Approaching a theory of galaxy formation

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Approaching a theory...

- There is not an established theory of galaxy formation
- Cosmology as a framework
- "Concordance cosmology" was established in the last decades of the XX century

ACDM cosmology

What makes a galaxy?

galaxy constituents

- dark matter, stars, gas, and dust
- galaxies are complex, multi-component systems

galaxy diversity

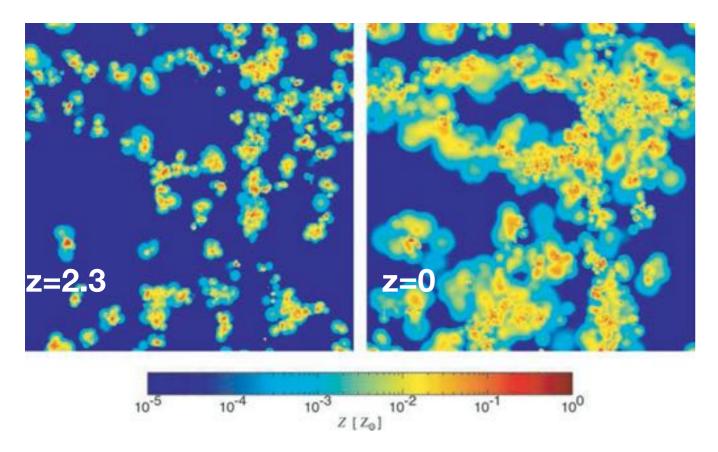


Galaxy Astrophysics

- Gas and dark-matter gravitational dynamics
- interactions between gas and radiation
- star formation

galaxy formation is a hard problem, <u>but</u>

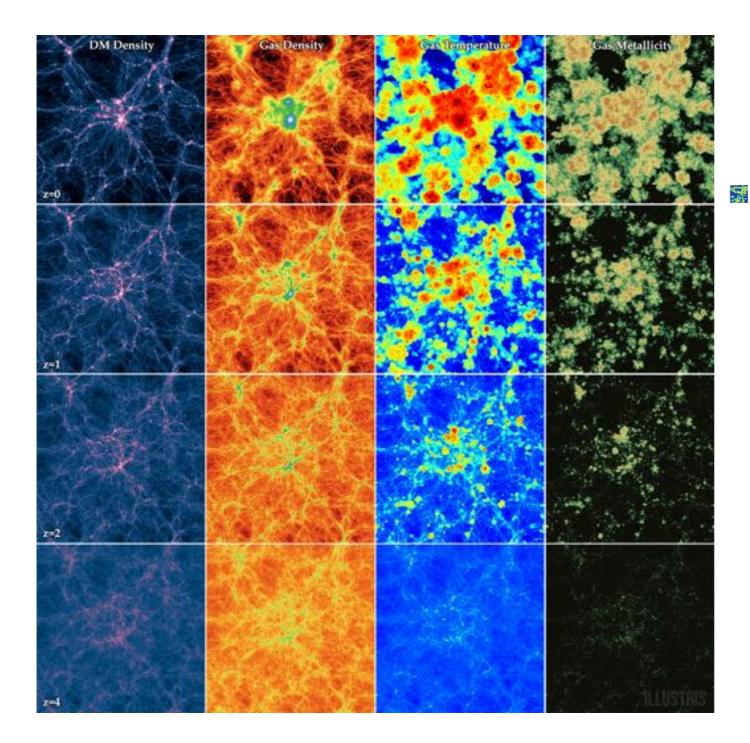
Fast developments (2003)



Projected mean metallicity of the gas in a 2 × 50³ simulation. The map is 11.3 h⁻¹ Mpc m_{dm}=8 10⁸ h⁻¹ Msun

Springel & Hernquist (2003)

Fast developments (2014)



ILLUSTRIS has 2 × 1820³ particles in a 100 Mpc box m_{dm}=6 10⁶ Msun

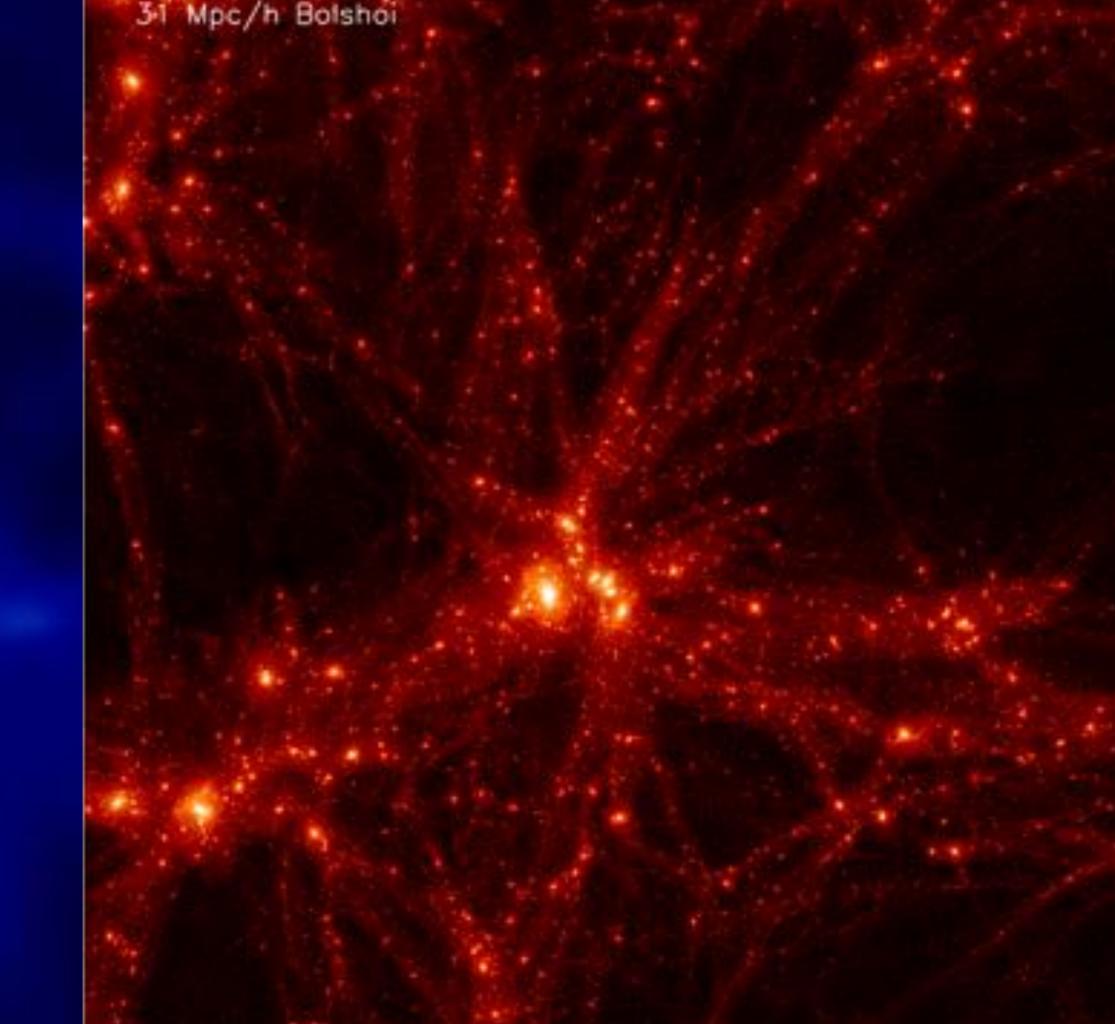
Outline of lectures

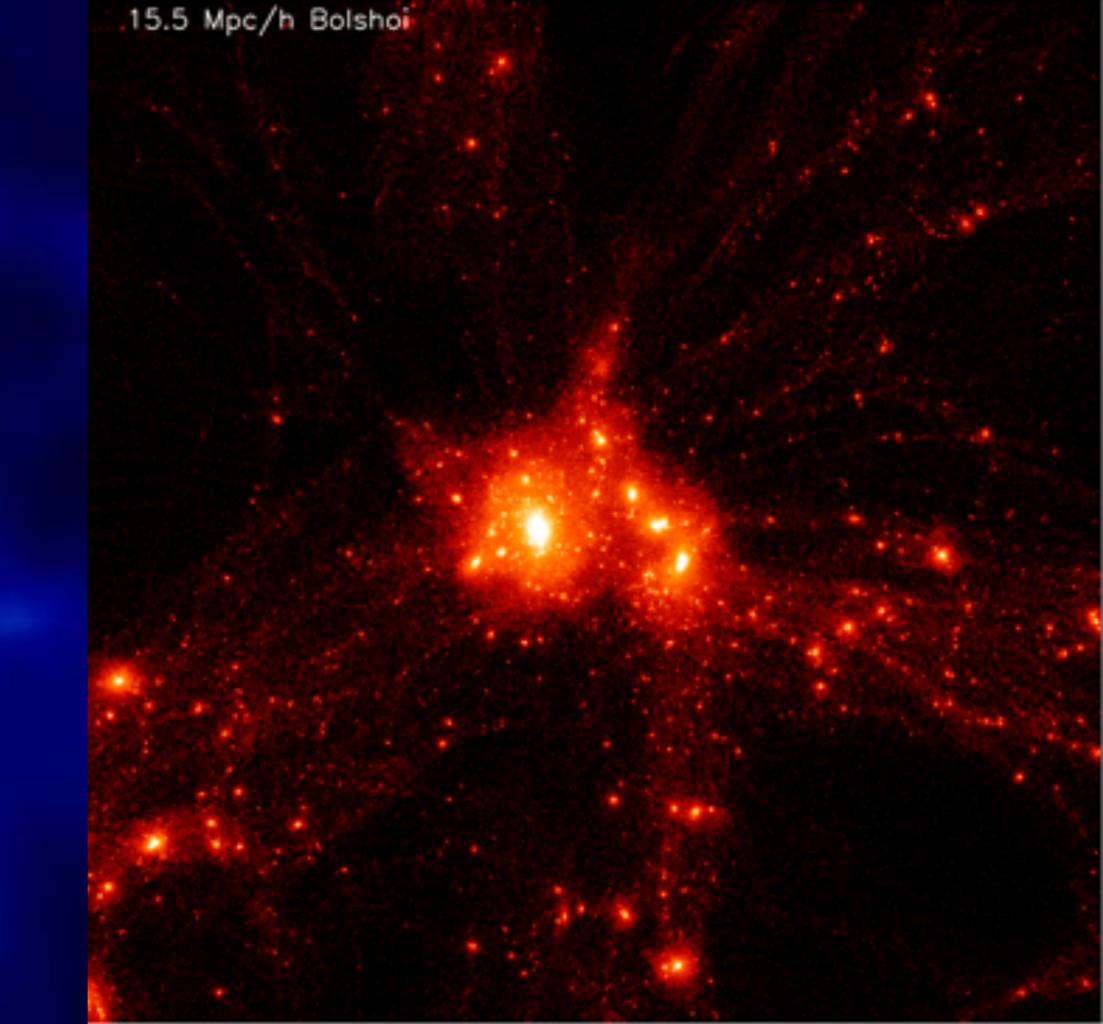
- Galaxies as crossroads between astrophysics and cosmology
- Cosmological simulations of galaxy formation
- The physics of star formation and feedback

Galaxies as crossroads between astrophysics and cosmology

bigBolshoi 1Gpc 8G particles

62 Mpc/h Bolshoi





Small Galaxy Group

 Galaxy Formation in a ΛCDM Universe.

7.7 Mpc/h Bolshoi

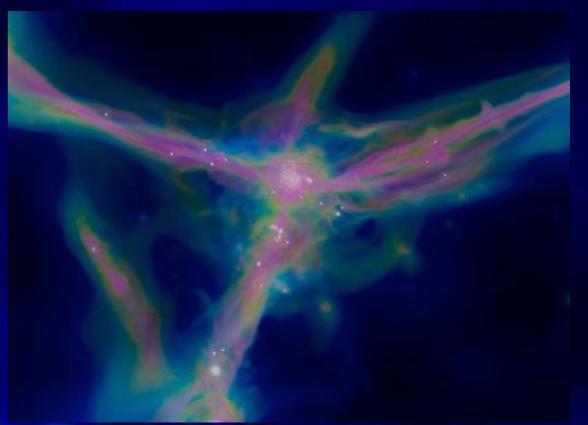
- Dynamic range:
- From Mpc to pc scales
- Physics:
- Gravity plus gas physics

-On Lenticular Galaxy NGC 5866

Our tool: Cosmological Simulations of Galaxy Formation



Galaxies as crossroads



- Different scales are related
- Different astrophysical processes are linked through non-linear mechanisms
- The evolution of the Universe matters for the formation of a single star
- Stars and galaxies are not isolated objects

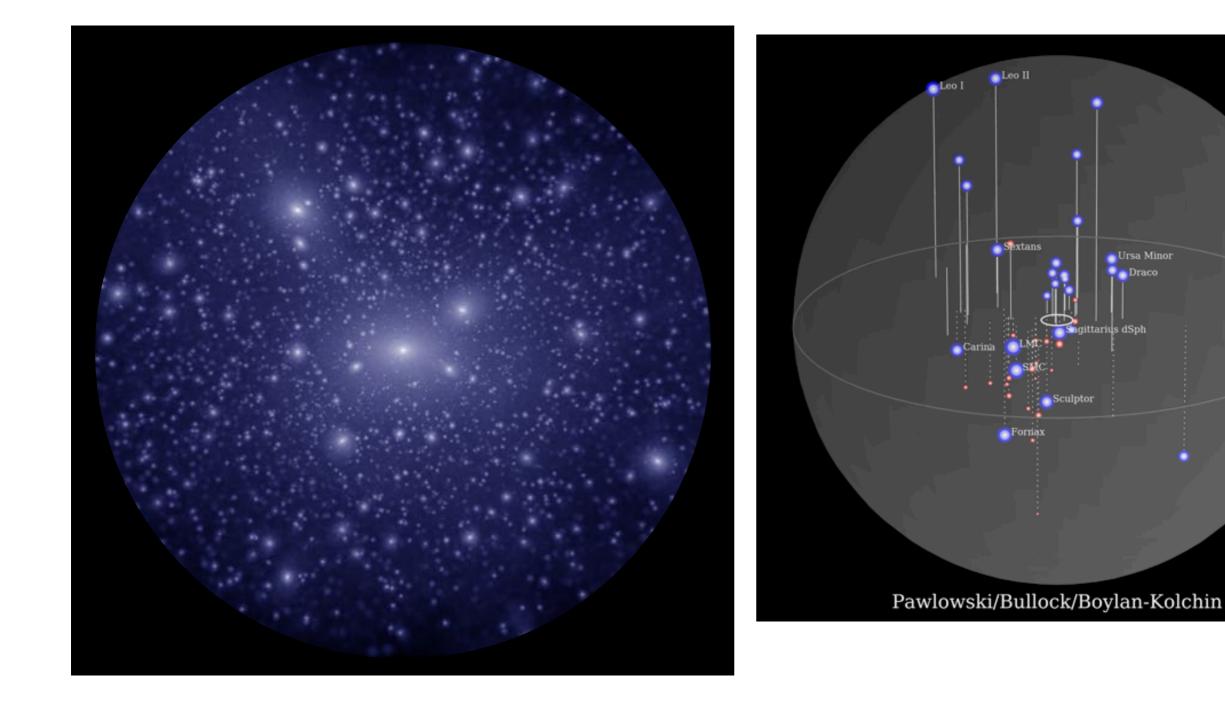
First application: Dwarf Galaxies as challenges to cosmology

Annual Review of Astronomy and Astrophysics Small-Scale Challenges to the ACDM Paradigm

James S. Bullock¹ and Michael Boylan-Kolchin²

Annu. Rev. Astron. Astrophys. 2017. 55:343–87
First published as a Review in Advance on June 28, 2017
The Annual Review of Astronomy and Astrophysics is online at astro.annualreviews.org
https://doi.org/10.1146/annurev-astro-091916-055313

low number of satellites



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WHERE ARE THE MISSING GALACTIC SATELLITES?

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AND

FRANCISCO PRADA Instituto de Astronomia, Apartado Postal 877, 22900 Ensenada, Mexico Received 1999 January 18; accepted 1999 April 15

> THE ASTROPHYSICAL JOURNAL, 524:L19–L22, 1999 October 10 © 1999. The American Astronomical Society. All rights reserved. Printed in U.S.A.

DARK MATTER SUBSTRUCTURE WITHIN GALACTIC HALOS

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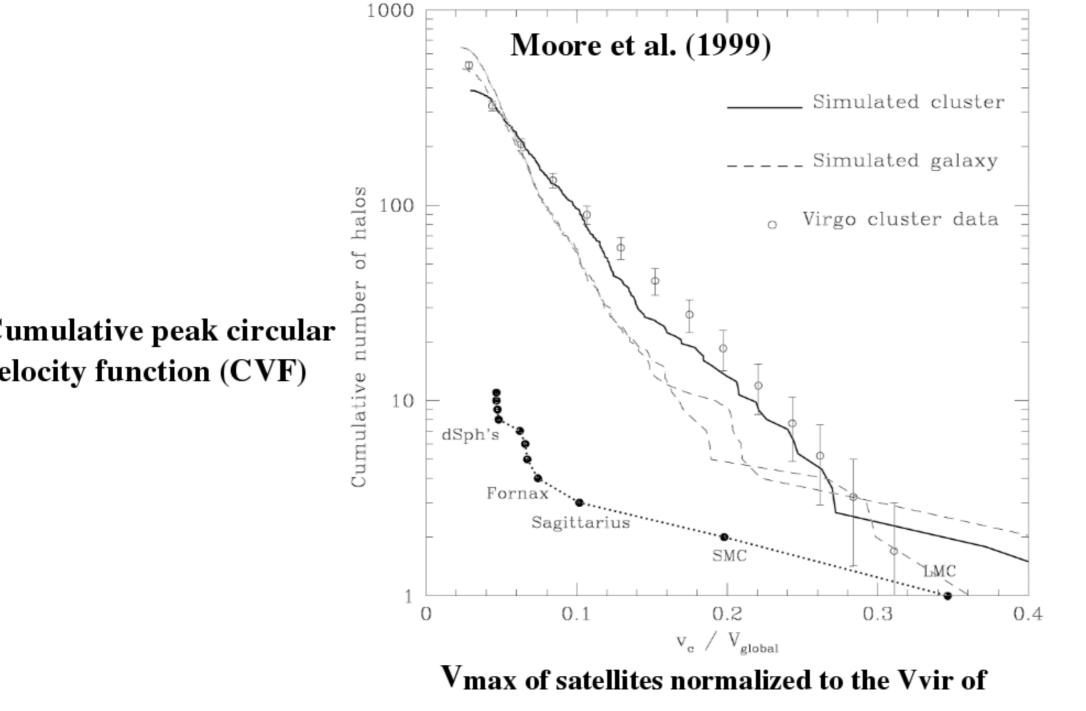
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AND

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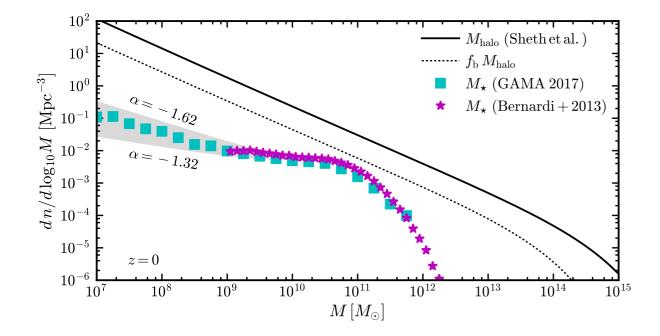
Quantifying the Missing Satellites Problem

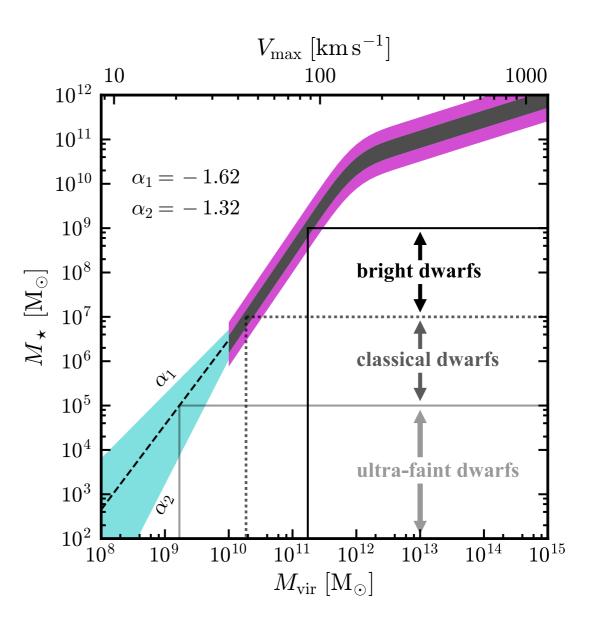


the parent halo

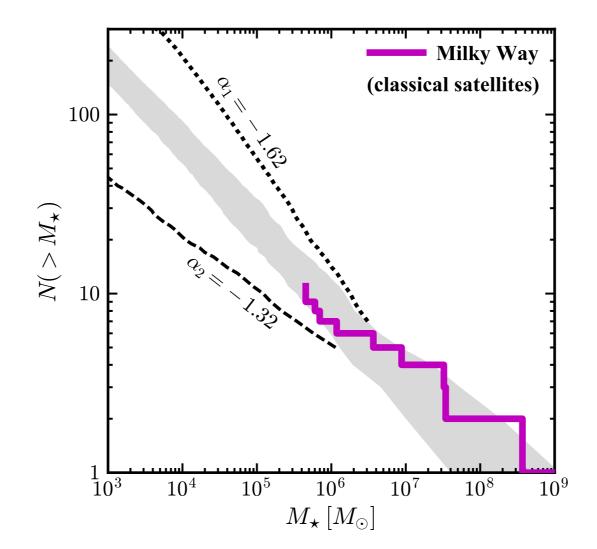
Discrepancy of ~ 2 orders of magnitude between the number of satellites predicted by numerical simulations of galactic systems and the Local Group.

Possible Solutions



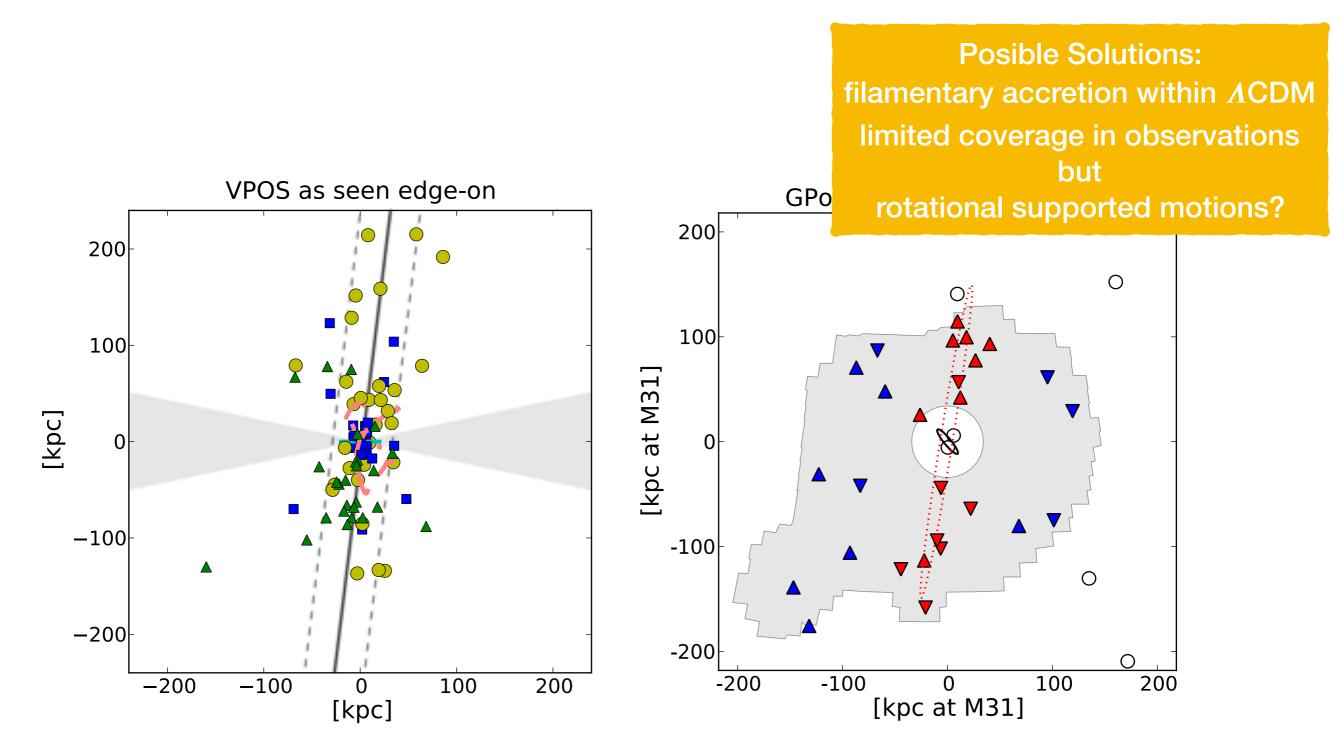


Possible Solutions



Which physical process drive this?

Planes of Satellites



First Tutorial Section

PROJECTS

- Main Goal: Get experience and skills with simulations data
- Beyond simple homeworks
- Scientifically relevant (and open) projects.

slice through galaxy plane

Radial momentum

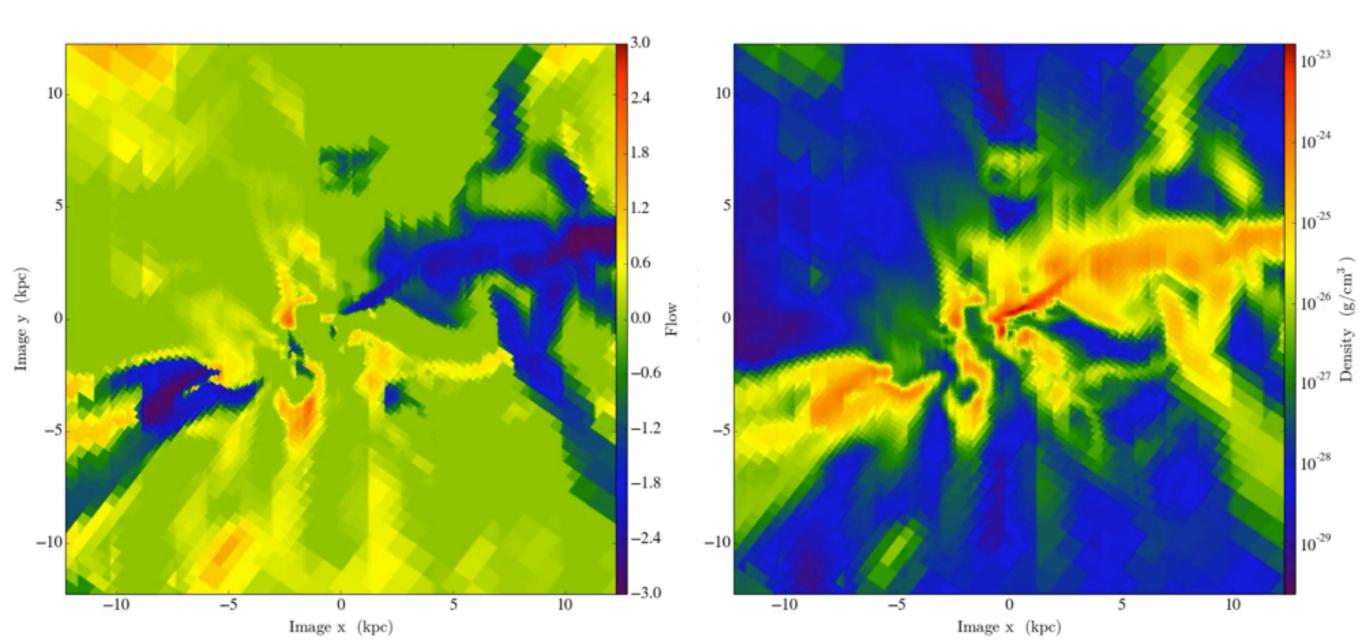
Gas density

DeGraf et al.

slice through galaxy plane

Radial momentum

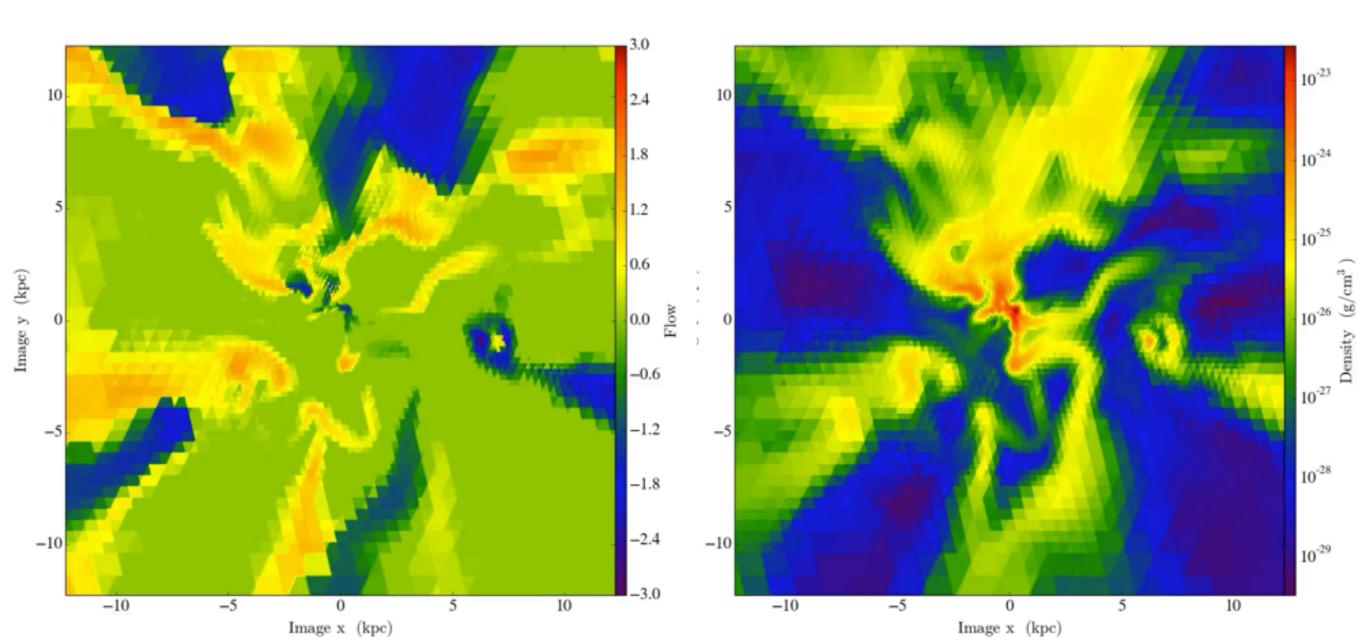
Gas density

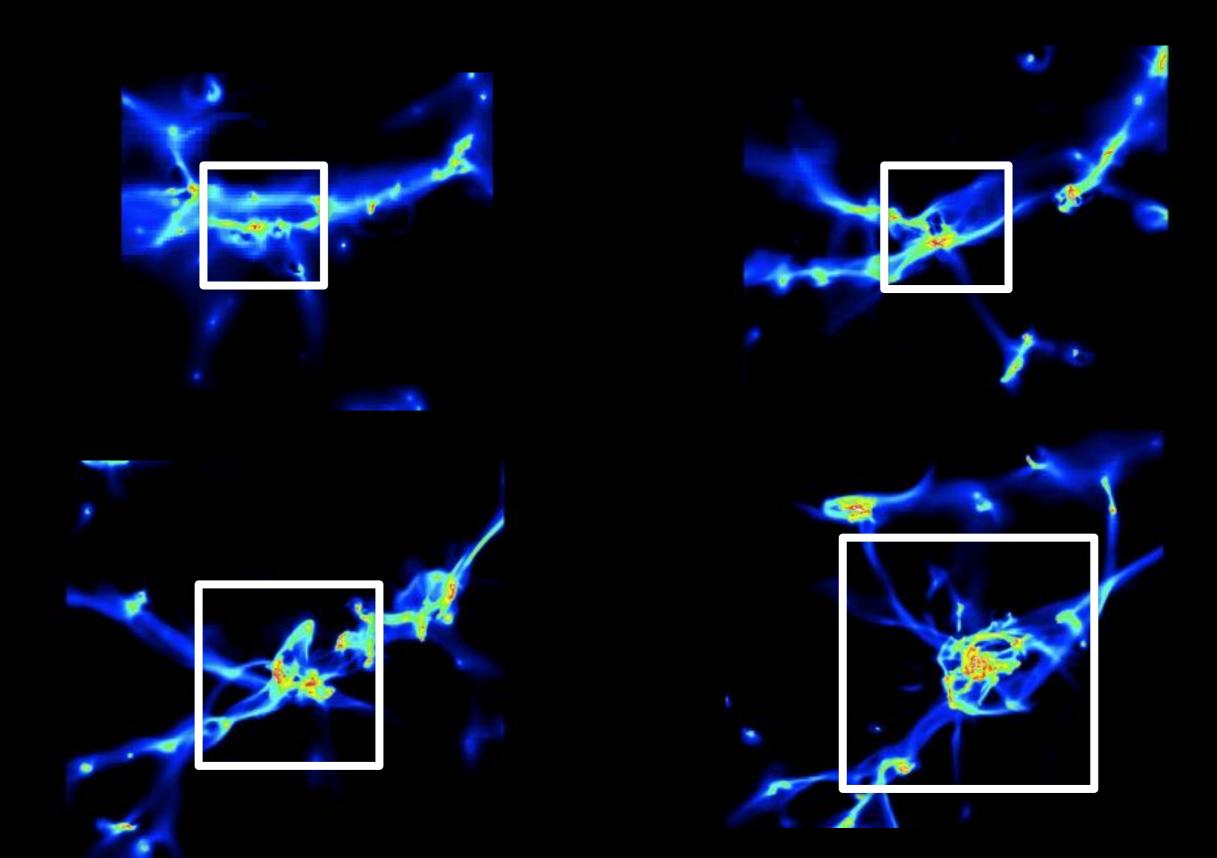


slice perpendicular galaxy

Radial momentum

Gas density

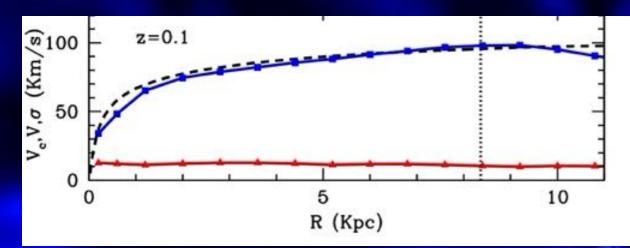




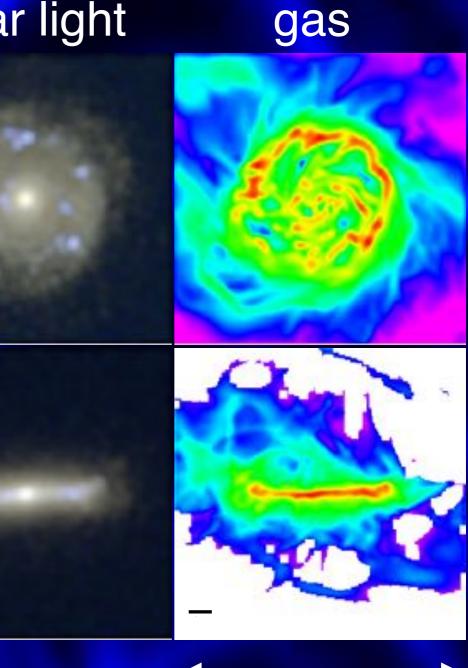
Data: Disc-dominated galaxy at z~0

stellar light

Rotation curve:



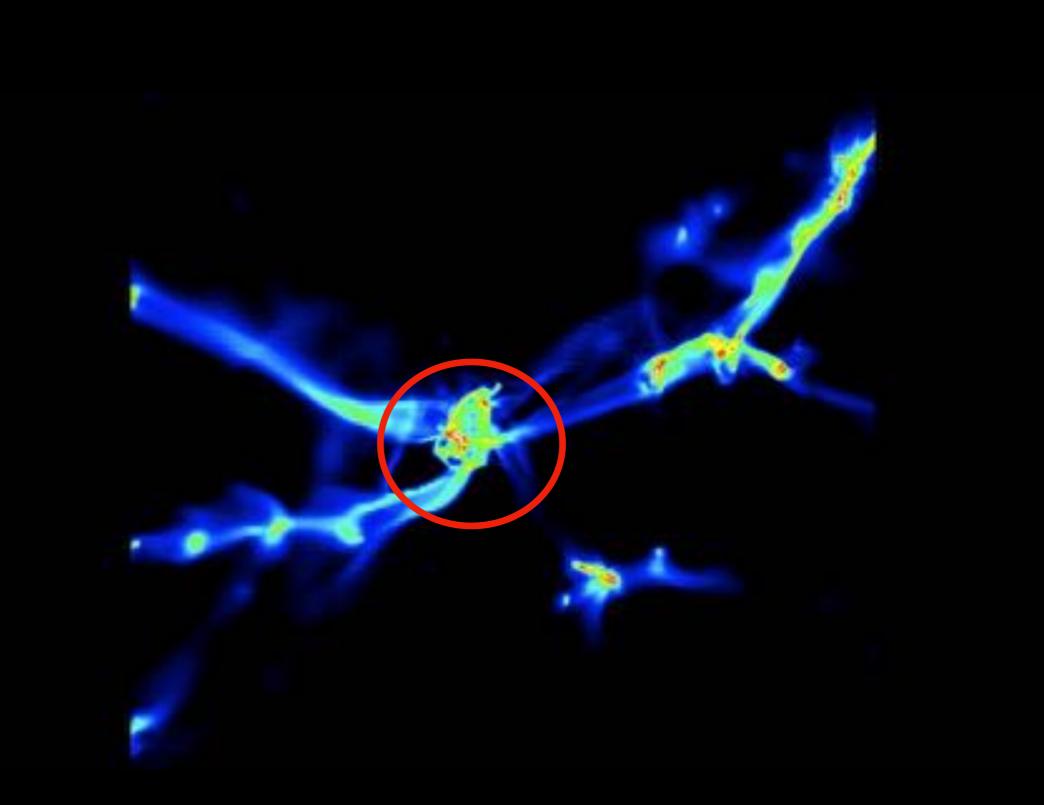
Ceverino et al. 2017a



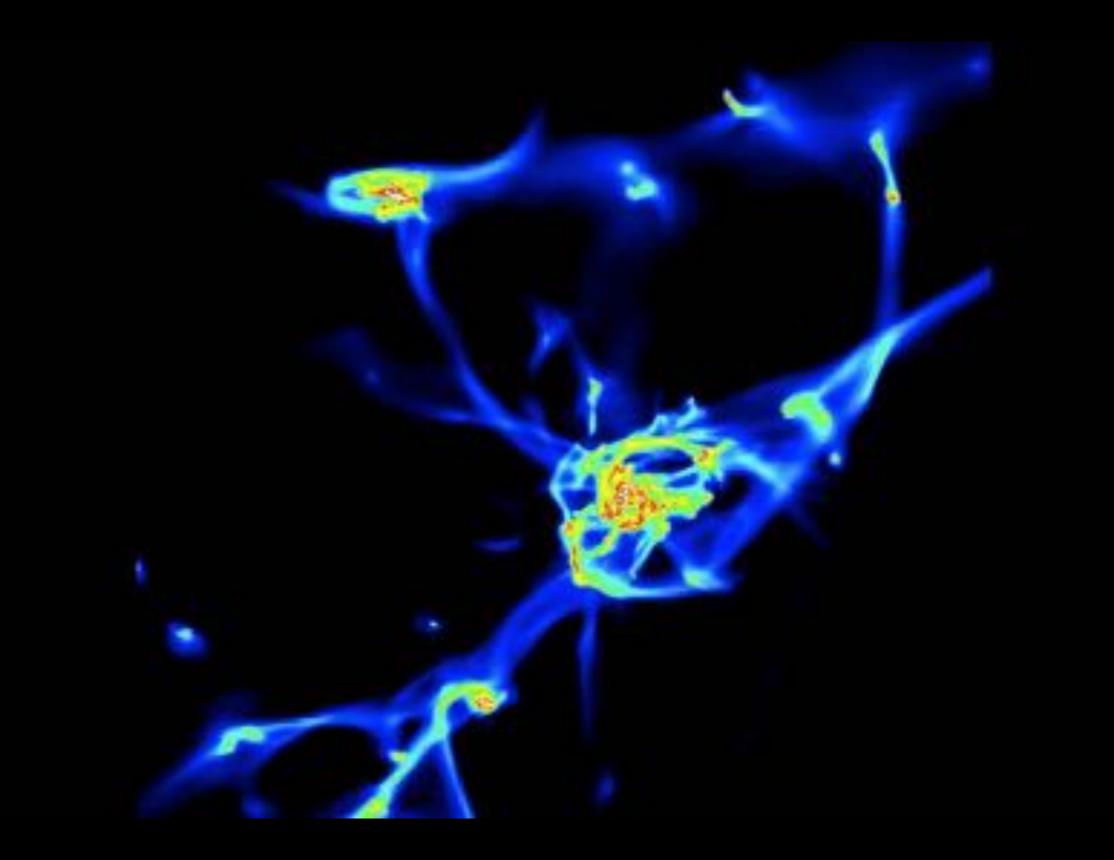
40 kpc

- 1. Accretion rate onto halos and onto galaxies: DM, gas, stars
- 2. Interaction of cold flows and Disk.
- 3. Angular momentum: in cold flows vs disk
- 4. Basic Structure of galaxies: Density profiles of gas, stars, DM. f_b?
- 5. Kinematics of gas: disk rotation curve, velocity dispersion
- 6. Kinematics of stars: bulge/disk decomposition
- 7. Gas outflows

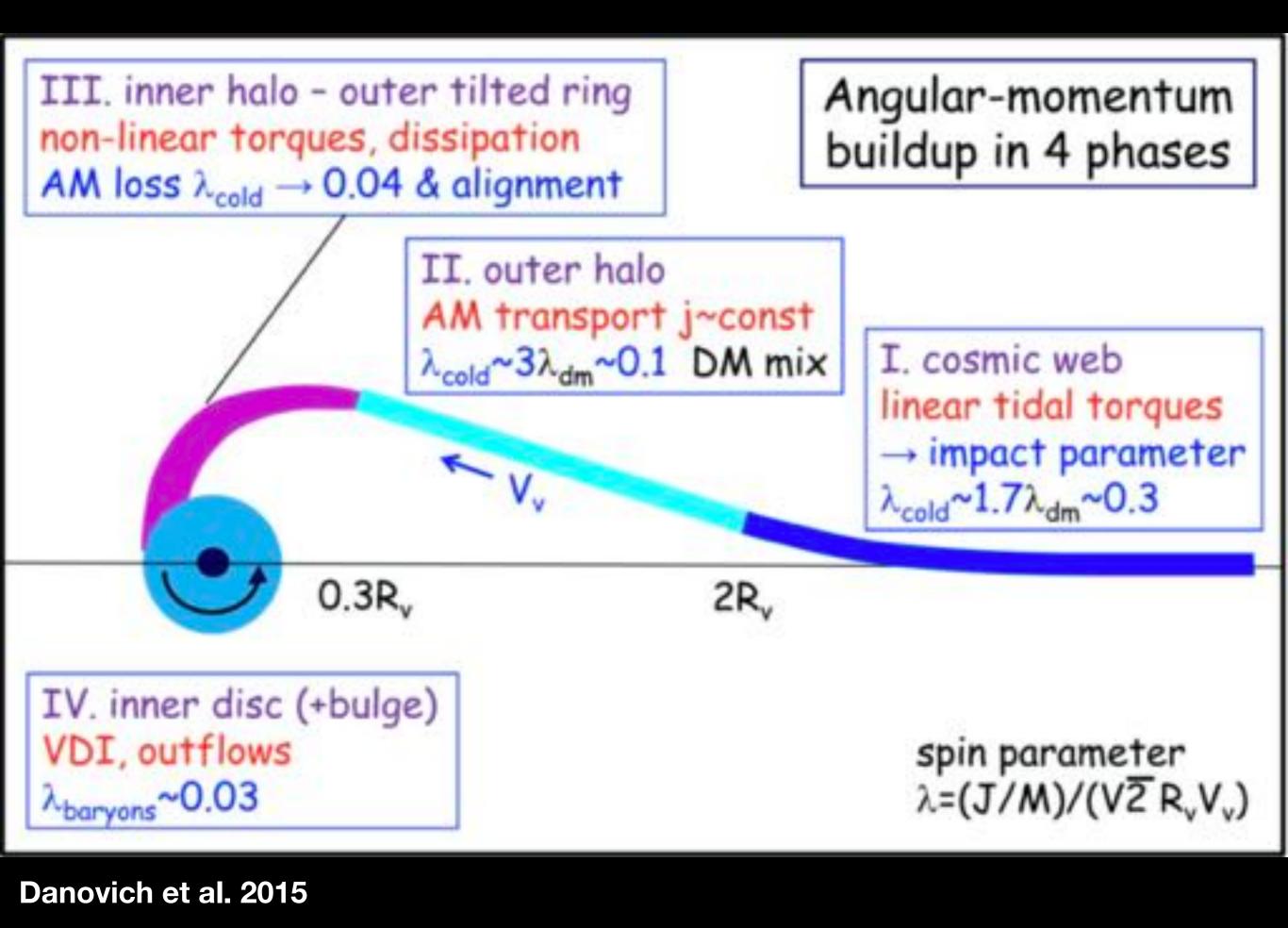
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ceverino@carina:~/MW3> more DataSet_README Brief description of the outputs from cosmological simulations of galaxy formation by Daniel Ceverino, Jerusalem, February 2010 These outputs files are binary fortran files that contains all the information inside a box of 4Rvir centered in th e major progenitor in cosmological simulations of galaxy formation. Name of the files: _____ The first part is the name of the simulation. Next, there are one or two letters that determine the type of data in the file: *_D* --> ID, Position, velocity and mass for DM particles (8 fields) *_S* --> ID, Position, velocity, mass and age for stellar particles (9 fields) *_Si*--> ID, Position, velocity, initial mass and age for stellar particles (9 fields) *_SZ*--> ID, Position, velocity, mass, age and SNII, SNIa metals mass fraction for stellar particles (11 fields) *_G* --> Cell size, position, velocity, density and temperature for gas cells (9 fields) *_GZ*--> Cell size, position, velocity, density, temperature and SNII, SNIa mass metals fraction for gas cells (11 fields) Next, there is a number that corresponds to the size of the cutout box. It is always equal to 4 times the virial radius (4Rvir). Finally, the file ends with the expansion parameter, a=1/(1+z), of the snapshot. For example, the file 'MW2_D120.a0.200.dat' contains the dark matter information of the major progenitor in simulat ion 'MW2' at a=0.200 (redshift z=4) inside a box of 120 proper kpc centered in that galaxy. Physical units: _____ Units are always in proper (not comoving) units: Cell size --> pc Position --> kpc Velocity --> km/s mass --> Msun age --> Gyr metals mass fraction --> dimensionless density --> H atoms / cm^3 temperature --> K Files format: These files are written in fortran binary format (big endian). All fields are single precision floating numbers, wi

th the exception of the fields of positions, velocities and masses for stars and dark matter: They have double-prec ision. The particle ID number is the only integer field.

Three take-home messages

- We are witnessing a rapid development towards a theory of galaxy formation
- Galaxies are crossroads for physics at different scales
- Research is fun

THANKS