ray Space Telescope



# Periodic Variability in gamma-ray Emitting Blazars with Fermi-LAT

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#### Introduction



- Blazars: variability in the overall electromagnetic spectrum:
  - Different timescales:
    - Long-term variations
    - Short variations
- Pattern  $\rightarrow$  **Periodicity**
- Astrophysical nature: <u>single or a binary SMBH</u>
- Find a sample with  $\gamma$ -ray periodic-emission



#### **Previous Results**



START AGNFile.txt 3FGL+2FHL+3FHL LSP + Discard AGN False **Power-Law Fitting** False blazars (351) True Removing **Read AGN properties** REDFIT True **Upper Limits** < 50% upper-limits DFT (Welch) -Selection Telescope time:  $\bullet$ END Candidate Criteria Aug. 2008-Sep. 2017 PDM -Save Candidate Save GLSP + Bootstrap False False Low-Significance Data Reduction: Candidate **MCMC** Sine Fitting Flux integrated  $\geq 1$  GeV False SP + Simulated LCs True True True 28-days binning **Bayesian QPO** Selection Candidate High-Significance Candidate wwz Constraints Candidate Criteria Methods: uneven Constraints Periodicity detection: 10 even Significance estimation: 4

Peñil et al. 2020

#### **Previous Results**



- 11 High-significance candidates
  (≥4σ)
  - 9 New detections
- 13 Low-significance candidates (>2.5σ)
  - 9 New detections
- 6 objects previously reported in the literature:
  - 5 with the same period
  - S5 0716+ 714 (2.9 yr)
- False-positive detection rate: 1 detection





- 24 periodicity candidates from previous work Peñil et al., 2020
- Light curves:
  - Telescope time: August 2008-December 2020
  - Extended with 3 extra years  $\rightarrow$  total of 12 years
- Data reduction:
  - Flux integrated ≥100 MeV (Reduction of upper-limits)
  - 28-days binning

## **Extended Pipeline**



- Full Width Half Maximum (FWHM)
- ARIMA/ARFIMA analysis
  - Robust against stochastic noise (Feigelson et al., 2018)
- Power Spectral Index ( $\beta$ ):
  - Power Spectrum Response method (Uttley et al. 2002)
- Flux Distribution:
  - Log-Normal & Normal
  - Shapiro-Wilk Test & MLE



## **Results:** 6 Blazars $\geq 5\sigma$ (local) periodicity detection





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#### Results

- Genuine periodicity: larger number of cycles, increase the significance (Vaughan, et al. 2015)
  - 14 blazars increase their significance
  - 4 blazars same significance
  - 6 blazars decrease their significance
- Power Spectral Index:
  - β=[0.9-1.5]
  - Jet modulations are likely coupled to the disk modulations (e.g, Abdo, et al. 2010)
- Flux Distribution:
  - Log-normal distribution
  - Fluctuations in the accretion disk to produce an aggregate multiplicative effect transmitted to the jet (e.g., Shah, et al., 2018)

# **Global Significance**



- Local Significance (pipeline techniques)
- Look-elsewhere effect:
  - PGlobal=1-(1-PLocal)^N
  - N: trial factor
- N =Nblazars\*indep. frequencies:
  Nblazars: 351



- Indep. frequencies:
  - Bottom-limit: 11 (# points in LC, samples/year)
  - Upper-limit: 100 (# frequencies in the periodograms)
  - Monte-Carlo simulations  $\rightarrow$  indep. frequencies=35
  - $\circ$  5 $\sigma \rightarrow$  3.5 $\sigma$



- Calibrated the significance of the methods:
  - Calculate the number of  $X\sigma$  detections with artificial LCs
  - LCs with same PSD and PDF (Emmanoulopoulos, D., et al. 2013)
    - ∆*o* = [8%-15%]
  - LCs based on a power-law [ $\beta$ - $\Delta$ ,  $\beta$ ,  $\beta$ + $\Delta$ ] (Timmer and Koenig, 1995)

 $\Delta \sigma = [9\%-25\%]$ 

- Evaluation of the method's detection against the noise:
  - Sinusoidal signal:
    - Different periods [1.5-4.5] yrs
  - Contaminated with noise :
    - white, pink, red
  - Methods more robust against pink\red noise: detection in [25%-65%]



- We confirm the evidence of periodicity in 18 blazars:
  - We find 6 blazars with periodicity detected with  $3.5\sigma$  (global significance)

- Future Work:
  - Multiwavelength and cross-correlation study of the 24 blazars
  - New periodicity analysis of all blazars in 4FGL catalogue: ~3000 blazars