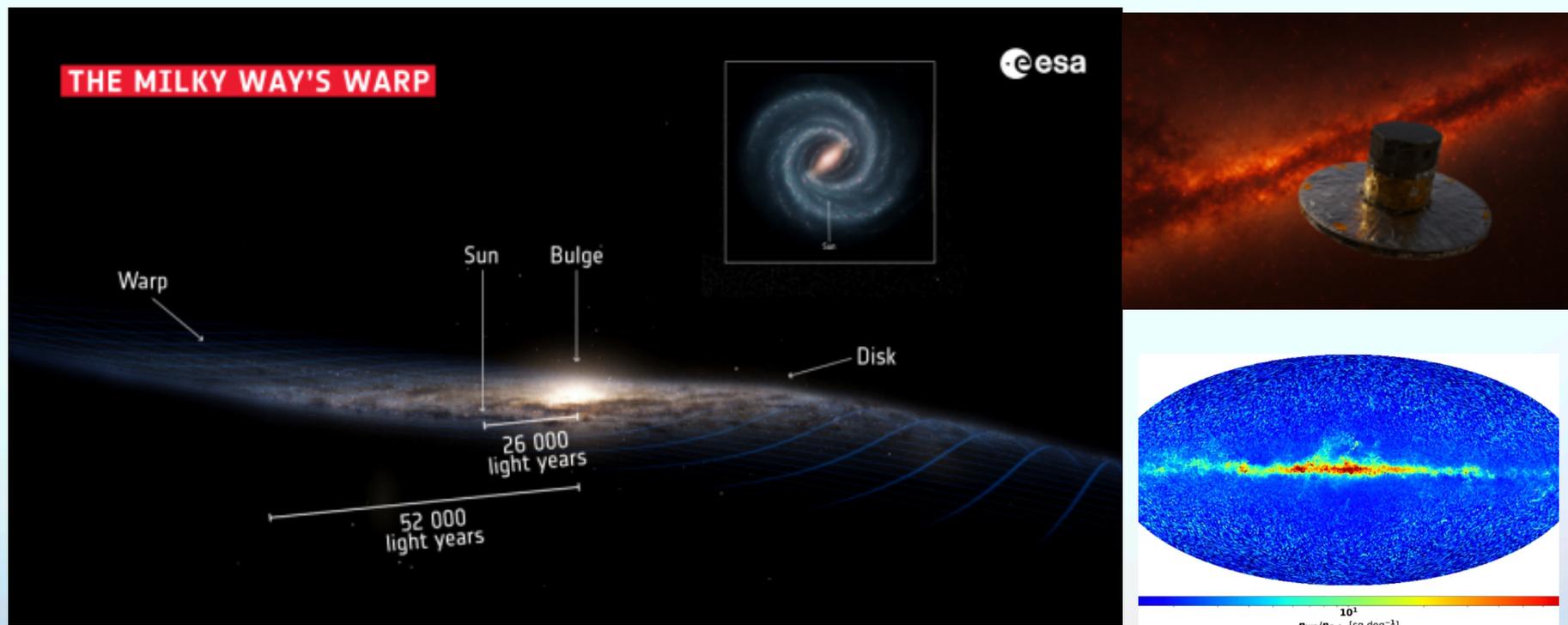


Vertical kinematics of young clusters in the Local Arm: from Gaia to GaiaNIR

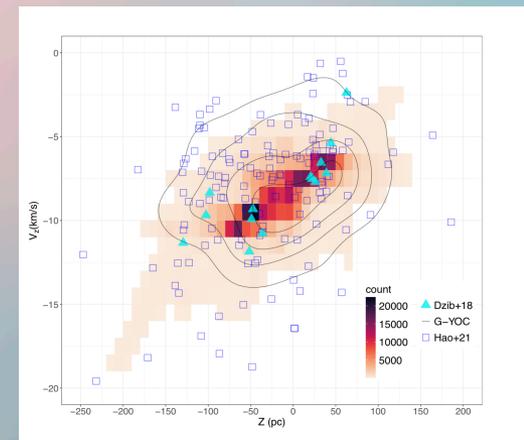
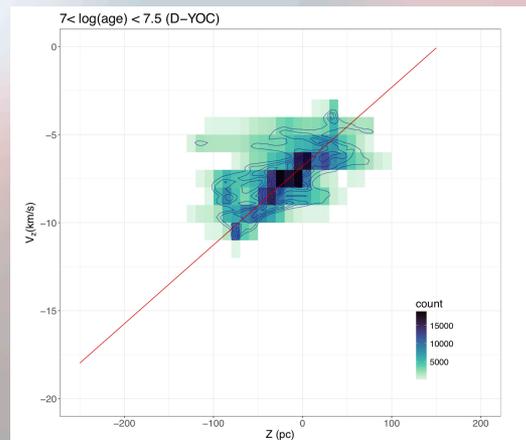
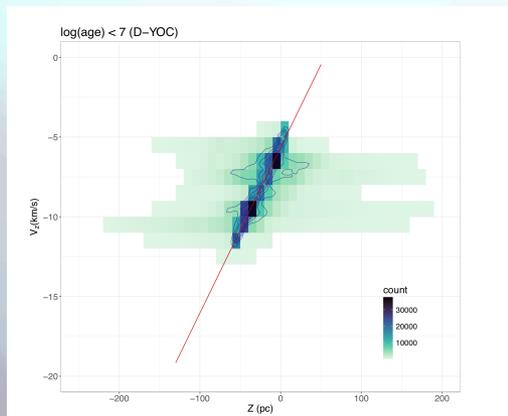
Linking astrostatistics, Galactic structure and future IR astrometry



Key result from Gaia:

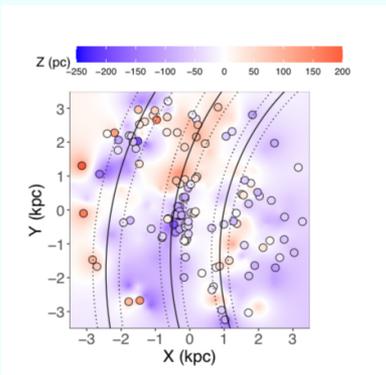
Young clusters as vertical oscillators

- ▶ Gaia + recent catalogues of **young open clusters** and **star-forming regions**
- ▶ Focus on clusters **younger than ~30 Myr** in the solar neighbourhood, sampling young clusters associated with the Local, Perseus and Sagittarius arms
- ▶ These clusters define a **tight linear relation in the V_z - Z plane**
- ▶ Consistent with **vertical harmonic motion**
- ▶ Provides an **independent estimate of the local mass density**
(Alfaro, Sánchez-Gil & Elmegreen 2025, MNRAS, 537, 3066)



Beyond the solar neighbourhood

Extending to the full Local Arm



▶ We build 4D (X, Y, Z, Vz) maps of young tracers in the Solar neighbourhood, in the Local, Perseus & Sagittarius arms

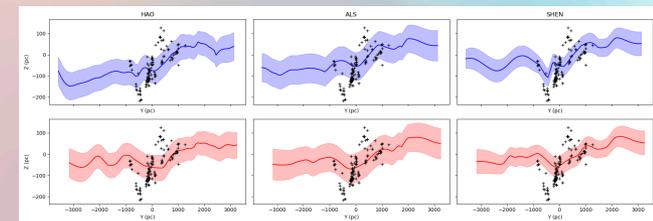
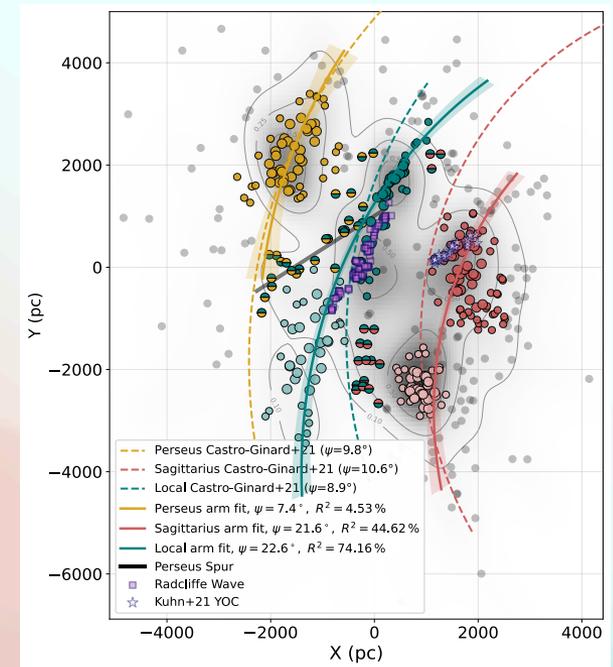
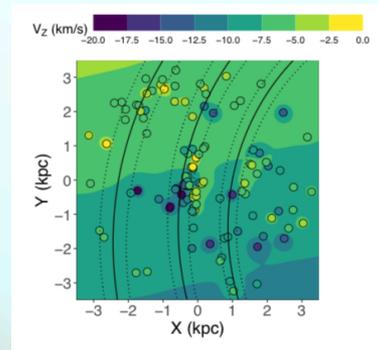
▶ Include OB associations and molecular clouds in addition to clusters

▶ We see vertical waves / corrugations linked to spiral-arm structure

▶ But Gaia alone is limited in dusty, crowded regions

▶ Strong extinction

▶ Incomplete sampling of embedded populations



Why GaiaNIR matters:

From Gaia to GaiaNIR: what we need

Near-IR astrometry turns our partial vertical maps into a complete picture of the young disc.

▶ Gaia (today):

- Limited in dusty, crowded regions
- Misses deeply embedded young populations
- Incomplete vertical picture of spiral arms

▶ GaiaNIR (tomorrow):

- Access to obscured segments of Local, Perseus & Sagittarius arms
- Better tracers for the local mass budget
- Time evolution of vertical corrugations

Take-home message

- ▶ Young clusters behave as **vertical oscillators** in spiral-like arms
- ▶ We are mapping **4D vertical structure** and **corrugations** with Gaia
- ▶ **GaiaNIR** is key to complete the picture in **dusty, embedded regions**

