Combined search in dwarf spheroidal galaxies for branon dark matter annihilation signatures with the MAGIC Telescopes

Tjark Miener¹, D. Nieto¹, V. Gammaldi², D. Kerszberg³, J. Rico³ for the MAGIC Collaboration

¹IPARCOS, Universidad Complutense de Madrid, ²Universidad Autónoma de Madrid, ³IFAE, The Barcelona Institute of Science and Technology (BIST)

Abstract

One of the most pressing questions for modern physics is the nature of dark matter (DM). Several efforts have been made to model this elusive kind of matter. The largest fraction of DM cannot be made of any of the known particles of the Standard Model (SM). We focus on brane world theory as a prospective framework for DM candidates beyond the SM of particle physics. The new degrees of freedom that appear in flexible brane world models, corresponding to brane fluctuations, are called branons. They behave as weakly interacting massive particles (WIMPs), which are one of the most favored candidates for DM. We present a multi-target DM search in dwarf spheroidal galaxies for branon DM annihilation signatures with the ground-based gamma-ray telescope MAGIC leading to the most constraining branon DM limits in the TeV mass range.

Introduction

Astrophysical and cosmological evidences propose that non-baryonic cold DM accounts for 84% of the matter density of the Universe. Nevertheless, the nature of DM is still an open question for modern physics, as this elusive kind of matter can not be made of any of the SM particles. In this contribution, we model the gamma-ray emission in the source region with the branon DM model and perform a multi dwarf spheroidal galaxies (dSphs) DM analysis using gLike^a and LikelihoodCombiner^b, open-source analysis tools for multi-target and multi-instrument DM searches.

Branon dark matter

Brane-world theory [1] has been put forward as a prospective framework for DM candidates. The characteristics of the suggested massive brane fluctuations in this theory match the ones of weakly WIMPs, which are one of the most favored candidates of cold DM. The left panel of Fig. 1 shows the branon branching ratios Br_i as a function of the branon mass m_{χ} . The differential photon yield per branon annihilation dN/dE is depicted for a set of branon DM masses (from light to dark: 0.1, 0.2, 0.5, 1, 2, 5, 10, 20, 50 and 100 TeV) in the right panel of Fig. 1.



^a<u>https://github.com/javierrico/gLike</u>

DOI 10.5281/zenodo.4601451

Dwarf observations with the MAGIC telescopes MAGIC recently published in [2] a combined model independent DM search in the dSphs Segue 1 (158h), Ursa Major II (95h), Draco (52h), and Coma Berenices (49h) with a total exposure of 354h.

Results

MAGIC Seque 1 limits (JCAP05(2022)00 dSphs (354 h We present the observational 95% confidence level (CL) upper limits (UL) on the Excluded by AMS-02 cluded by CMS A sensitivity (300 h) thermally-averaged cross-section $\langle \sigma v \rangle$ (see Fig. 2) and on the brane tension f (see 95% containmen Fig. 3) for branon DM annihilation obtained with 354h of multi-dSph observations 10⁰ 10^{1} 10^{-} with the MAGIC telescopes. We perform a model dependent search for branon DM $m_{\chi}[TeV]$ particles of masses between 100 GeV and 100 TeV. As expected from the no significant gamma-ray excess in the data, our constraints for branon DM annihilation Fig. 2: Observational 95% CL UL on the thermally-averaged cross-section are located within the 68% containment band, which is consistent with the nodetection scenario. We obtain **our strongest limit** $\langle \sigma v \rangle \approx 5.9 \times 10^{-24} \text{ cm}^3 \text{s}^{-1}$ for a Conclusion



~0.7 TeV branon DM particle mass.



tmiener@ucm.es



Acknowledgements

-82729-C6-5-R. FPA2017-90566-REDC. PID2019-PID2019-107847BB-C44 PID2019-107988GB-C22): the Indian Department of Atomic Energy Tokyo, JSPS, and MEXT: the Bulgarian Mi aría de Maeztu" CEX2019-000918-M. the Unidad de Excelencia "María de Maeztu 18-2 and the "la Caixa" Foundation (fellowship LCF/BQ/PI18/11630012), by the Croatian Science Foundatio HrZZ) Project IP-2016-06-9782 and the University of Rijeka Project 13,12,1,3,02, by the DFG Collaborative Research Centers 3823/C4 and SFB876/C3, the Polish National Research Centre grant UMO-2016/22/M/ST9/00382 and by the Brazilia ICTIC, CNPg and FAPERJ

Economy, Industry, and Competitiveness / European Regional Development Fund grant FPA2015-73913-JIN. VG's contribution to this work has been supported by Juan de la Cierva-Incorporacion IJC2019-040315-I grant, and by the PGC2018-095161-B-I00 and CEX2020-001007-S projects, both funded by MCIN/AEI/10.13039/501100011033 and by "ERDF A way of making Europe". VG thanks J.A.R. Cembranos for useful discussions.

DK is supported by the European Union's Horizon 2020 research and innovation programme under the Marie Sklodowska-Curie grant agreement No. 754510. DK and JR acknowledge the support from MCIN/AEI/ 10.13039/501100011033 under grants PID2019-107847RB-C41 and SEV-2016-0588 ("Centro de Excelencia Severo Ochoa"), and from the CERCA institution of the Generalitat de Cataluny

^D<u>https://github.com/TjarkMiener/likelihood_combiner</u>



Fig. 2: Observational 95% CL UL on the brane tension.

This work leads to the **most constraining branon DM limits in the TeV mass range**, superseding previous constraints by CMS, AMS-02 and MAGIC Segue1 limits [3]. We can achieve even more stringent exclusion limits by adding further dSph observations of the MAGIC telescopes or other gamma-ray or neutrino telescopes to this analysis [4].

References

[1] J. A. Cembranos et al., Brane-World Dark Matter [arXiv:0302041]

[2] MAGIC Collaboration, Combined searches for dark matter in dwarf spheroidal galaxies observed with the MAGIC telescopes, including new data from Coma Berenices and Draco [arXiv:2111.15009]

[3] T. Miener et al., Constraining branon dark matter from observations of the Segue~1 dwarf spheroidal galaxy with the MAGIC telescopes [arXiv:2201.03344]

[4] C. Armand et al., Combined Dark Matter Searches Towards Dwarf Spheroidal Galaxies with Fermi-LAT, HAWC, HESS, MAGIC and VERITAS [arXiv:2108.13646]

DOI 10.5281/zenodo.4597500





