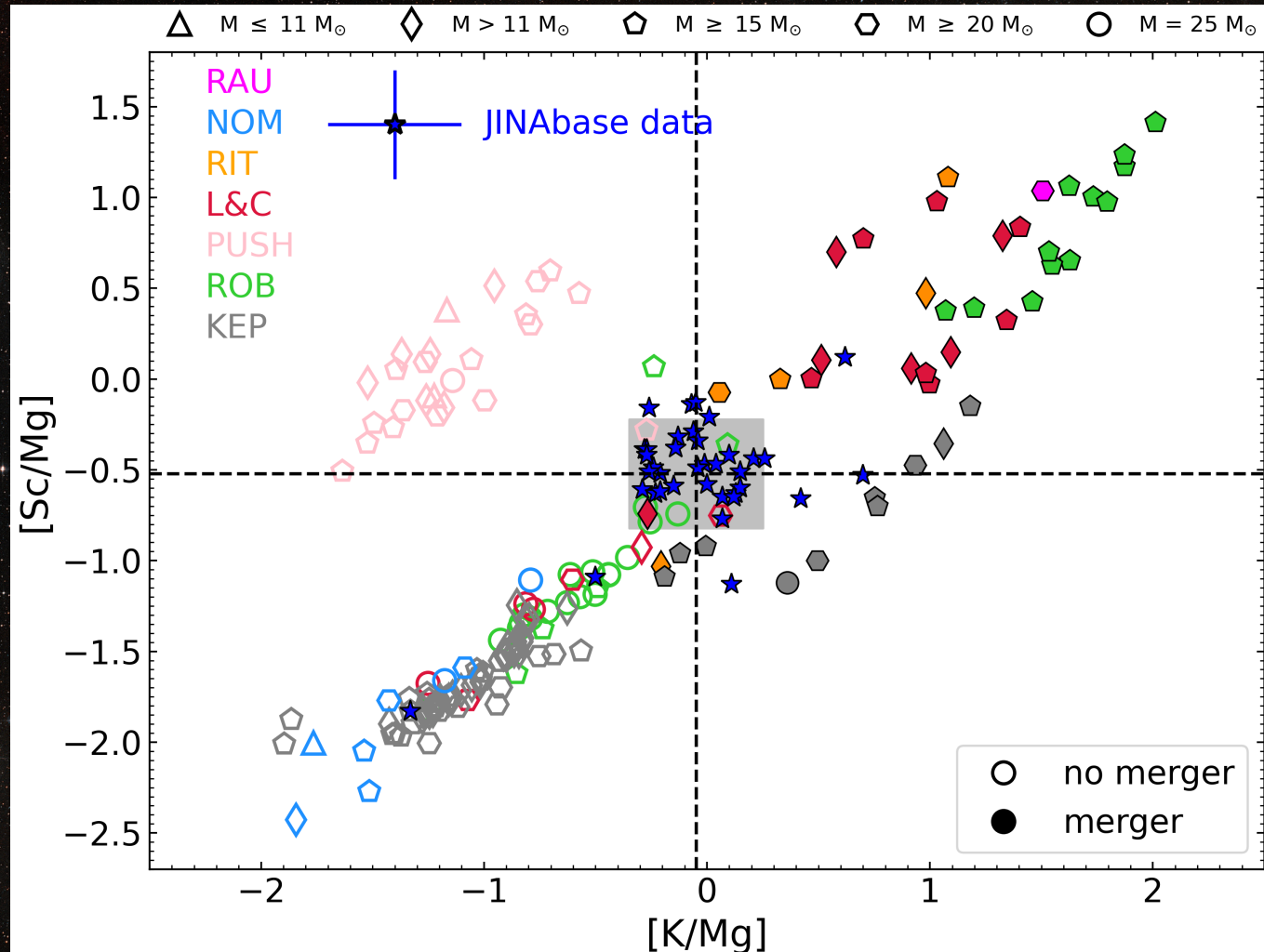
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- Mixing between more stars does not allow the ratios to move to the observations!
- An extra (super-solar) component is required;
- Models with C-O shell mergers naturally populate the required quadrant!



# Summary and conclusions

- Carbon-oxygen (C-O) shell mergers in massive stars have a crucial impact on the explosion and on the ejecta composition of CCSNe;
- Indications from 3D modeling seem to confirm their existence;
- Production site for odd-Z elements: possible enrichment already in the early universe!
- Large density jump at the Si/O interface because of the extended mixed convective region: easier explosion? In progress, stay tuned!
- Questions: frequency? Trend with mass/metallicity/rotation? Nuclear physics?

THANK YOU!