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Nuclear Astrophysics Activities at CENS

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Introduction of CENS(Center for Exotic Nuclear Studies)



- The Center for Exotic Nuclear Studies (CENS) was launched in December 2019.
- We have 28 researchers and 8 Ph.D. students as of today. The Nuclear Astrophysics group is lead by Tony Ahn.



- First approved experiment by RIKEN PAC in Dec.
 2020
- 14O(a,p) experiment at CRIB/CNS in Mar. 2023
- CENS organized NIC XVII in September 2023.
- First RAON experiments in Sep. 2024
- CENS co-organized INPC 2025 in May 2025.







Nucleosynthesis processes



Nuclear chart and the major nucleosynthetic processes in the universe X. Tang *et al.*, *Association of Asia Pacific Physical Societies* 31, 19 (2021)

Nucleosynthesis process can explain the observation.
 Nuclear Physics plays an important role!



Calculated r-process yields for solar abundance patterns F. Kappeler *et al. Rep. Prog. Phys.* 52 945 1989





Astrophysically important nuclei on CENS Nuclear Chart







CENS detector and device development







Development of STARK silicon detector array

- STARK: Silicon Telescope Array for Reactions in inverse Kinematics
- Three rings with 12-16-12 polygons. 96-116-107 mm from the center of the target
- Scattering chamber: 580(X) x 400(Y) x 600(Z) mm3
- CryoSTAR (low T. gas cell target) compatible
- \rightarrow (α ,p) reaction studies, transfer reaction studies, OMP studies





⁷Li(d,p)⁸Li reaction simulation with
 40 MeV ⁷Li beam energy











Conceptual Design of STARK chamber

Solid Targets: CH₂, CD₂

Gas Targets: H₂, D₂, ⁴He, ³He, N₂

beams

기초과학연구원

CENS detector and device development (selected)



Silicon Telescope Array for Reactions in inverse Kinematics

[VOICE]







Poster #136 by Soomi Cha Development of the Active Target TPC for Multiple nuclear physics eXperiments (AToM-X)

[ATOM-X] Active target TPC for Multiple nuclear eXperiments

[JETTSTAR]



gas JET Target system for nuclear STructure and Astrophysical Research







CENS *ap*-explorer Project



Rank	Reaction	Type ^a	Sensitivity ^b	Category	Rank	Reaction	Type ^a	Sensitivity ^b	Category
1	⁵⁶ Ni(α, p) ⁵⁹ Cu	U	12.5	1	1	$15\Omega(\alpha,\alpha)^{19}Ne$	D	16	1
2	${}^{59}Cu(p, \gamma){}^{60}Zn$	D	12.1	1	1	5611 () 590		10	1
3	$^{15}O(\alpha, \gamma)^{19}Ne$	D	7.9	1	2	$^{50}Ni(\alpha, p)^{50}Cu$	U	0.4	1
4	$^{30}S(\alpha, p)^{33}Cl$	U	7.8	1	3	$^{59}Cu(p, \gamma)^{60}Zn$	D	5.1	1
5	$^{26}\text{Si}(\alpha, p)^{29}\text{P}$	U	5.3	1	4	${}^{61}Ga(p, \gamma){}^{62}Ge$	D	3.7	1
6	${}^{61}\text{Ga}(\text{p}, \gamma){}^{62}\text{Ge}$	D	5.0	1	5	$^{22}Mg(\alpha, p)^{25}Al$	D	2.3	1
7	²³ Al(p, γ) ²⁴ Si	U	4.8	1	6	$140(0, p)^{17}F$	D	5.9	1
8	${}^{27}P(p, \gamma){}^{28}S$	D	4.4	1	7	$23_{A1(a, b)}^{24_{C}}$	D	5.0	1
9	⁶³ Ga(p, γ) ⁶⁴ Ge	D	3.8	1	/	23 Al(p, γ) 23 Si	D	4.0	1
10	60 Zn(α , p) 63 Ga	U	3.6	1	8	$^{16}Ne(\alpha, p)^{21}Na$	U	1.8	1
11	$^{22}Mg(\alpha, p)^{25}Al$	D	3.5	1	9	63 Ga(p, γ) 64 Ge	D	1.4	2
12	⁵⁶ Ni(p, γ) ⁵⁷ Cu	D	3.4	1	10	$^{19}F(p, \alpha)^{16}O$	U	1.3	2
13	$^{29}S(\alpha, p)^{32}Cl$	U	2.8	1	11	$^{12}C(\alpha, \gamma)^{16}O$	U	2.1	2
14	$^{28}S(\alpha, p)^{31}Cl$	U	2.7	1	12	26Si(a, p)29P	U	1.9	2
15	$^{31}Cl(p, \gamma)^{32}Ar$	U	2.7	1	12	$17r(\alpha, p)$ r	U	1.0	2
16	³⁵ K(p, γ) ³⁶ Ca	U	2.5	2	13	$r(\alpha, p)$ where $r(\alpha, p)$	U	3.5	2
17	$^{18}\text{Ne}(\alpha, p)^{21}\text{Na}$	D	2.3	2	14	$^{24}Mg(\alpha, \gamma)^{28}Si$	U	1.2	2
18	$^{25}\text{Si}(\alpha, p)^{28}\text{P}$	U	1.9	2	15	${}^{57}Cu(p, \gamma){}^{58}Zn$	D	1.3	2
19	⁵⁷ Cu(p, γ) ⁵⁸ Zn	D	1.7	2	16	60 Zn(α , p) 63 Ga	U	1.1	2
20	34 Ar(α , p) 37 K	U	1.6	3	17	$17E(p_{c})^{18}Ne$	U.	17	2
21	$^{24}\text{Si}(\alpha, p)^{27}\text{P}$	U	1.4	3	10	40c a(r a) 41m	D	1.7	2
22	$^{22}Mg(p, \gamma)^{23}Al$	D	1.1	3	18	$^{\circ}$ Sc(p, γ) $^{\circ}$ II	D	1.1	2
23	⁶⁵ As(p, γ) ⁶⁶ Se	U	1.0	3	19	⁴ °Cr(p, γ) ⁴⁹ Mn	D	1.2	2
24	$^{14}O(\alpha, p)^{17}F$	U	1.0	3					
25	40 Sc(p, γ) 41 Ti	D	0.9	3					
26	${}^{34}Ar(p, \gamma){}^{35}K$	D	0.8	3					
27	${}^{47}Mn(p, \gamma){}^{48}Fe$	D	0.8	3					
28	$^{39}Ca(p, \gamma)^{40}Sc$	D	0.8	3					

R. H. Cyburt et al. ApJ 830:55

• Key Research Question: (2016) Direct measurements of key (α , p) reaction cross sections which important for α p-process and pprocess.

• Methods:

- 1. Thick Target in Inverse Kinematics (TTIK) using TexAT_v2, AToM-X or VOICE
- 2. (α, p) Reaction in Inverse Kinematics using JENSA, CryoSTAR or JETTSTAR with STARK





Direct measurement of ${}^{14}O(\alpha,p){}^{17}F$ cross section at CRIB

R. H. Cyburt et al. 2016 Rank Reaction Type Sensitivity ¹⁵O(a,y)¹⁹Ne 1 D 16 ⁵⁶Ni(α.p)⁵⁹Cu U 6.4 2 ⁵⁹Cu(p,y)⁶⁰Zn 5.1 3 D 61Ga(p,y)62Ge 4 3.7 D $^{22}Ma(\alpha,p)^{25}AI$ 5 D 2.3 $^{14}O(\alpha,p)^{17}F$ 5.8 6 D 7 ²³Al(p,y)²⁴Si 4.6 D ¹⁸Ne(α,p)²¹Na 8 U 1.8 9 ⁶³Ga(p,γ)⁶⁴Ge D 1.4 ¹⁹F(p,a)¹⁶O 10 U 1.3

Reactions that impact the burst light curve in the multi-zone X-ray burst model



"A direct measurement of the ¹⁴O(α ,p)¹⁷F reaction with the Texas Active Target detector" approved by RIKEN PAC (2020)

CRIB

Beam time was very hard to get due to the Covid-19. We performed the experiment in Mar.

CNS Radio-Isotope Beam Separator

(p,n)

¹⁴O beam (RI)



Previous measured data and calculated total cross sections of ${}^{14}O(\alpha,p)$ reaction











Csl Det.

Si. Det.

Si. Det



Proposed by S. Ahn & Data analysis by C. Park

Data Analysis Results - Preliminary



- Cross sections of the (a,p0) & (a,p1) reactions
- 2 2.8 MeV Shows good agreement
- 1 1.5 MeV Higher cross sections
 - → contributions from new resonances or the restricted resolution at low energies
- 2.8 MeV -
 - No resonance features
 - \rightarrow limited E_{cm} resolution

Lower cross sections than AZURE calculation



Direct measurement of ${}^{34}Ar(\alpha, p){}^{37}K$ cross section at CRIB/CNS

Motivation: a key reaction for understanding the luminosity curve of the double peak and nucleosynthesis mechanism in X ray bursts. ¹⁸Ne(α,p)²¹Na(p,γ)²²Mg(α,p)²⁵Al(p,γ)²⁶Si,(α,p)²⁹P(p,γ)³⁰S(α,p)³³Cl(p,γ)³⁴Ar(α,p)³¹K(p,γ)³⁸Ca(α,p)⁴¹Sc



Bird's eye-view of RAON

• Accelerator System • RI production System Conventional Utilities • Experimental System Assembly Bd. Waste Storage SRF Test Bd. **Guest House High Energy B** IF Accelerator **HQ Office** 100 **Control Center** ISOL SCL₂ **High Energy** an and SCL₃ Low Energy A Low Energy B Cryogenics Bd. Utility Bd. Electricity Bd. Courtesy of IRIS, (from T. Shin)





RAON



The RAON at the IRIS has a potential for the unique feature of combining the ISOL and In-Flight for the production of very rare isotopes.



KoBRA : Korea Board acceptance Recoil spectrometer & Apparatus

IF : In-flight Fragmentation separator **SCL 3**: (Low-energy) Super Conduction Linac 3 **SCL2** : (High-energy) SCL2





ISOL : Isotope-Separator On-Line

Experimental Setup of ⁴⁰**Ar+p elastic scattering**



Analysis results of the ⁴⁰Ar+p scattering data





Extracted new optical model parameters using SFRESCO.

8.96

1.29

0.540

52.0

1.19

0.672

D. Kim, S. Ahn et al., submitted for publication





5.75

0.590

0.996

Summary

- The objectives of CENS are to conduct theoretical studies and perform experiments using RI and stable beams at accelerator facilities around the world including RAON in Korea. Our nuclear astrophysics group investigates key nuclear reactions relevant to various stellar environments.
- We have developed major devices to be used at RAON and other RI accelerator facilities. ATOM-X, STARK, CENS silicon detectors, CryoSTAR, JETTSTAR, Clover HPGe detectors, Wien Filter, etc
- We recently focus on (a,p) reactions related to the rp-process Direct cross section measurements of ¹⁴O(a,p)¹⁷F and ³⁴Ar(a,p)³⁷K reactions using TexAT_v2 and AToM-X at CRIB/CNS.
- We performed commissioning experiments at RAON last year.
 Optical Model Potential study of ⁴⁰Ar+p elastic scattering at low energy region.





Acknowledgements

All the CENS members







Thank you for your attention!

Optical Model Potential Study with ⁴⁰Ar + p elastic scattering

Theoretical predictions for the cross-section in KD and DF models at low energies ($E_{lab} = 2.80, 4.95, \text{ and } 6.70 \text{ AMeV}$)



[Main Goal] Compare the phenomenological and microscopic optical model for ${}^{40}\text{Ar}+p$ elastic scattering in lower energy region.





Experimental Setup

CRIB CNS Radio-Isotope Beam Separator

F0 target: H2 gas (90K, 80 mm, ~500 Torr)

- TexAT_v2 (Texat Active Target TPC version 2)



 \rightarrow 3 MeV/u, 10⁵ pps after thin scintillator

Reaction Rate



• 0.3 - 2 GK

 Higher than REACLIB and TALYS
 ∵ Higher cross sections in the low energy region (Gamow window)

AToM-X: CENS Active Target TPC

- AToM-X : Active target TPC for Multiple nuclear eXperiment
- Active area: 256(X) x 180(Y) x 288(Z) mm³
- Scattering chamber: similar to TexAT chamber to make portable!
- Octagonal shape field cage, larger active area than TexAT Micromegas
- New micromegas with a good position resolution (0.5mm expected wt 4x4 mm² pixels)
- 5658 channels total (4608 from Micromegas plate and 1050 from aux detectors)
- GET electronics will be applied for signal processing: new AsAd board production.
- Online Data Analysis Cluster System (48 CPUs, 96GB Memory and 11 TB SSDs)







Top View of Micromegas





copied from S. Cha's KPS Fall 2023

Courtesy of Tony Ahn



Development of CENS silicon sensors

- Silicon detectors: good efficiency, angular resolution and energy resolution
- Issues:
 - \checkmark Most of them are made by Micron \rightarrow expensive and difficult to customize
 - Few experts who can make them in the field of low energy experiments. \checkmark
- Single-sided strip silicon sensors are developed at CENS (H.Y. Lee):
 - ✓ Pad design → Fabrication → Pad dicing → PCB design → PCB manufacturing → Assembly → Test



A fabricated 6-in silicon wafer



8strips sensor (BB10 design rule)



Wire bonding









Optical Model Potential Study of ⁴⁰Ar + p elastic scattering

- Optical model potential (OMP) parameters are required to predict cross-section for each energy.
- Lack of optical model parameters at low energies, especially near the Coulomb barrier.



[Main Goal] Compare the global optical models with the experimental data in low energy region and extract OMP parameters.

Copied from D. Kim's INPC 2025 presentation

Cryogenic Stable TARget (CryoSTAR)

- Goals: 40 K, 0.3~2 atm of $H_2/D_2/^3$ He/⁴He gas
- Reaction Gas Cell Dimensions: 3/6 mm thick, 20 mmD with Havar/Mo/Ti/Mylar window (2~50 um)









- ✓ First performance test completed with He gas.
- ✓ 40 K achieved with 1.3 bar Helium gas.
- ✓ To Do:
 - Various target test
 - target thickness measurement with alpha source
 - Performance test with light beams









Neutron transfer reactions

Proposed by S. Ahn

- The (n,γ) total cross section has a large uncertainty.
- The (n,γ) cross sections via discrete states or resonances.
- The (d,p) can provide direct capture part of the (n,y) cross section.
 - \rightarrow We need spin/parity, excitation energy and spectroscopic factor.







Sunghoon(Tony) Ahn