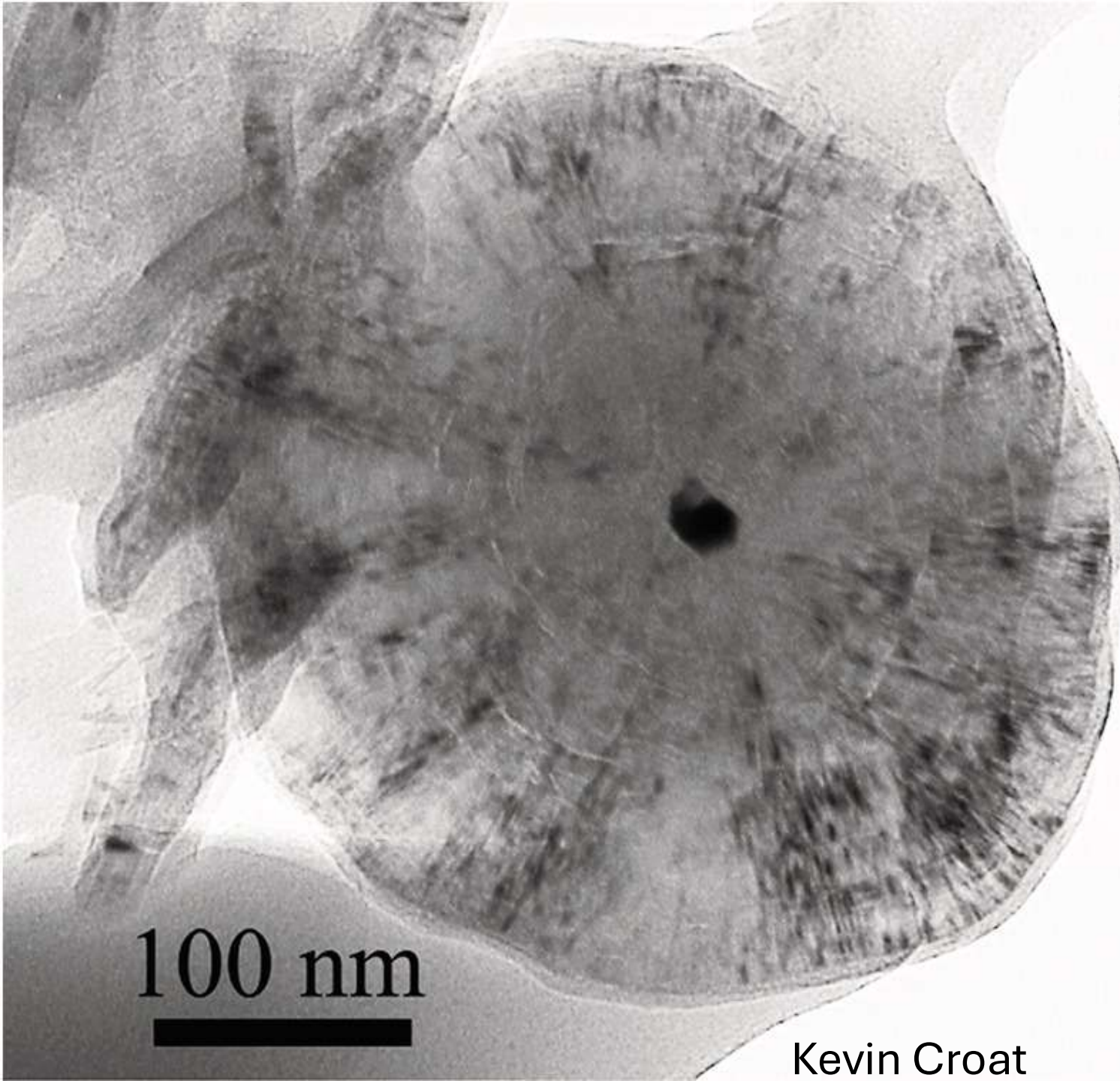




Effects of Metallicity on TiC, SiC, and Graphite Condensation

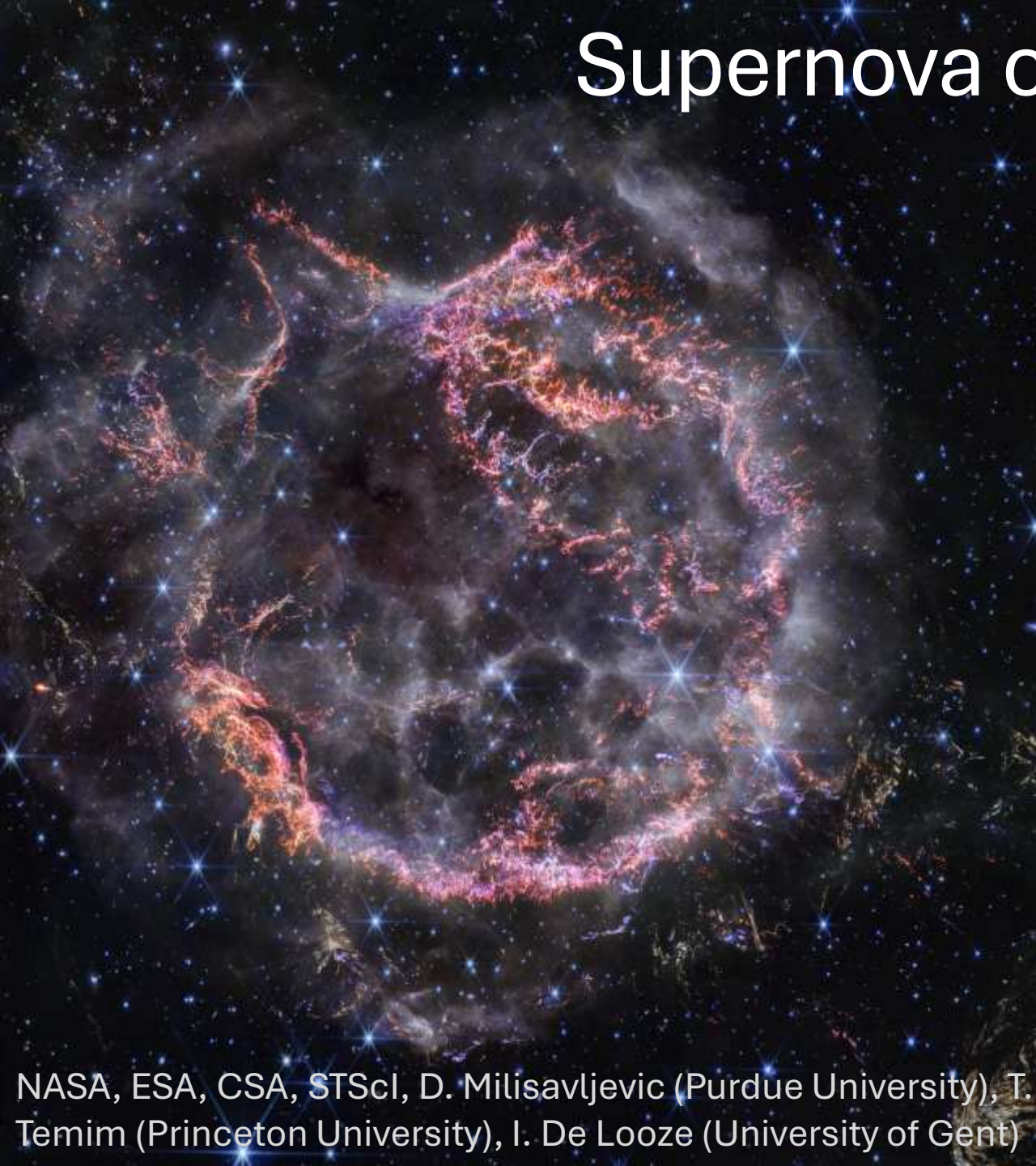
Gabrielle Adams and Katharina Lodders
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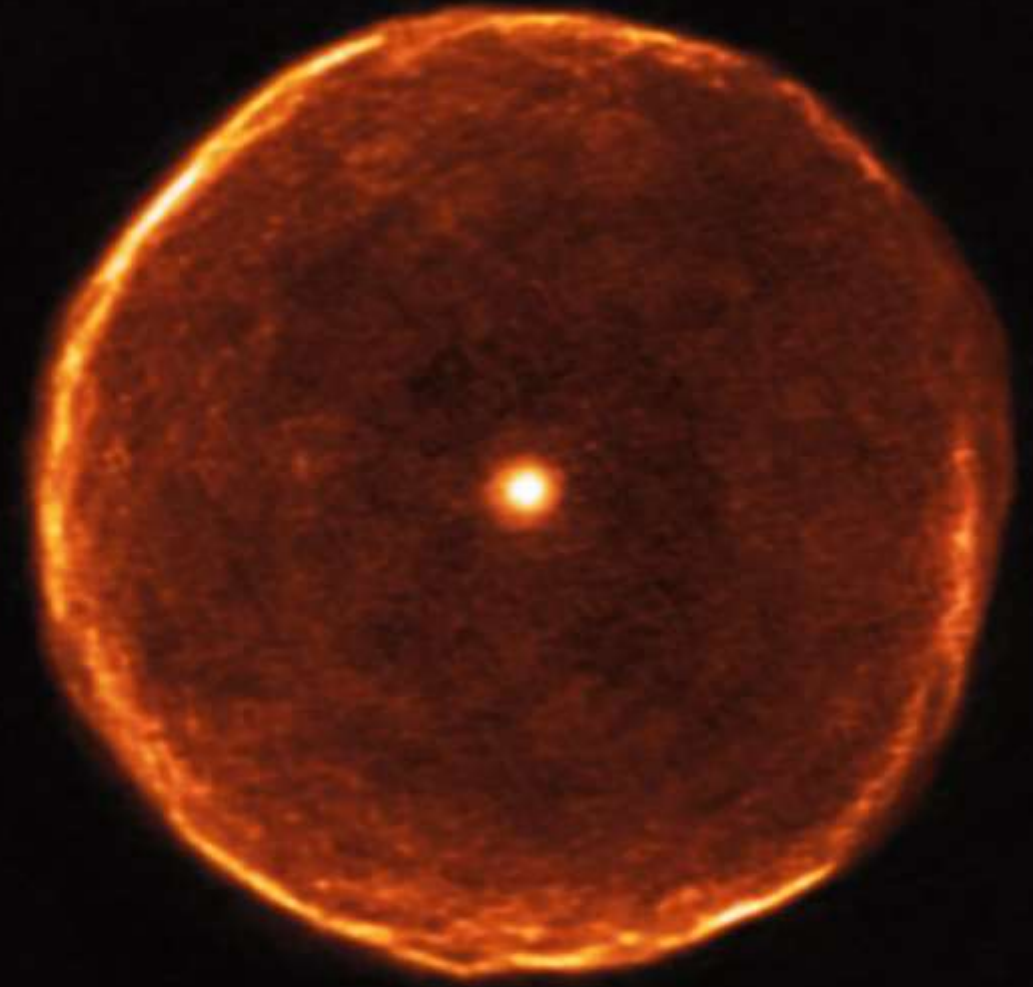
a

- Ancient dust preserved in primitive meteorites
- Present in solar nebula
- A (biased) sample of presolar material

Supernova or AGB star



NASA, ESA, CSA, STScI, D. Milisavljevic (Purdue University), T. Temim (Princeton University), I. De Looze (University of Gent)

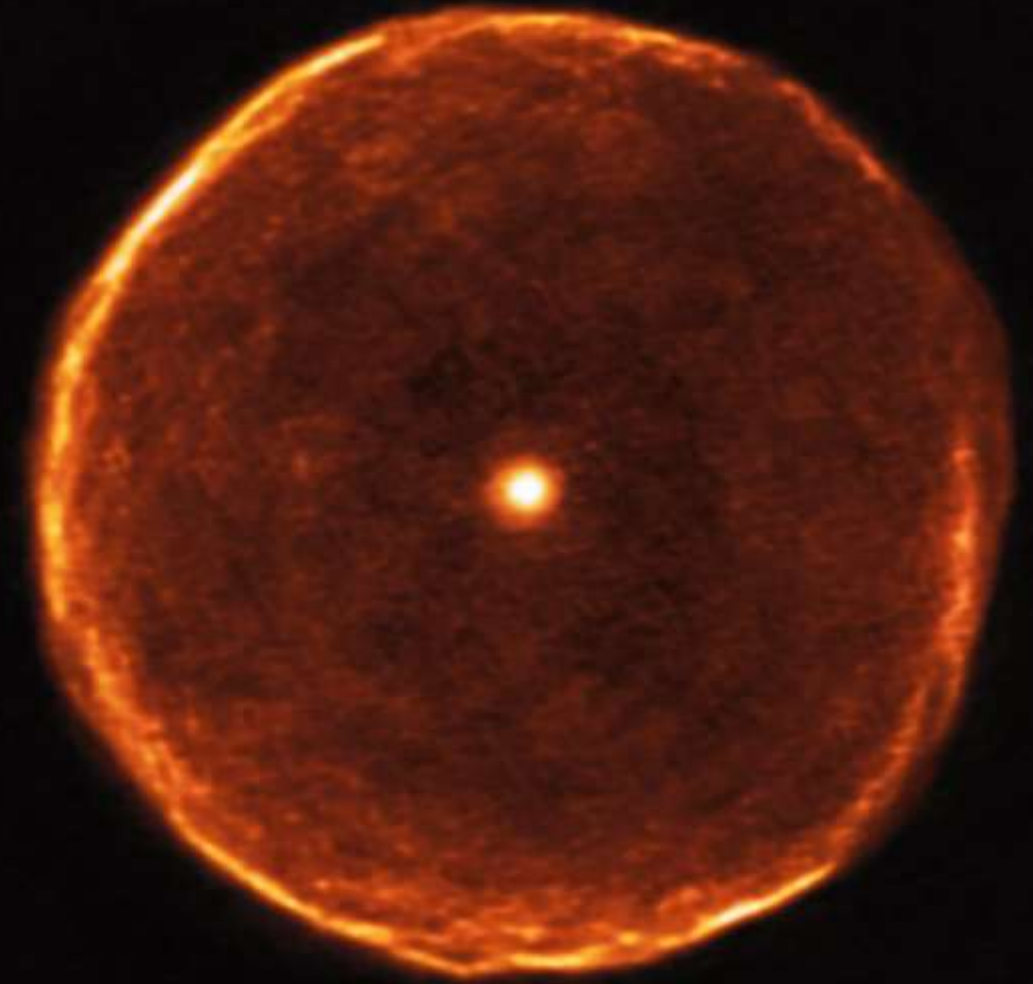


ALMA (ESO/NAOJ/NRAO)/F. Kerschbaum

Supernova or AGB star



NASA, ESA, CSA, STScI, D. Milisavljevic (Purdue University), T. Temim (Princeton University), I. De Looze (University of Gent)



ALMA (ESO/NAOJ/NRAO)/F. Kerschbaum

Three types of AGB stars:

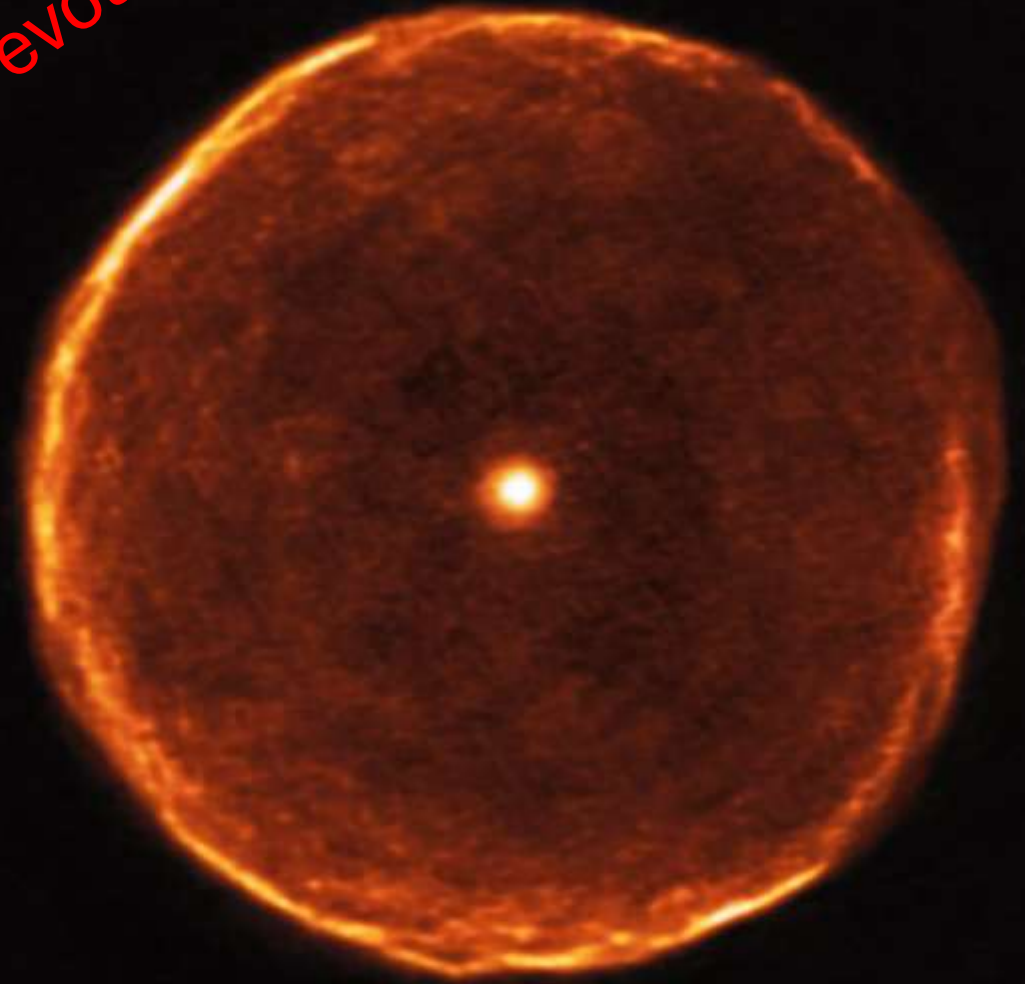
M-type: $C/O < 1$

S-type: $C/O \sim 1$

C-type: $C/O > 1$

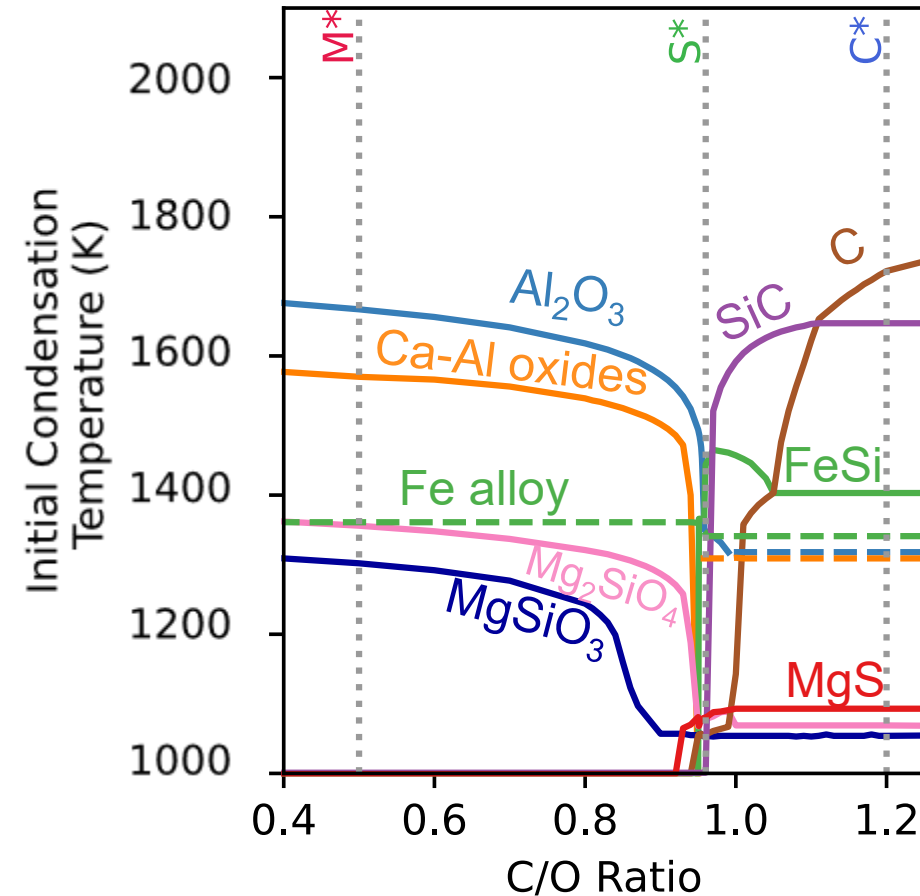


More evolved



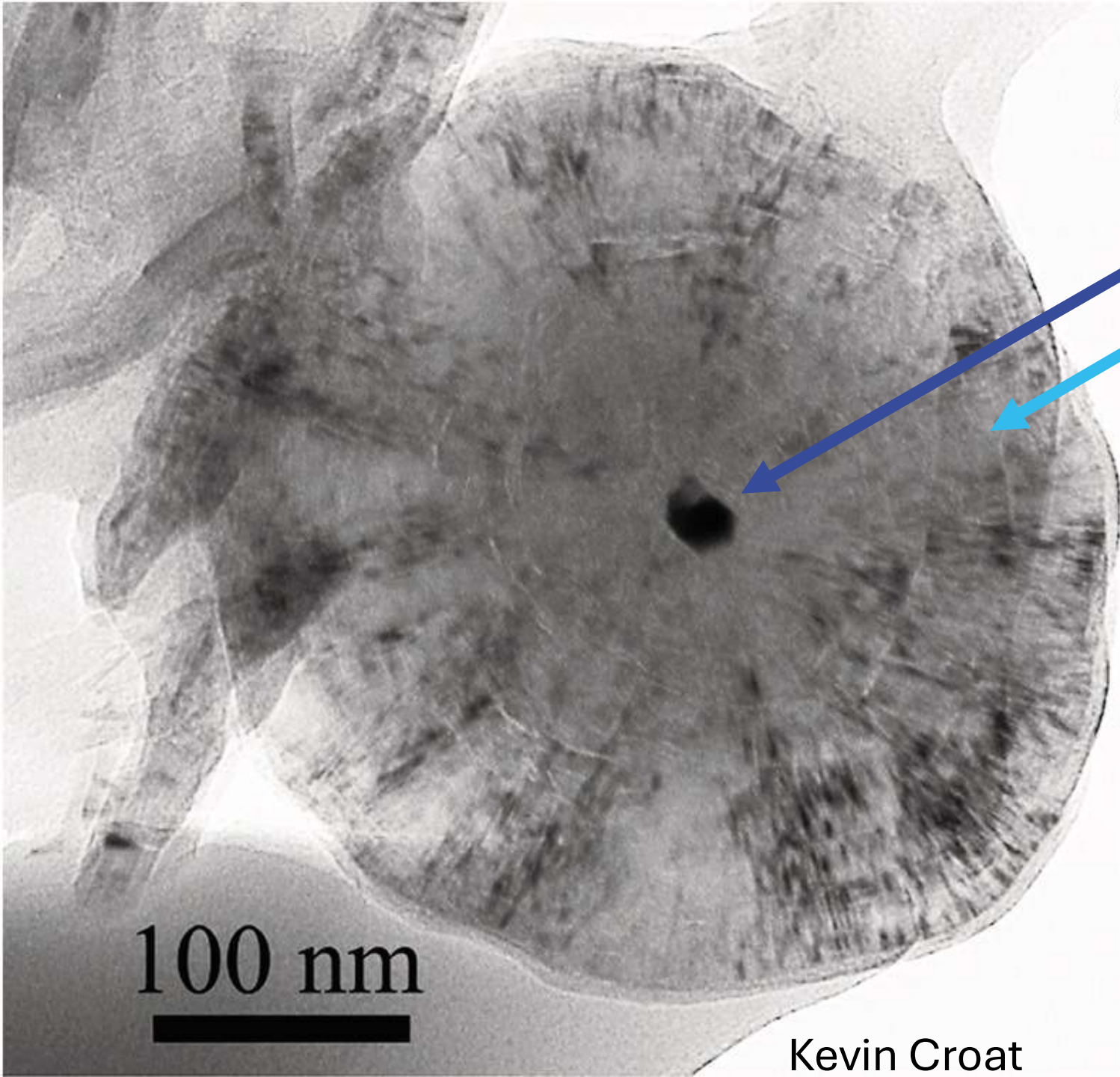
Condensation Computations

- Useful to interpret and predict grain composition
- Depend on total pressure and bulk composition
- C/O ratio
 - Solar C/O is ~ 0.56
 - O-rich: silicates and oxides
 - C-rich: carbides and graphite



How to change bulk elemental composition

- C/O ratio
- **Metallicity**
 - Atomic ratio of a heavy element to hydrogen relative to the respective ratio in the Sun
 - Condensation at non-solar metallicity has not been investigated beyond a few exploratory calculations



a TiC-C-SiC

Graphite grain with
central refractory
carbide crystal

100 nm

THE ASTROPHYSICAL JOURNAL

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Effects of Metallicity on Graphite, TiC, and SiC Condensation in Carbon Stars

Gabrielle M. Adams and Katharina Lodders

Published 2025 May 13 • © 2025. The Author(s). Published by the American Astronomical Society.

[The Astrophysical Journal](#), [Volume 985](#), [Number 1](#)

Citation Gabrielle M. Adams and Katharina Lodders 2025 *ApJ* **985** 35

DOI 10.3847/1538-4357/adc40f

Method

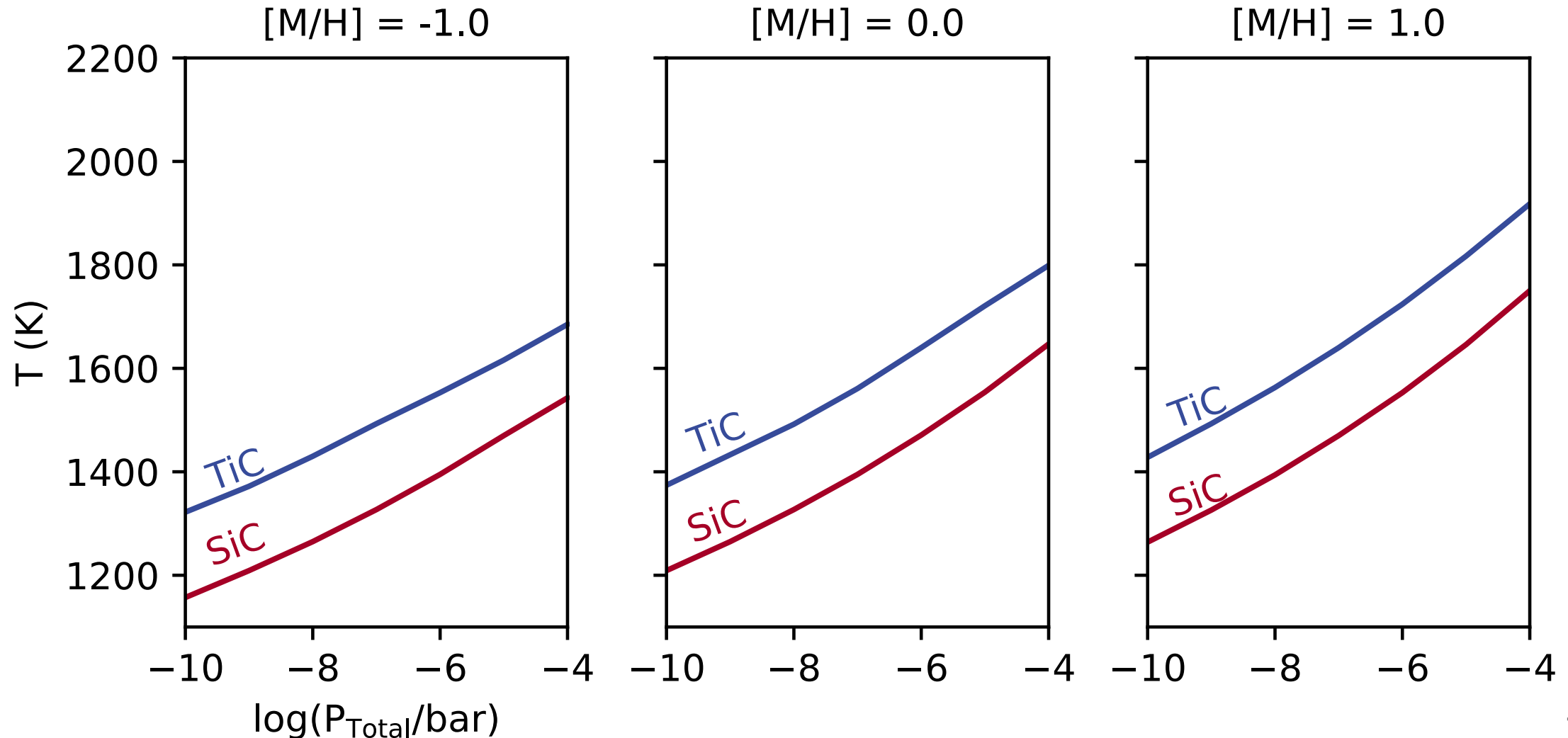
- Thermodynamic equilibrium calculations
- Reducing conditions: $C/O = 1.2$ (Lambert et al. 1986)
- Uniform metallicity factor for all elements heavier than He
- Ranging from 0.01 to 100 times solar metallicity, $[-2]$ to $[2]$
- $C/O = 1.1$ and 3.0 also investigated

Key points

- Increasing metallicity increases condensation temperatures
- TiC-C-SiC is favored at subsolar to solar metallicities and intermediate to high pressures
- Metallicity determines the pressure dependence of graphite condensation

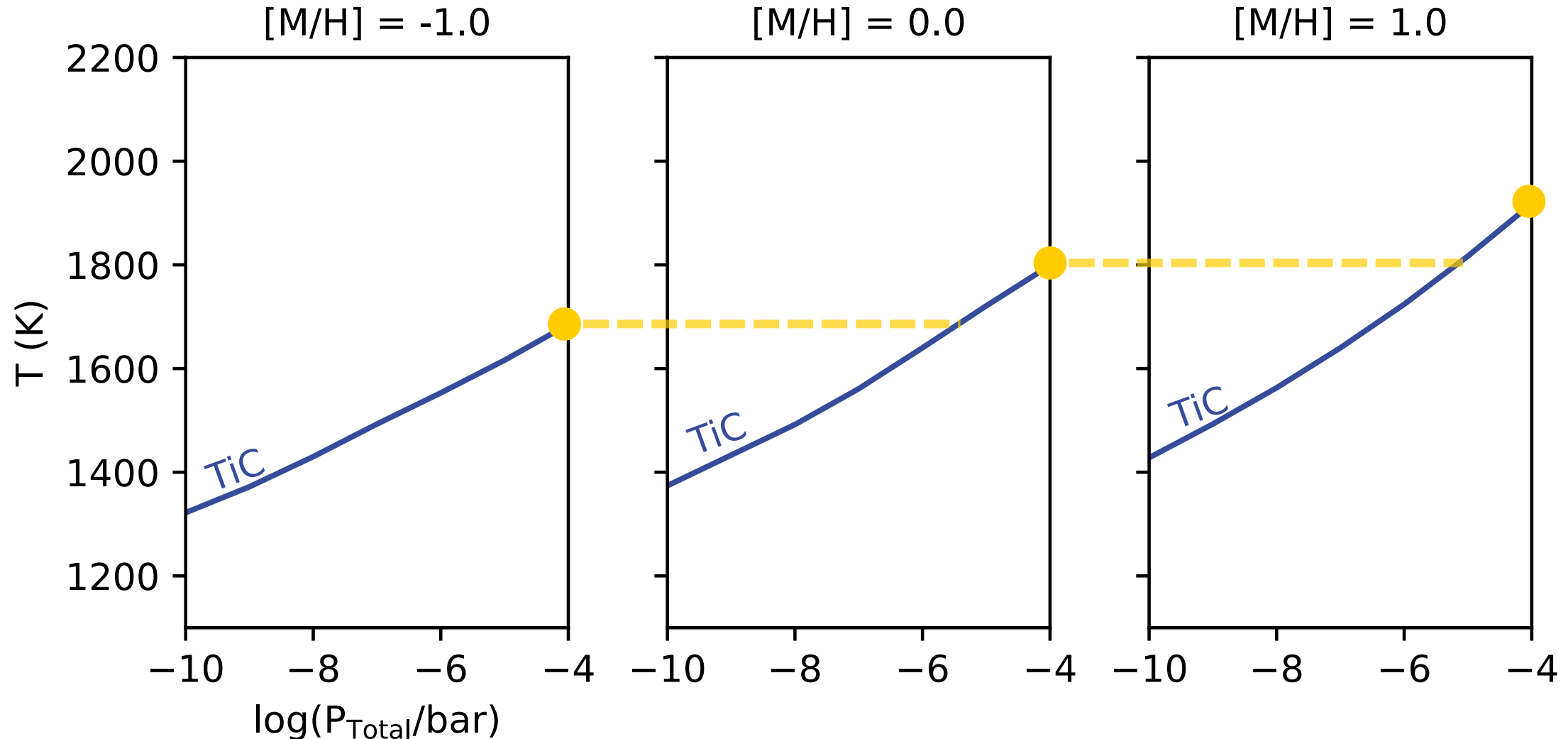
Increasing Metallicity Increases Condensation Temperatures

Starting Condensation Temperatures for TiC, SiC, and C, C/O = 1.2



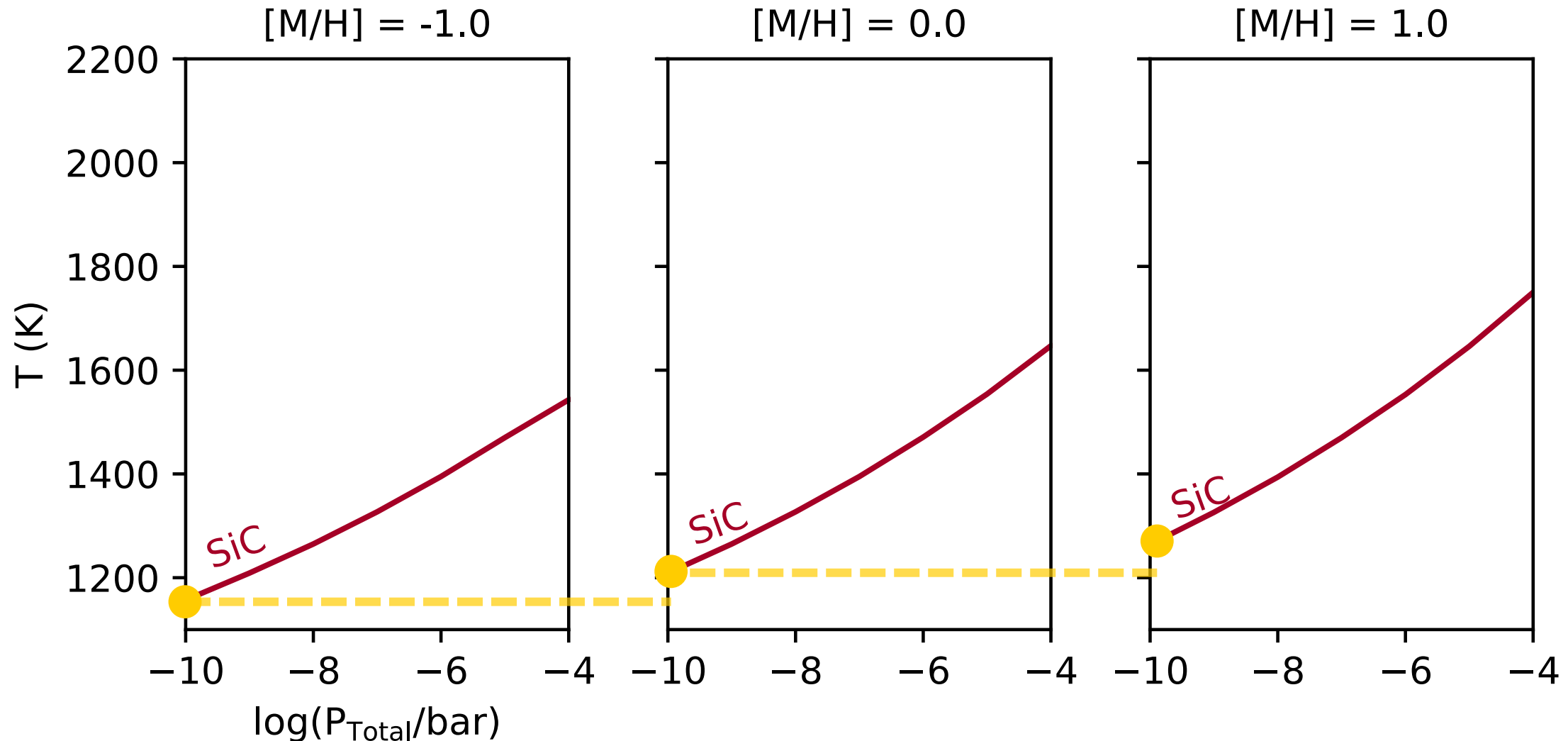
Increasing Metallicity Increases Condensation Temperatures

Starting Condensation Temperatures for TiC, SiC, and C, C/O = 1.2



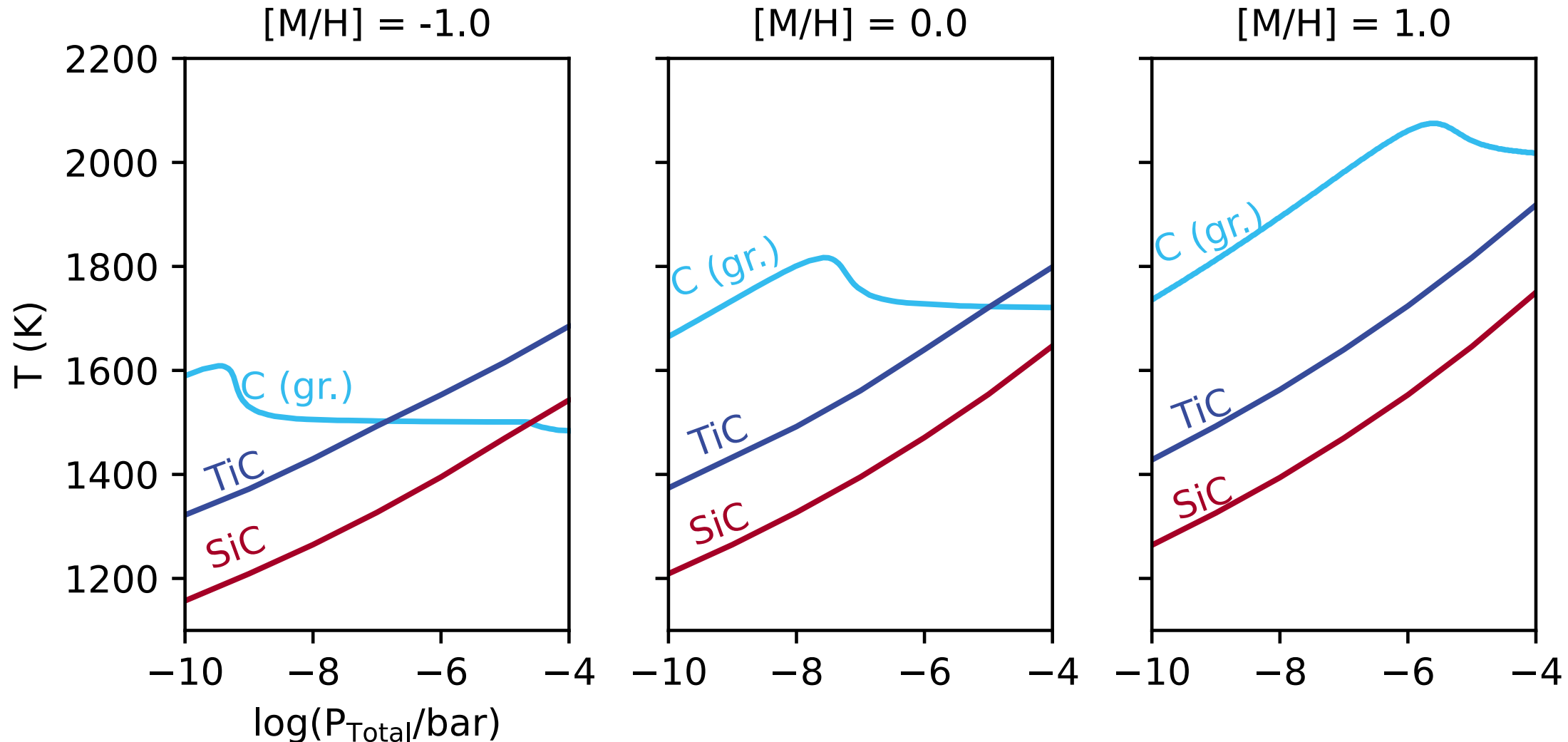
Increasing Metallicity Increases Condensation Temperatures

Starting Condensation Temperatures for TiC, SiC, and C, C/O = 1.2



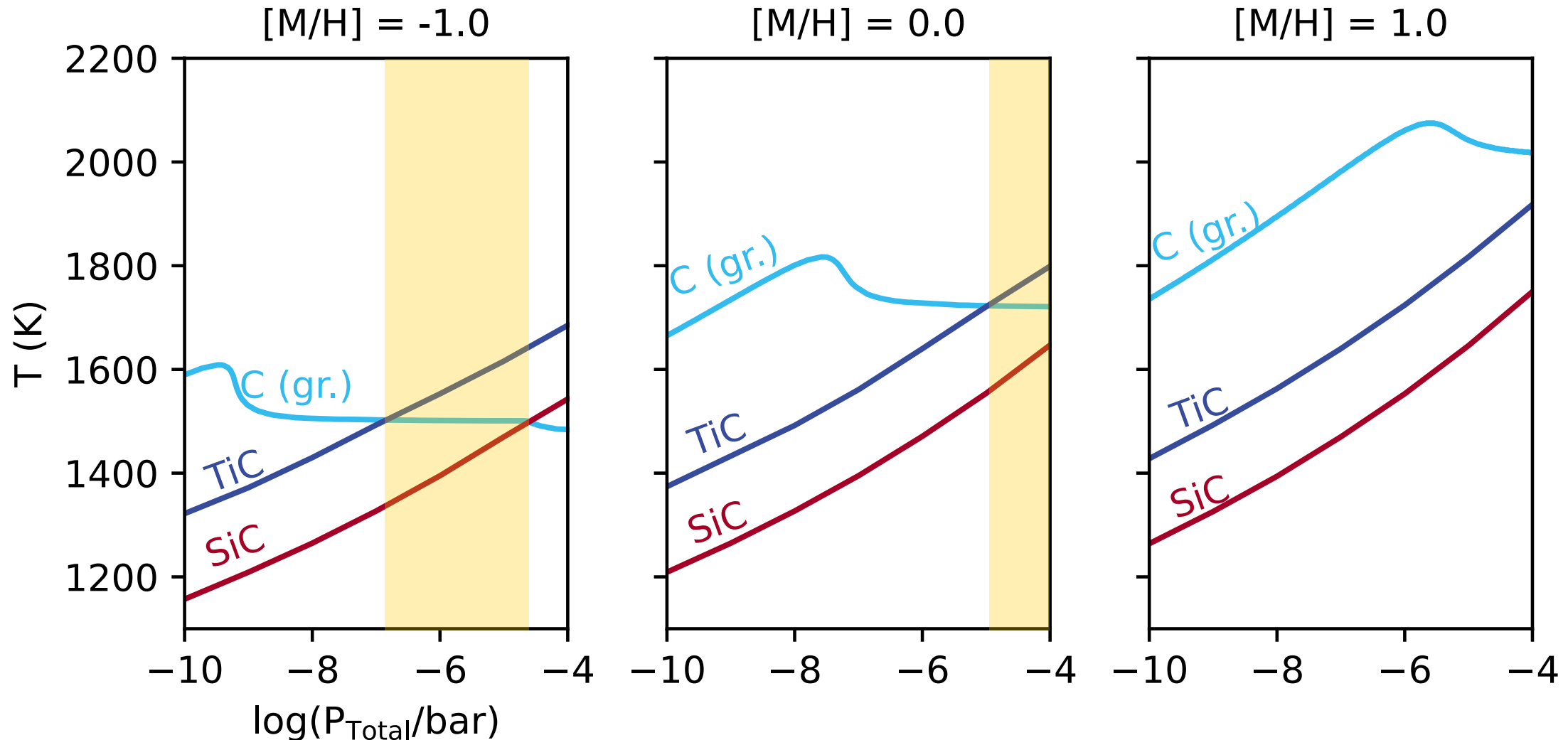
TiC-C-SiC is Favored at Subsolar to Solar Metallicities

Starting Condensation Temperatures for TiC, SiC, and C, C/O = 1.2



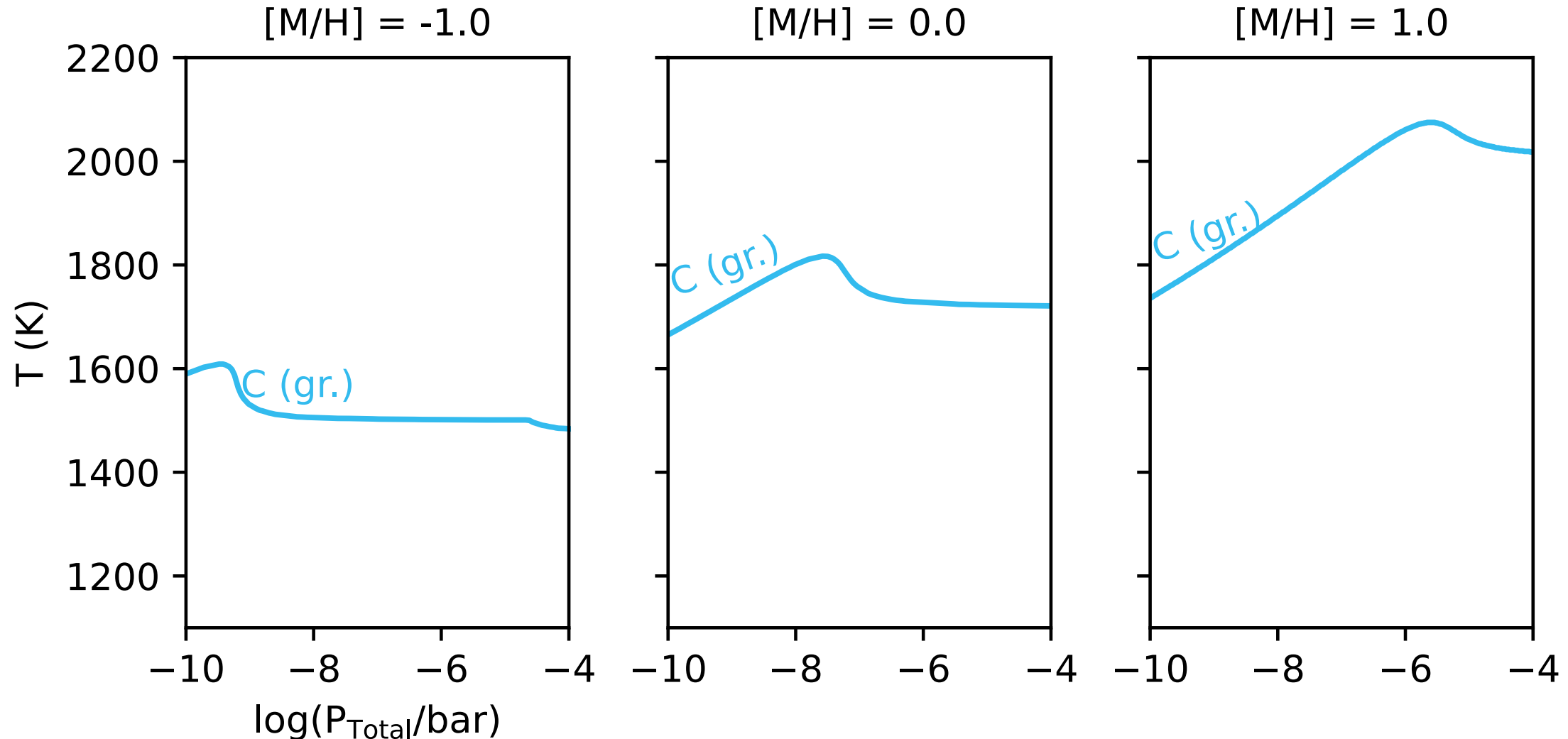
TiC-C-SiC is Favored at Subsolar to Solar Metallicities

Starting Condensation Temperatures for TiC, SiC, and C, C/O = 1.2

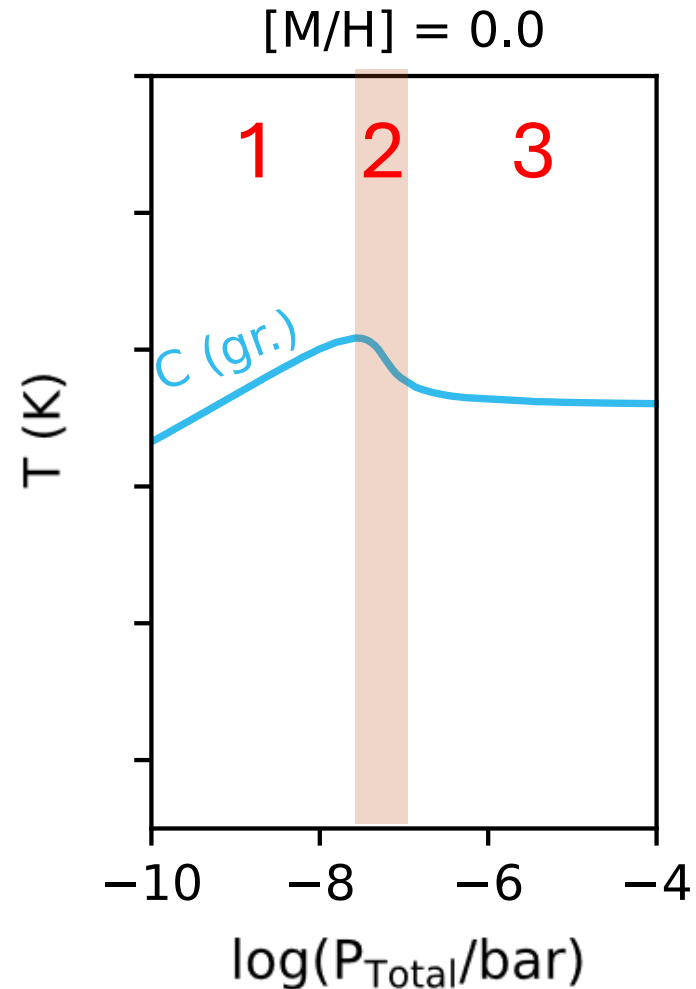


Metallicity Determines the Pressure Dependence of Graphite Condensation

Starting Condensation Temperatures for TiC, SiC, and C, C/O = 1.2

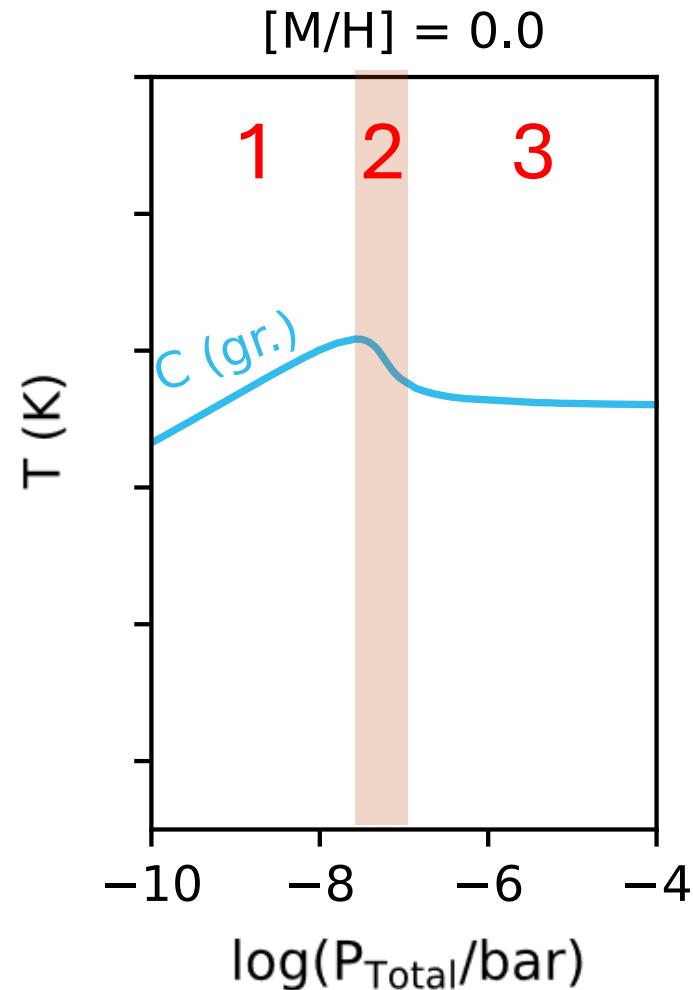


Metallicity Determines the Pressure Dependence of Graphite Condensation



3 Regions

Metallicity Determines the Pressure Dependence of Graphite Condensation

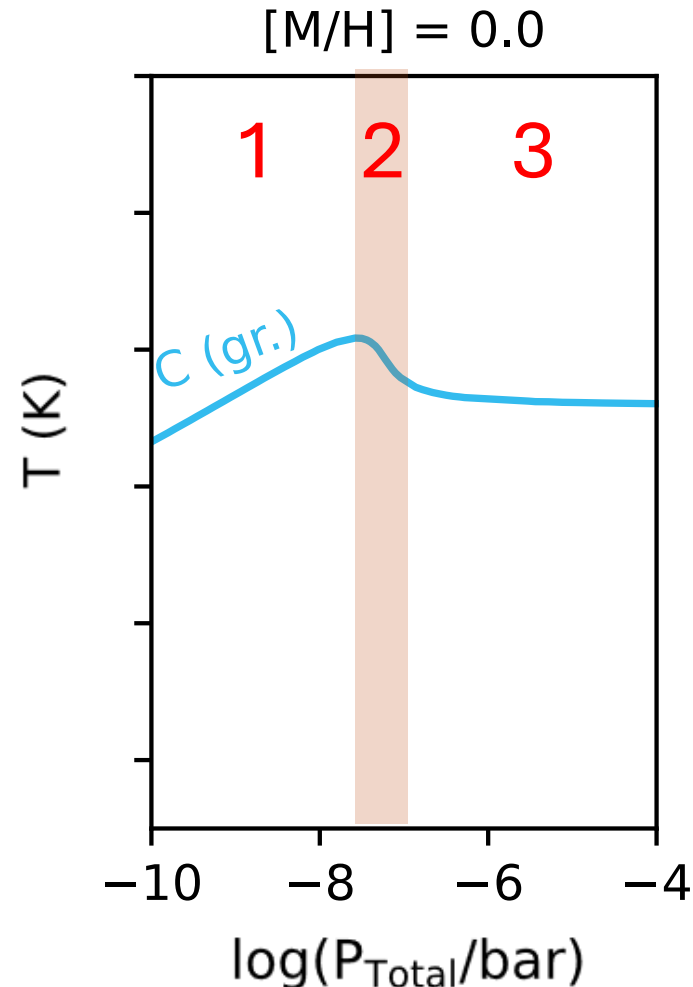


1 Monatomic H and C

2

3

Metallicity Determines the Pressure Dependence of Graphite Condensation

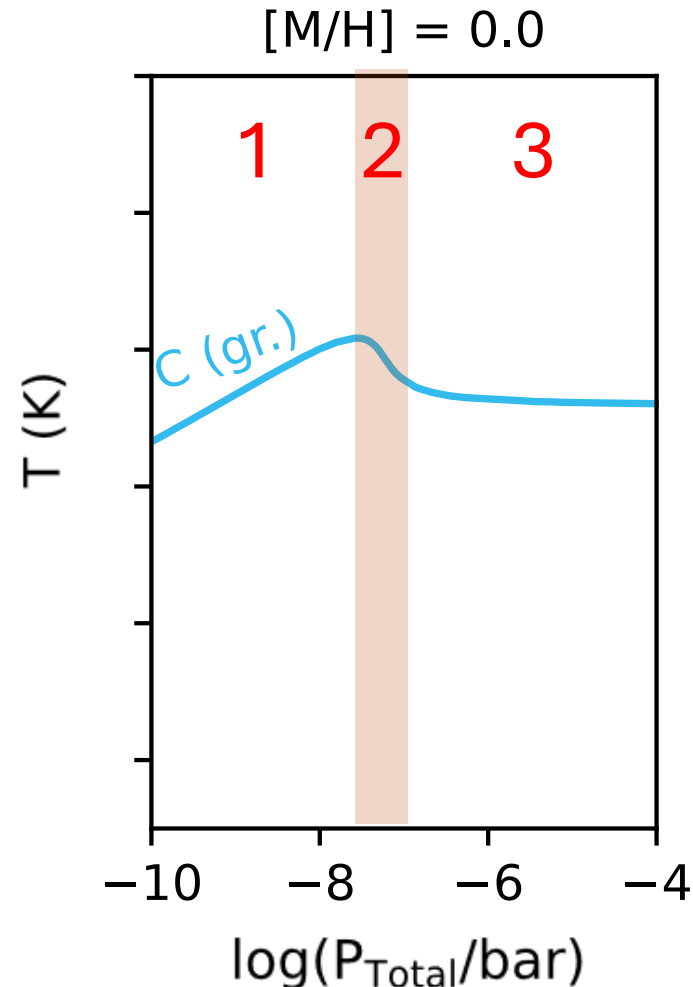


1 Monatomic H and C

2 C_2H_2 and monatomic H

3

Metallicity Determines the Pressure Dependence of Graphite Condensation

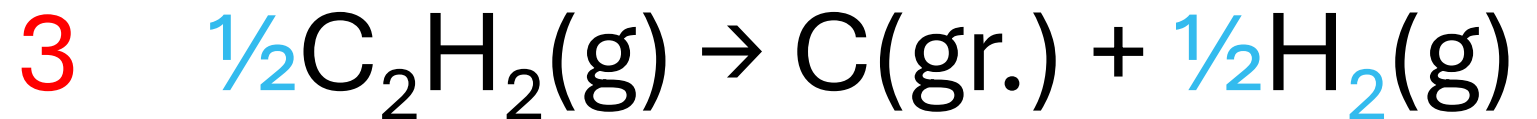
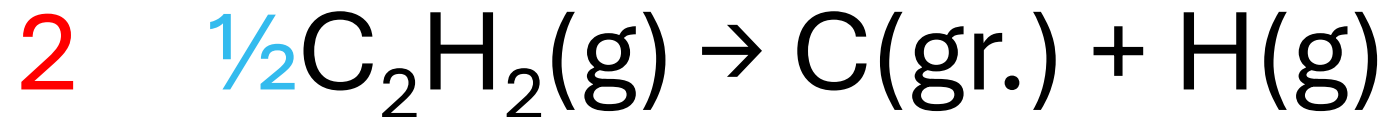
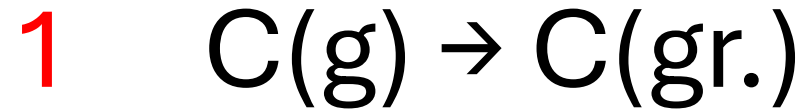
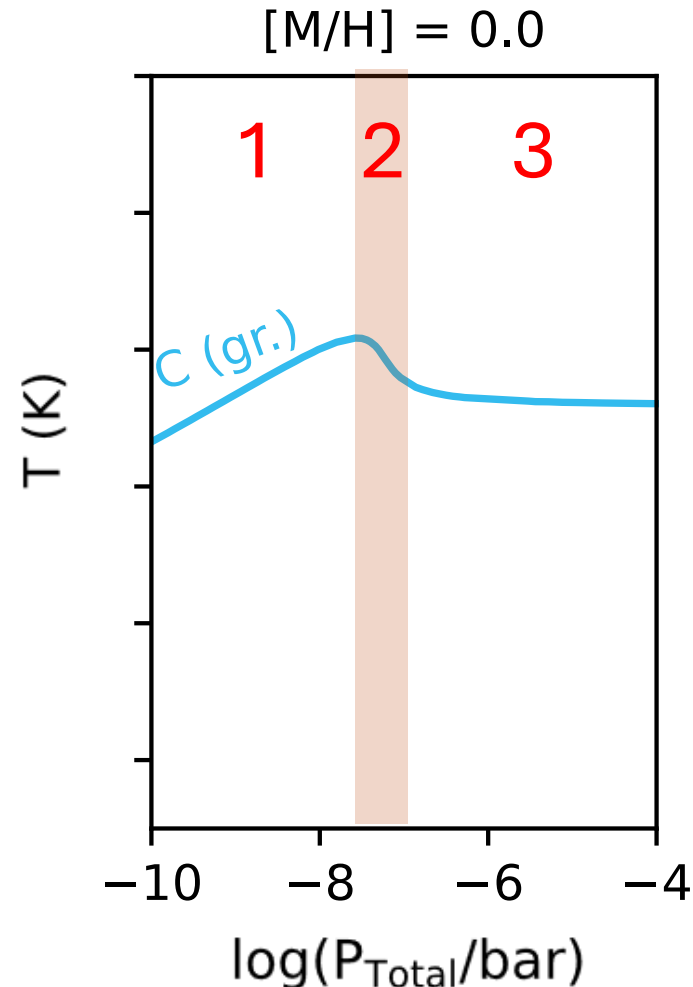


1 Monatomic H and C

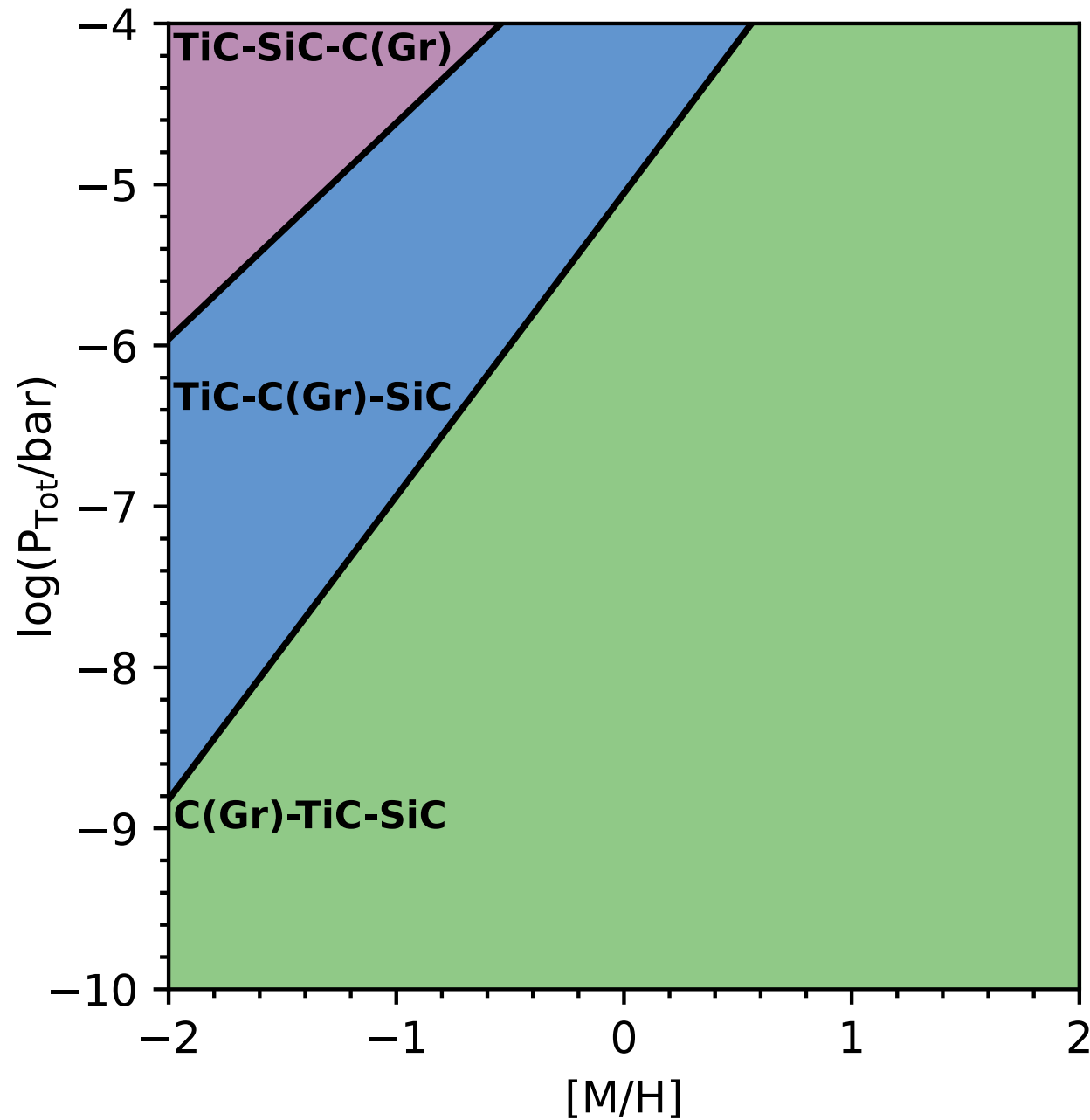
2 C_2H_2 and monatomic H

3 C_2H_2 and H_2

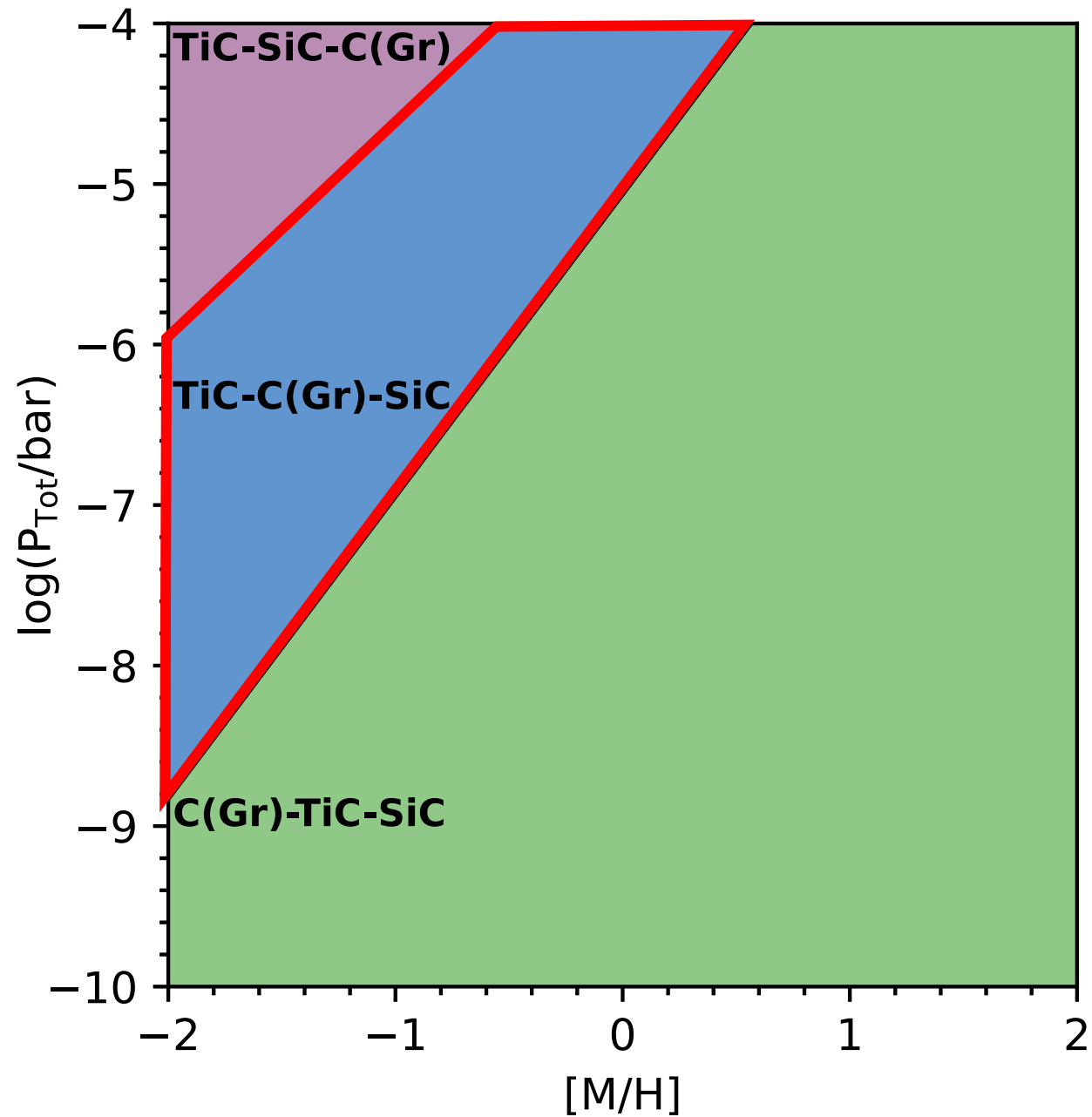
Metallicity Determines the Pressure Dependence of Graphite Condensation



C-bearing Condensation Sequence



C-bearing Condensation Sequence



Key points

- Increasing metallicity increases condensation temperatures
- Metallicity determines the gas chemistry, which determines the pressure dependence of graphite condensation
- TiC-C-SiC is favored at subsolar to solar metallicities and intermediate to high pressures

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In progress: change C/O ratio and [M/H]

