

Gravitational waves and Nucleosynthesis in Binary Neutron-stars simulations

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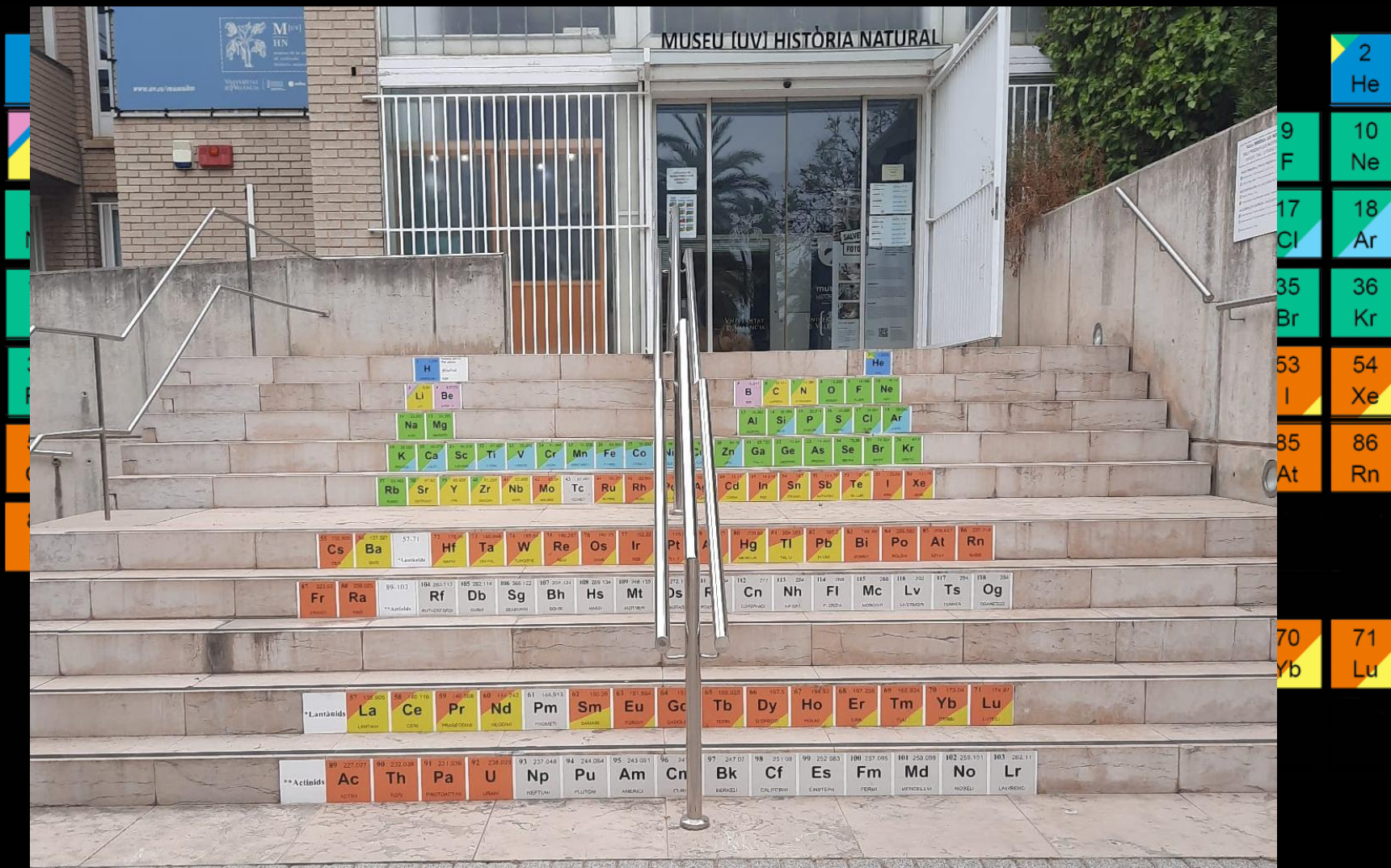
WHY NEUTRON STARS MERGER?

1 H	big bang fusion 					cosmic ray fission 					2 He						
3 Li	4 Be	merging neutron stars 					exploding massive stars 					5 B	6 C	7 N	8 O	9 F	10 Ne
11 Na	12 Mg	dying low mass stars 					exploding white dwarfs 					13 Al	14 Si	15 P	16 S	17 Cl	18 Ar
19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr
37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe
55 Cs	56 Ba		72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn
87 Fr	88 Ra																
		57 La	58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu	
		89 Ac	90 Th	91 Pa	92 U												

GRAPHIC BY JENNIFER JOHNSON

IMAGE CREDITS: ESA/NASA/ASSNOVA

WHY NEUTRON STARS MERGER?

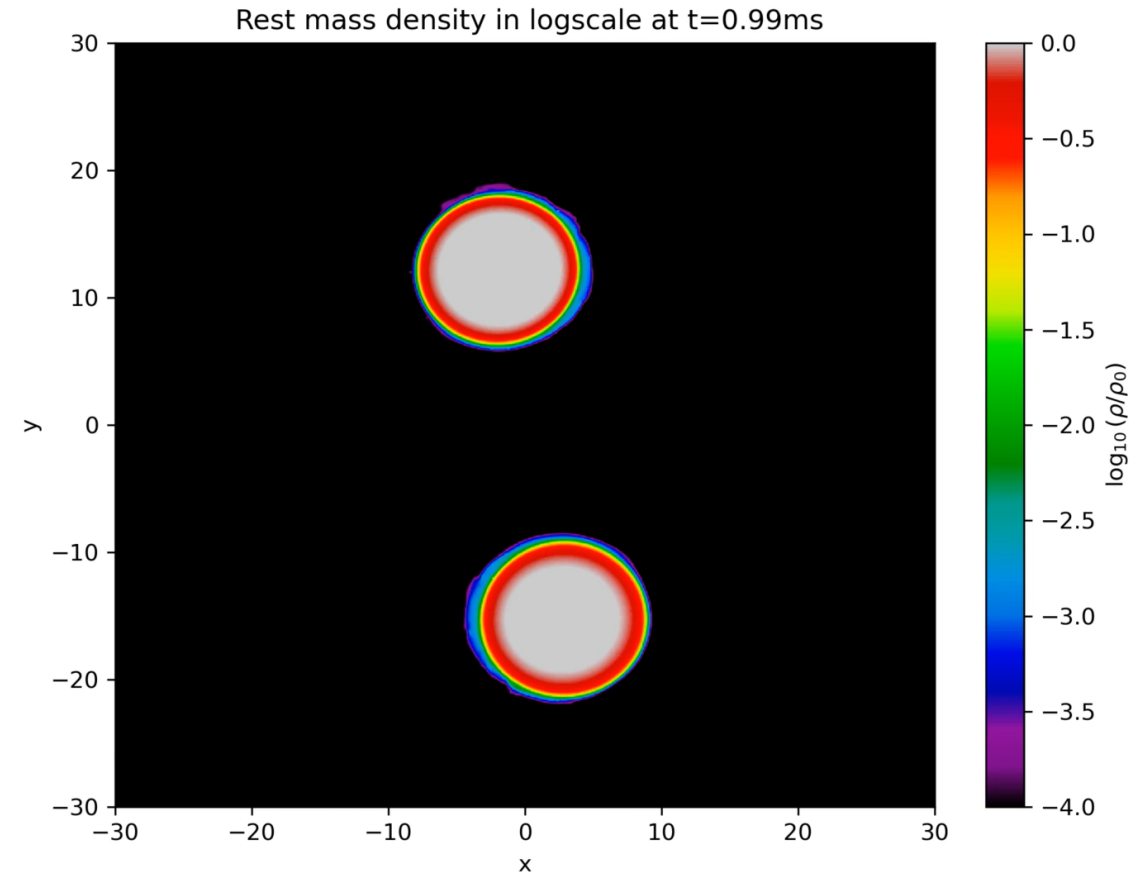


WHY NEUTRON STARS MERGER?



MERGER PHASES

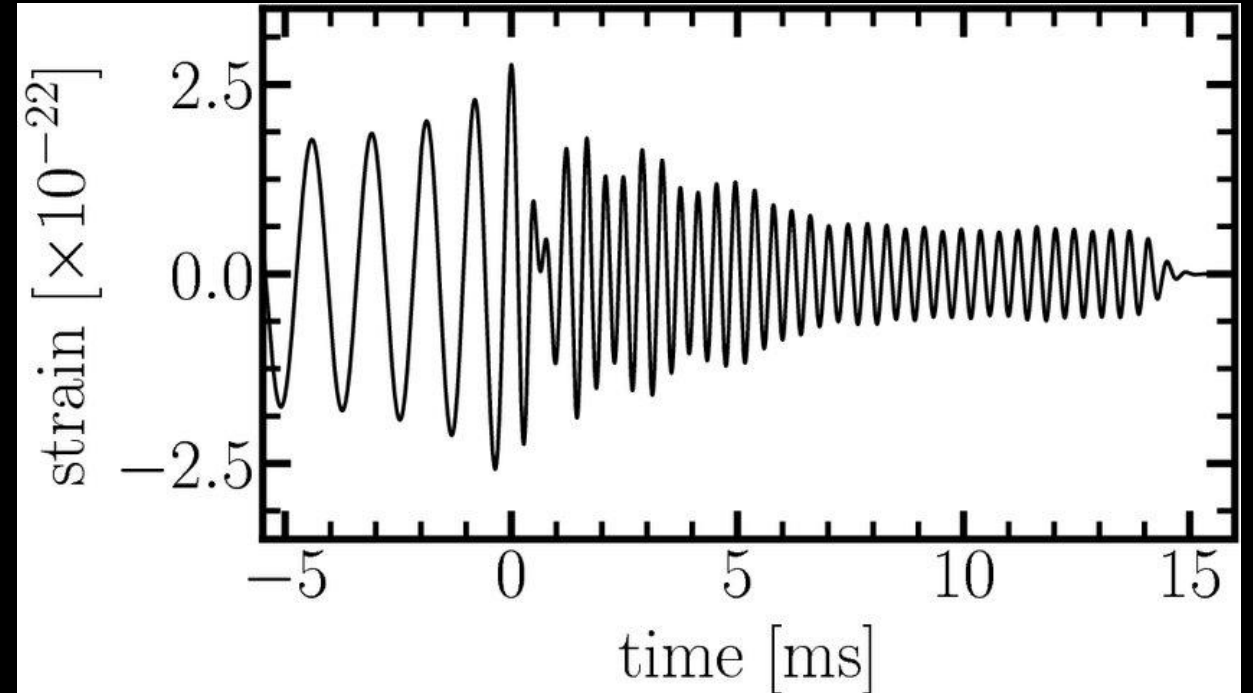
- **Inspiral**
- **Merger**
- **Post merger phase/Collapse**
- **Stable NS/Collapse**



Venturi, GR et al. (in prep)

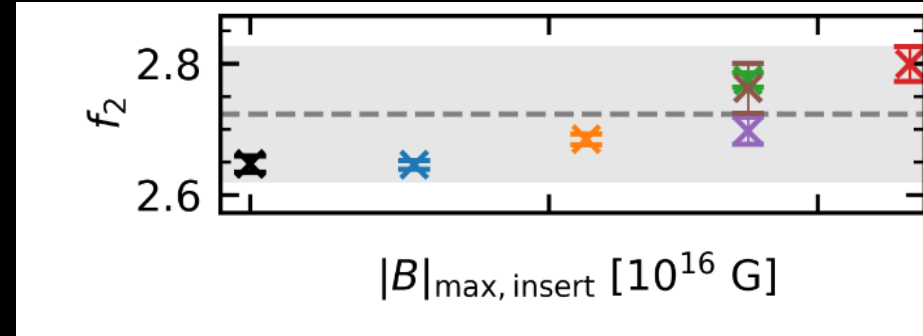
MERGER PHASES

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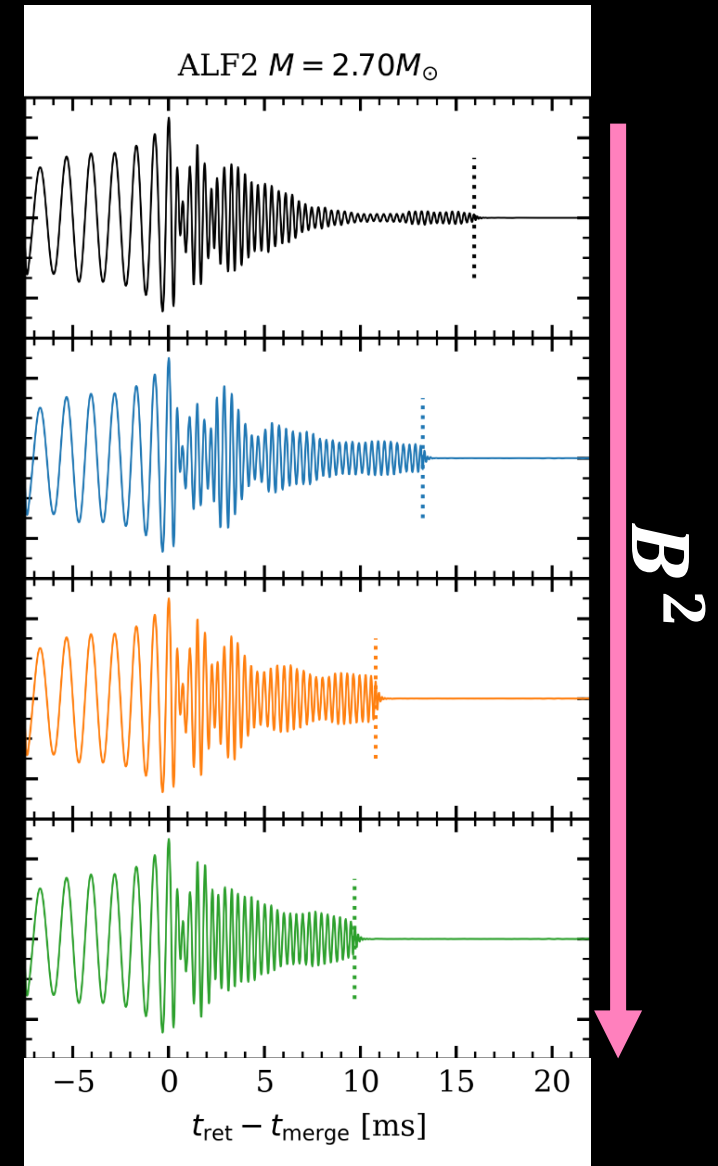
B-FIELD IN BNS



- Faster collapse to black hole
- Frequency shift in GWs

Other references about B-field and nucleosynthesis:

- De Haas et al. (2025), 10.1093/mnras/stad2931
- Kelsey A. Lund et al 2024 *ApJ* 964 111



BAMBER ET AL. (2025), PhysRevD.111.044038

OUR SETUP

- EoSs:

- SFHo (soft) $q = 1,25$
- DD2 (intermediate) $q = 1,25$
- HShen (stiff) $q = 1$

Toroidal



Poloidal



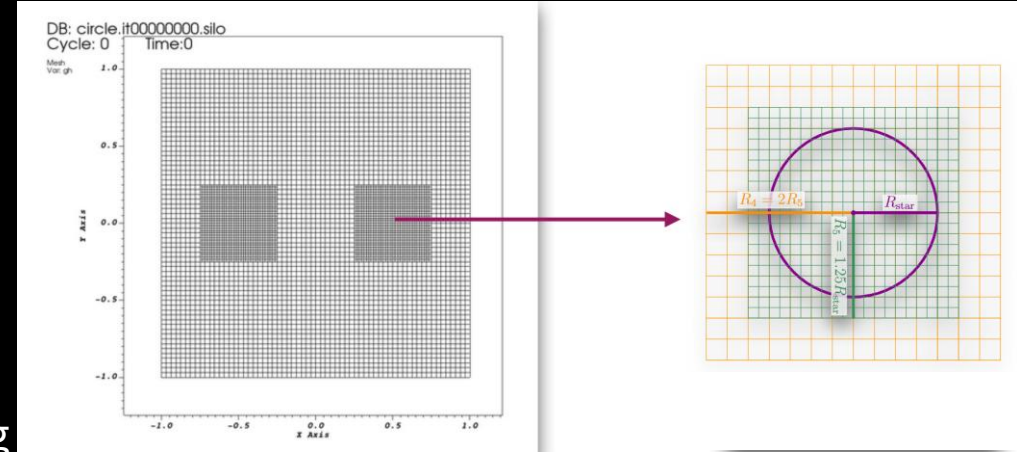
Venturi, GR et al. (in prep)

OUR SETUP

- BNS simulation using:
Einstein Toolkit + IllinoisGRMHD
adding tracer particles evolution
(up to 40 ms after merger)



www.einsteintoolkit.org
www.illinoisgrmhd.net

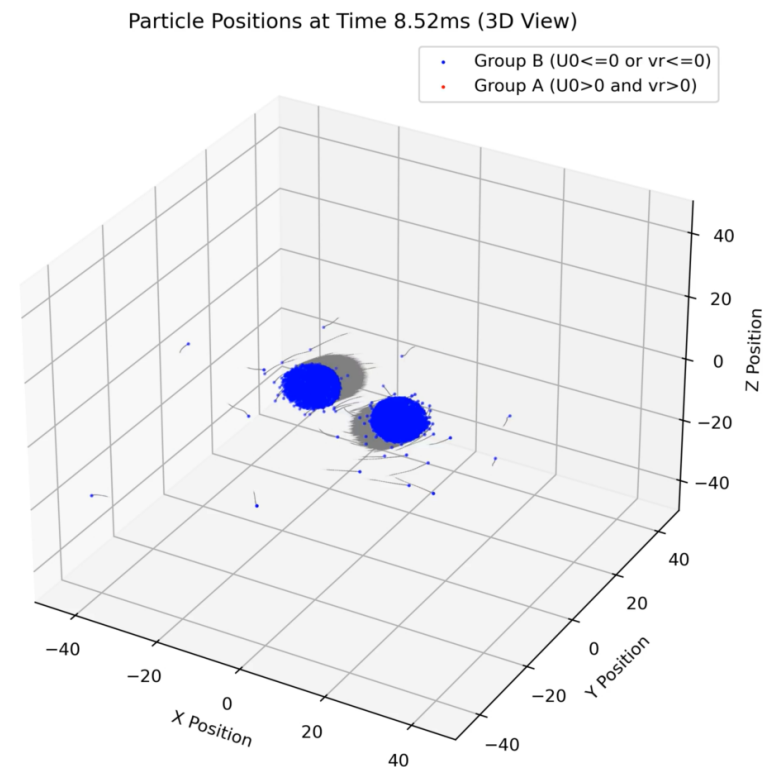
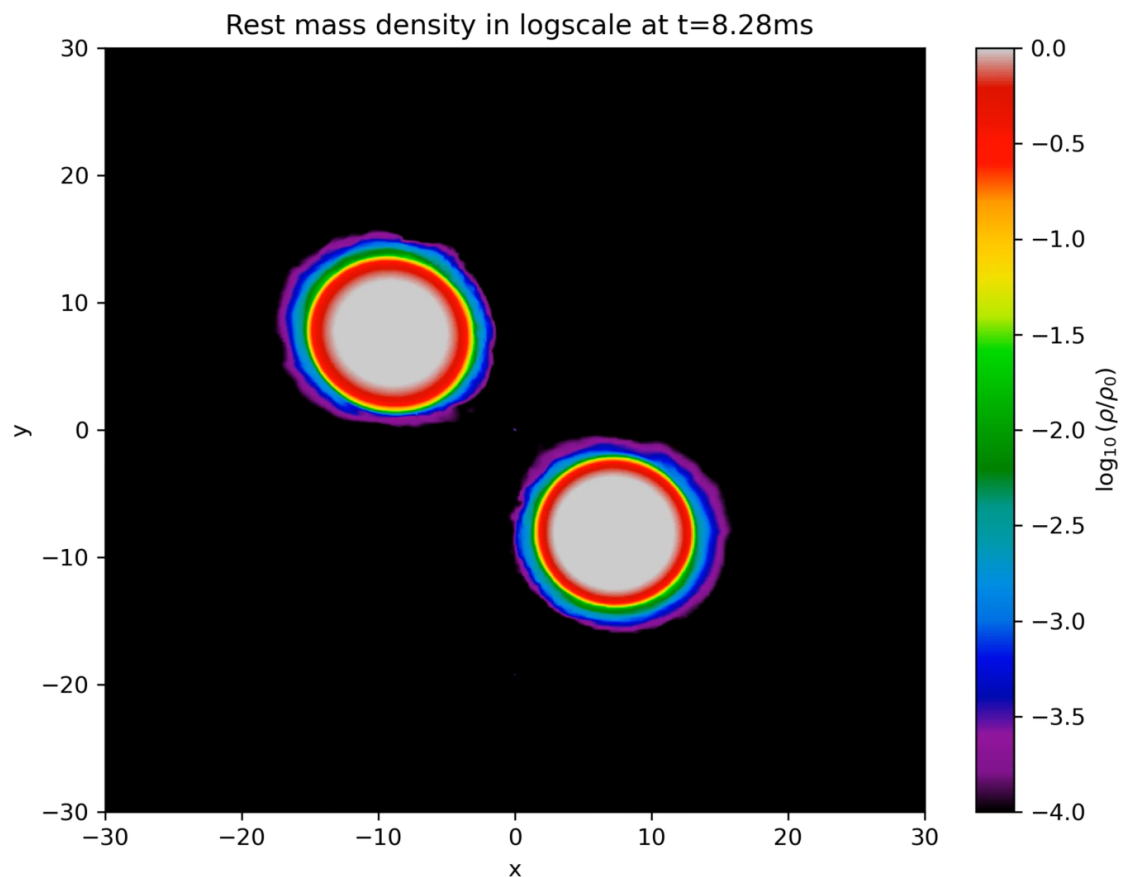


- Nucleosynthesis using WinNet
Of the ejected tracers



nuc-astro.github.io/WinNet/

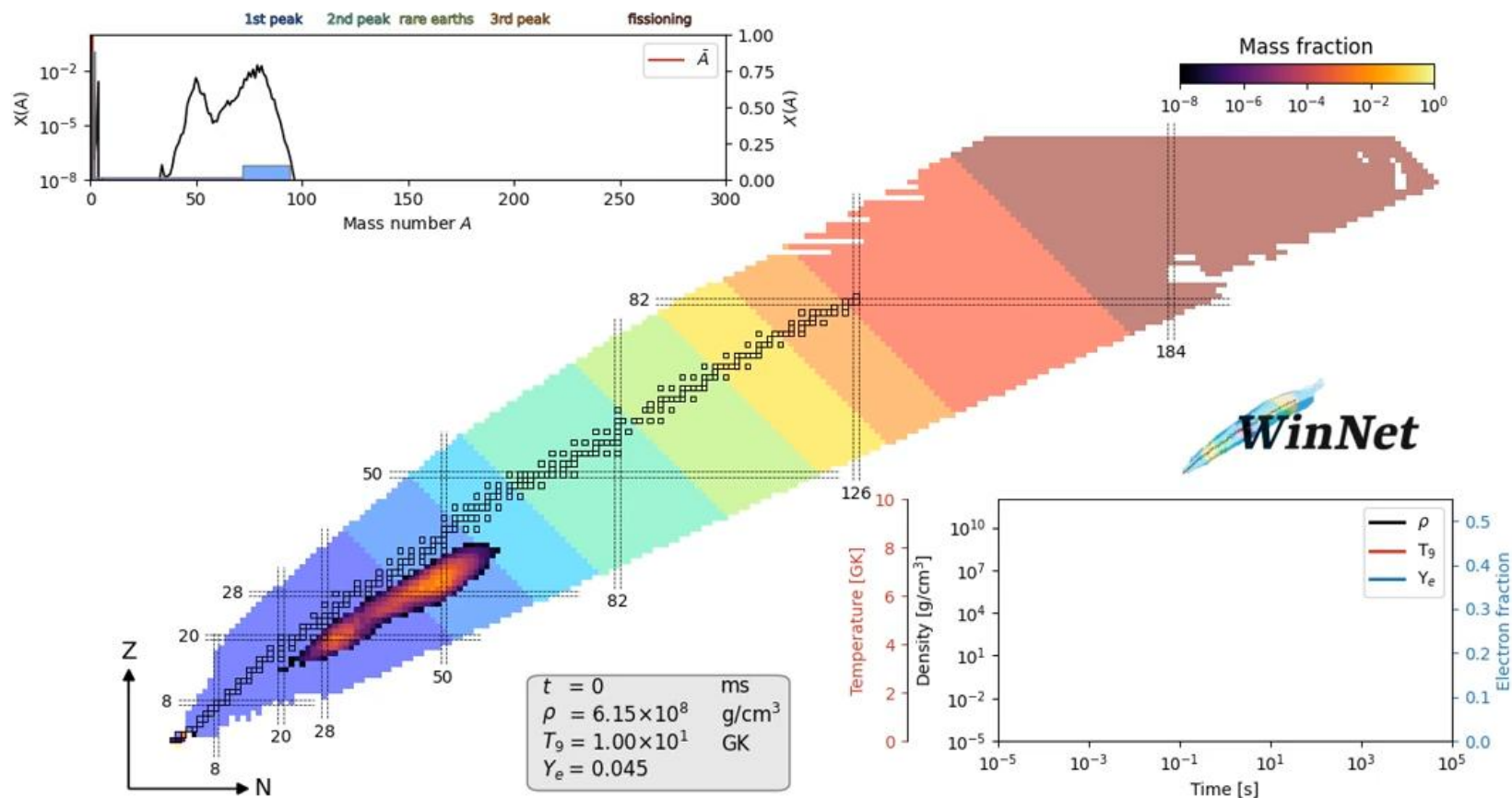
OUR SETUP



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OUR SETUP

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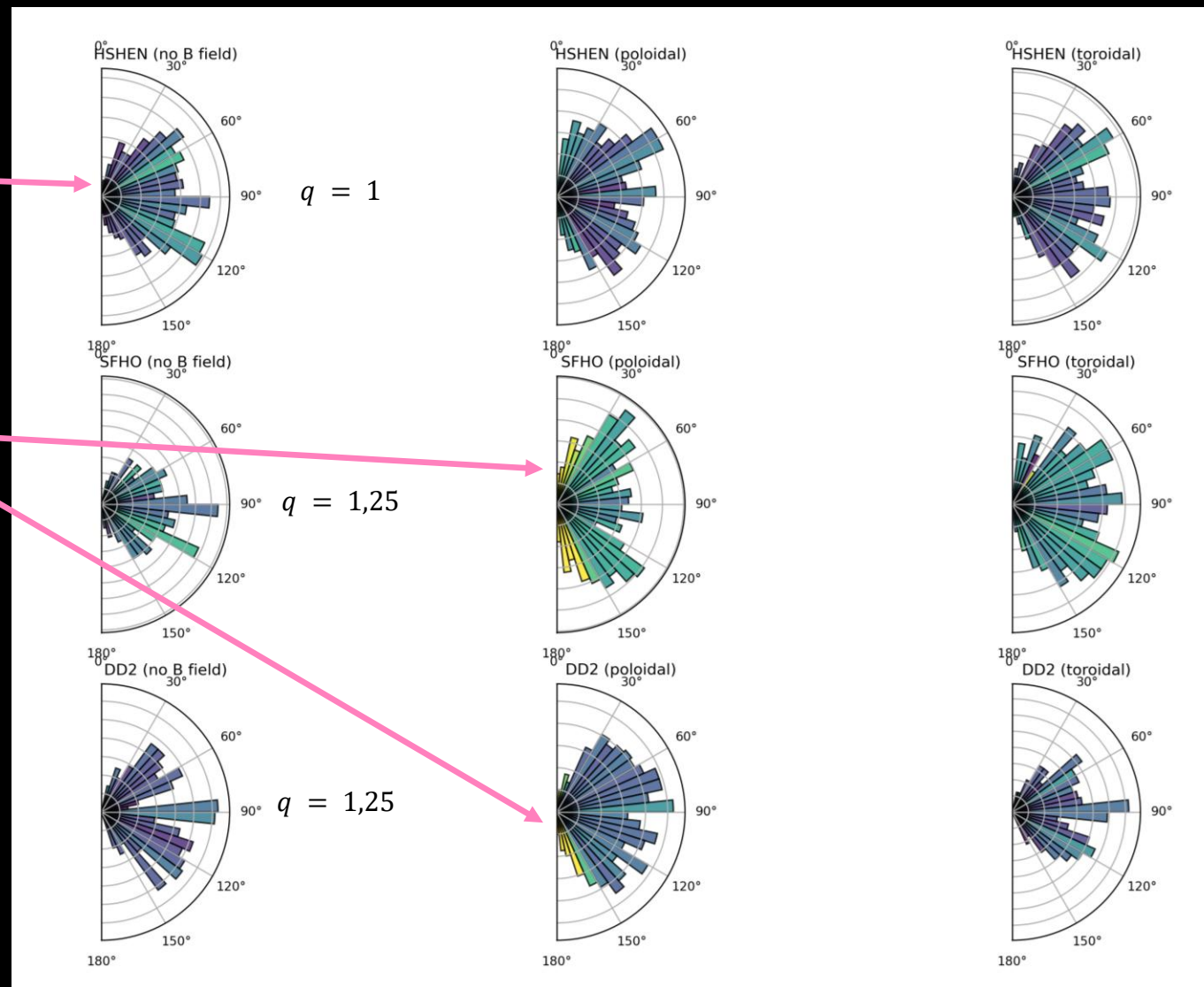
EJECTA PROPERTIES

Venturi, GR et al. (in prep)

- $q = 1$ seems not affected!

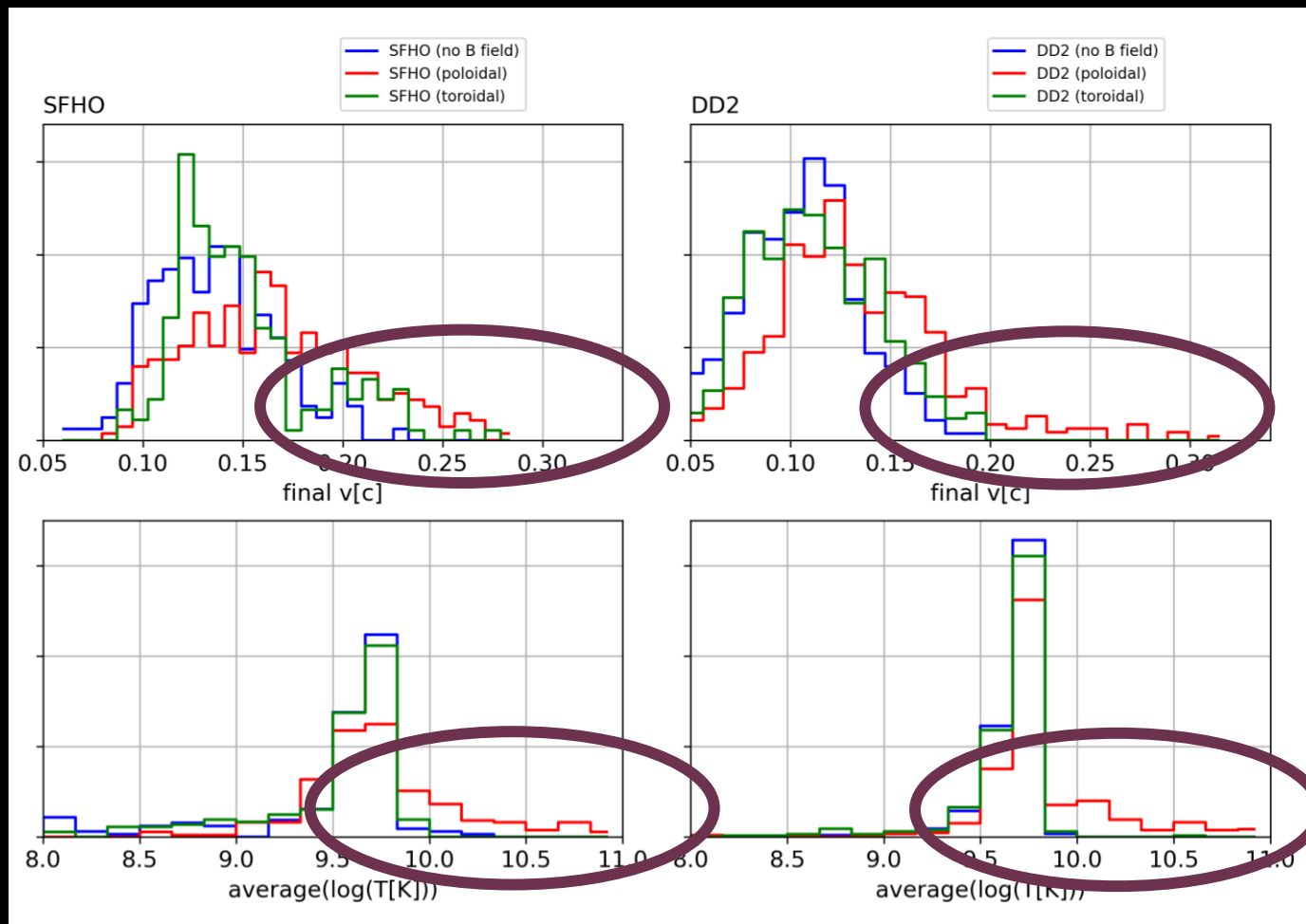
- $q = 1,25$ Poloidal shows fast ejection at the poles!

- $q = 1,25$ Toroidal is not affected!



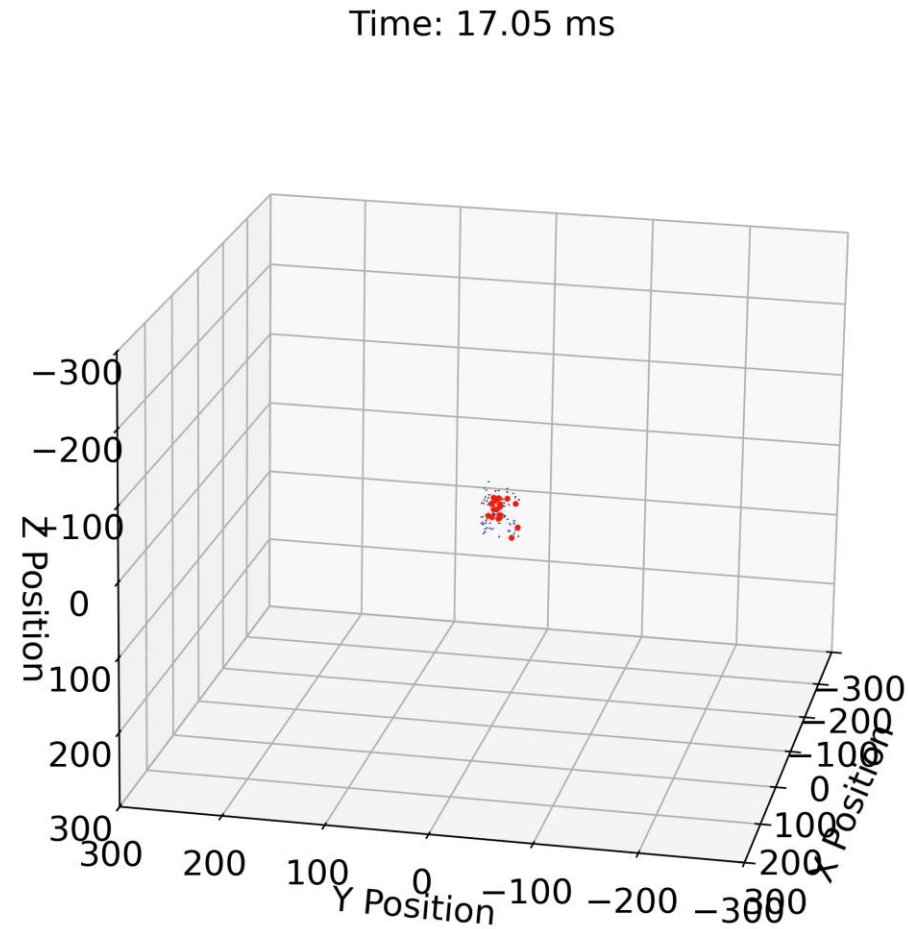
EJECTA PROPERTIES

Higher Velocities and Temperature for Poloidal and Unequal Masses



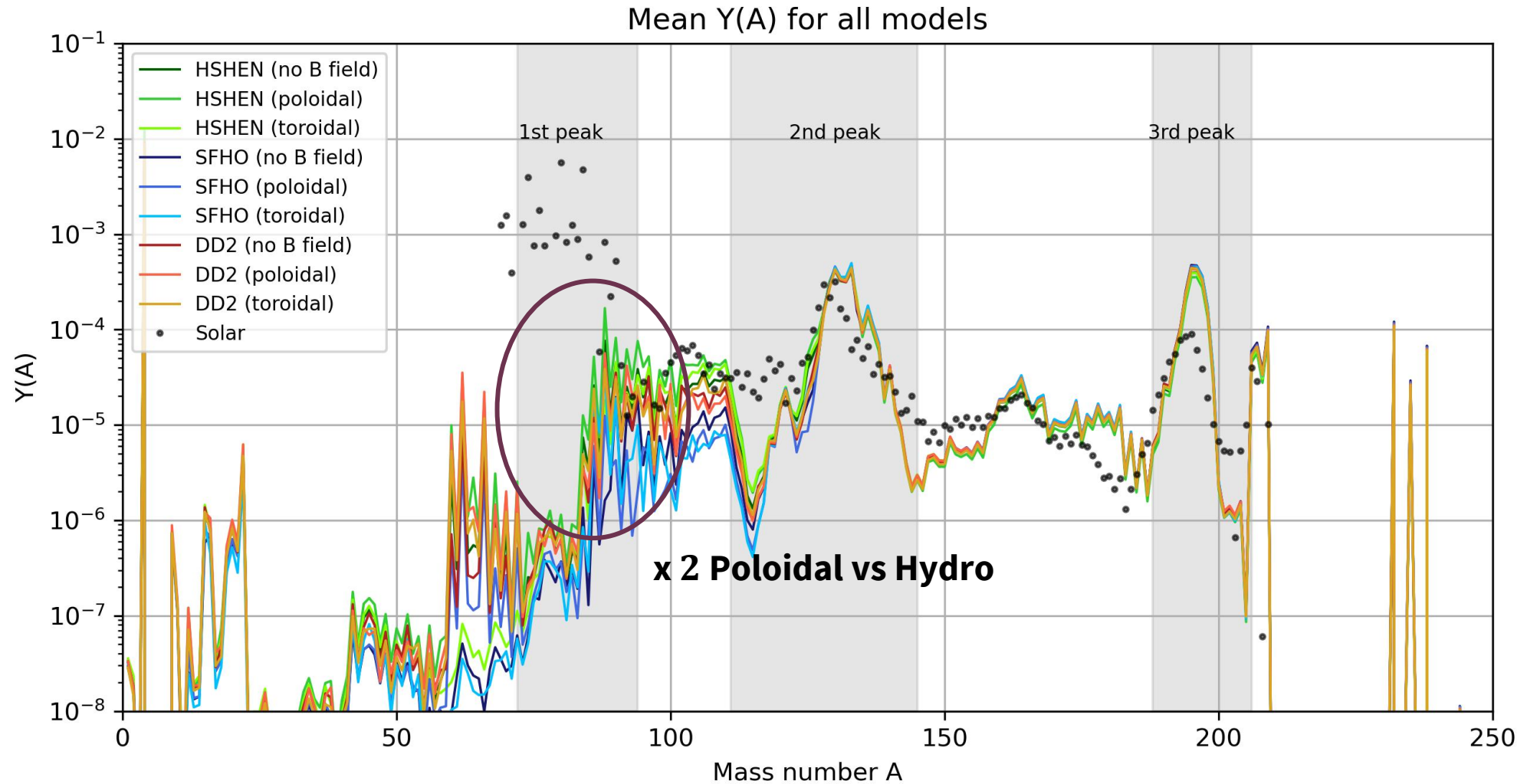
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EJECTA PROPERTIES: JET?

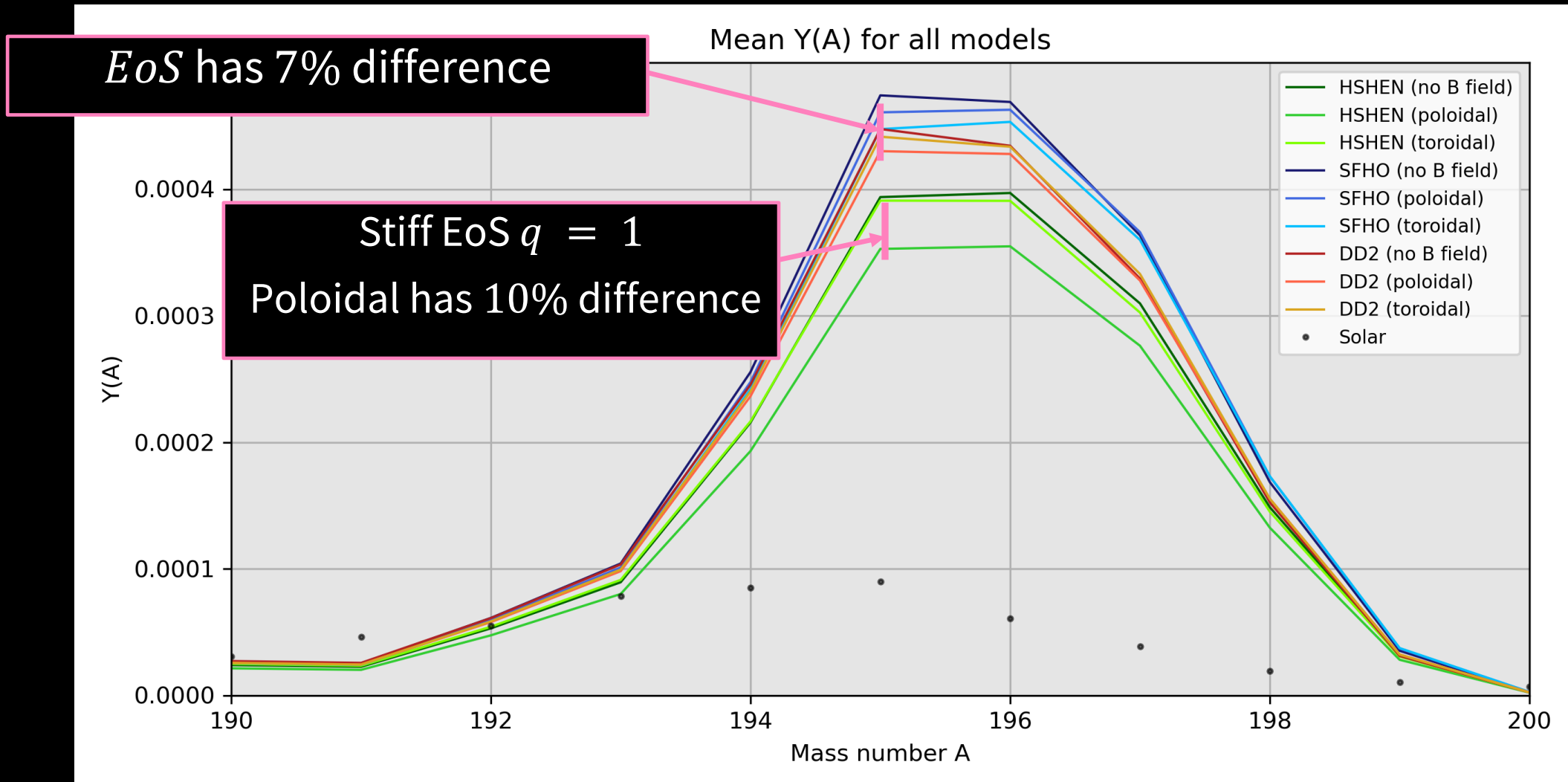


Venturi, GR et al. (in prep)

NUCLEOSYNTHESIS

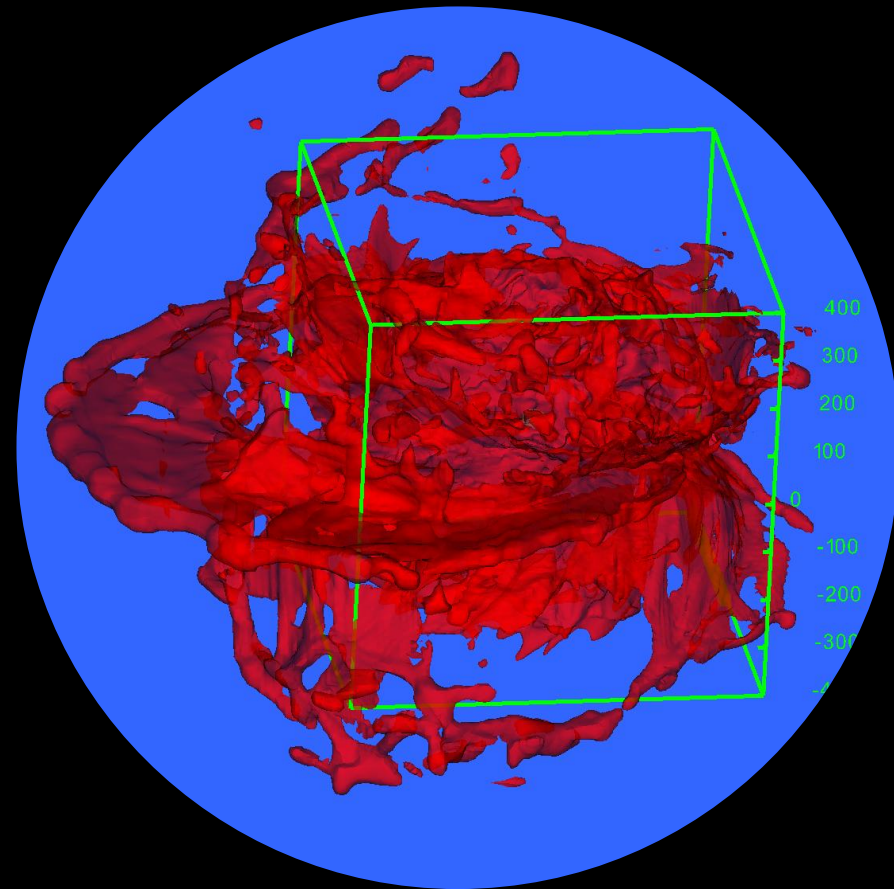


NUCLEOSYNTHESIS: 3rd Peak



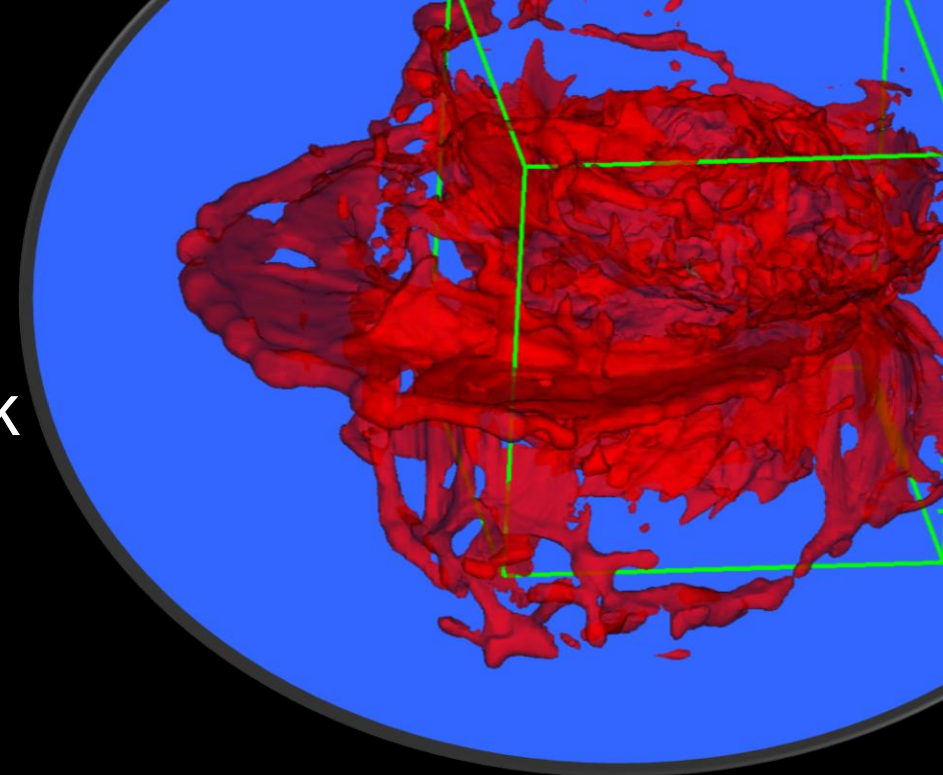
Venturi, GR et al. (in prep)

Conclusion



Conclusion & Future Prospective

- Poloidal Magnetic field suppress the 3rd peak
- Effect comparable to EoS contribution
- Enhanced ejection in the pole
 - Nucleosynthesis in the "Jet"/Angular distribution
- Post-processing code to extract tracer data!



Venturi, GR et al. (in prep)



Fabrizio Venturi



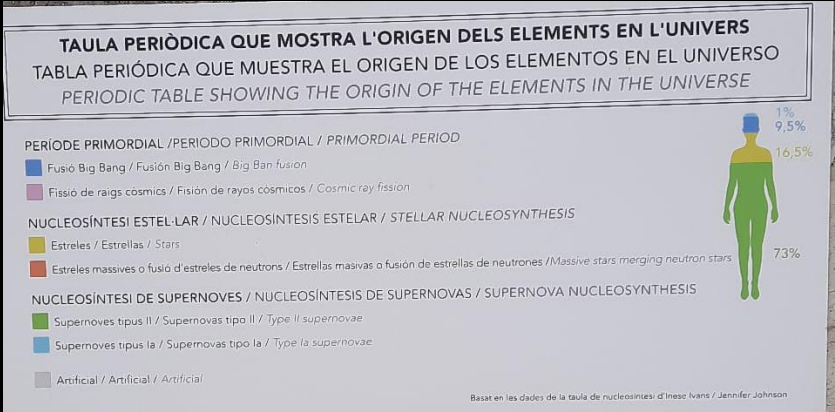
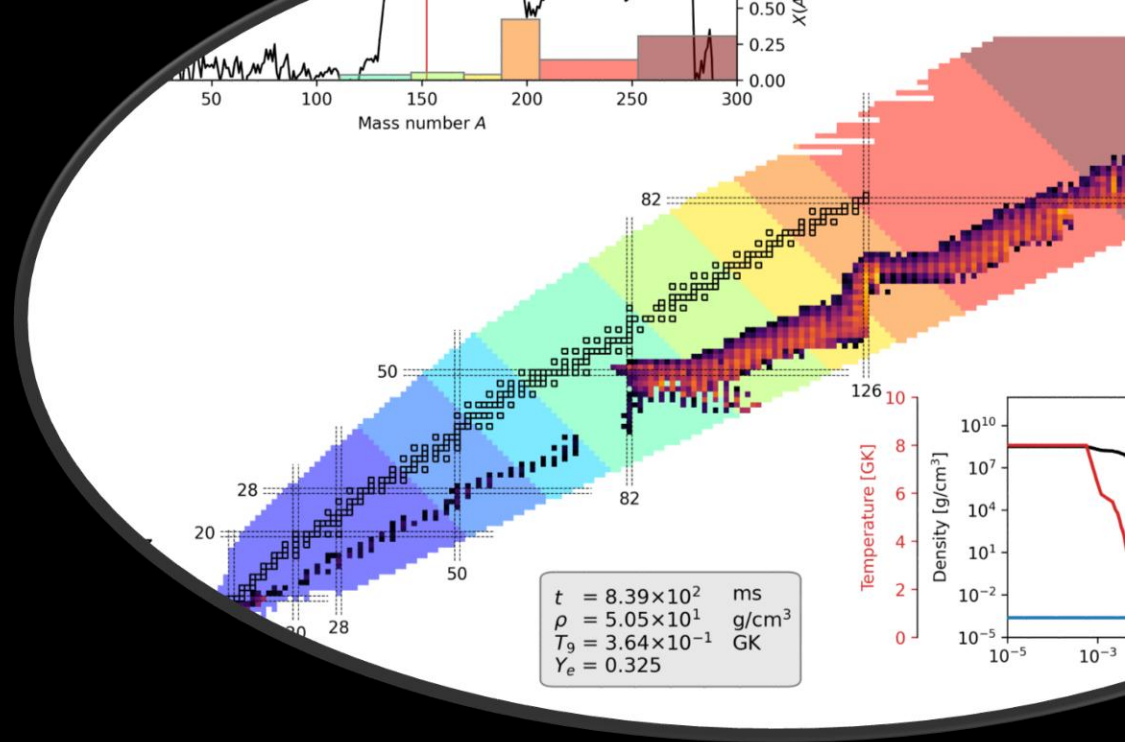
Elina Ghai



Milton Ruiz



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Thanks!
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